# Some remarks on stochastically perturbed (Hamiltonian) systems 

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#### Abstract

We shortly mention several perturbation problems of classical dynamical systems by stochastic forces. We look more closely to the case of an Hamiltonian system consisting of a particle moving in $\mathbb{R}^{d}$ under the action of a force derived from a potential $V$ and an additional stochastic force. We report on a recent extension of ours with Zehnder of results by Potter, Mc Kean, Markus and Weerasinghe. Under restrictions on the growth of $V$ at infinity or attractiveness of the force towards the origin we give existence, uniqueness and stability results for the solution of the (stochastic) equations of motion. We also give a comparison theorem with solutions of a corresponding linearised system, via a Cameron-Martin-Girsanov-Maruyama type of formula. We also discuss the asymptotic behavior of the solution for large times, as well as the existence of a $\sigma$-finite, not finite invariant measure, the Lebesgue measure in phase space.


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## 1. Introduction

Stochastic perturbations of classical dynamical systems arise in various contexts. We would like to mention briefly some of them, somewhat related methodologically, and concentrate then on the presentation of some new results concerning stochastic perturbations of finite dimensional Hamiltonian systems.
Historically one can trace the origins of the theory of stochastic perturbations of dynamical systems in considerations involving random walks and Markov chains on one hand (e.g. in work by Bachelier and Markov) and statistical mechanics on the other hand (e.g. Maxwell, Boltzmann, Einstein and Gibbs). Work by A. Einstein (1905), Smoluchowski (1905), A. Einstein and E. Hopf (1910), P. Langevin (1911), N. Wiener (1923), S. Bernstein, P. Lévy, L. Ornstein-G. Uhlenbeck (1930) (cfr. e.g. [DeH-L], [ Ne 1$]$ ) was very influential for later developments (they were in fact. forerunners of the modern theory of stochastic differential equations). Typically in this work the time evolution of position $x$ and velocity $v$ of a classical system, a Newton particle, is considered. The evolution is described by equations of the type

$$
\begin{equation*}
\frac{d x}{d t}=v(t), \quad \frac{d v(t)}{d t}=K(x(t), t)-\gamma v(t) \tag{1.1}
\end{equation*}
$$

with $K(\cdot, t)$ a given (possibly space and time dependent) vector field over the state space $\mathbb{R}^{d}$ of the system, the deterministic force field. $\gamma \geq 0$ is a constant (damping). The stochastically
perturbed system is obtained by adding on the right hand side a stochastic term, e.g. of the white noise type, $\frac{d \omega(t)}{d t}$, with $\omega(t)$ a Brownian motion on $\mathbb{R}^{d}$, or more generally, of the form $\sigma(x(t), v(t), t) \frac{d \omega(t)}{d t}$ with $\sigma$ a (possibly space and time dependent) $d \times n$-matrix and $\omega(t)$ a Brownian motion on $\mathbb{R}^{n}$ :

$$
\begin{equation*}
\frac{d x}{d t}=v(t), \quad \frac{d v(t)}{d t}=K(x(t), t)-\gamma v(t)+\sigma(x(t), v(t), t) \frac{d \omega(t)}{d t} . \tag{1.1}
\end{equation*}
$$

Mathematically the discussion of the above equations is part of the theory of stochastic differential equations of Ito's or Stratonovich's type (with known rules to commute between the two) (see e.g. [ Ar ], $[\mathrm{Ga}]$ for introductions and [ $\mathrm{Ik}-\mathrm{Wa}],[\mathrm{RoW}]$ for more advanced topics).

In the study of properties of behavior of solutions for large times, which interests us most here, methods can be differentiated according to whether the damping term is present or not and whether $\sigma$ is degenerate or not. For the case with damping term $(\gamma>0)$ with $\sigma$ non degenerate see e.g. [Kh].
For the case of degeneration see e.g. [Ar], [Klie].
The case $\gamma=0$ (Hamiltonian conservative systems) is the main topics to be discussed below.
Let us note however at this point that a more abstract formulation of the problem is to look at stochastic perturbations of deterministic systems of 1 . order differential equations of the form

$$
\begin{equation*}
\frac{d y(t)}{d t}=\beta(y(t), t) \tag{1.2}
\end{equation*}
$$

with $y(t)$ a $\mathbb{R}^{n}$-valued function and $\beta(\cdot, \cdot)$ a given (time and space dependent) vector field. The stochastic perturbations can be of the type

$$
\tilde{\sigma}(y(t), t) \frac{d \tilde{\omega}(t)}{d t}
$$

with $\tilde{\omega}(t)$ a standard Brownian motion in $\mathbb{R}^{m}$ and $\tilde{\sigma}$ a $d \times m$-matrix-valued (time and spacedependent) function. The stochastically perturbed system is thus of the form

$$
\begin{equation*}
\frac{d y(t)}{d t}=\beta(y(t), t)+\tilde{\sigma}(y(t), t) \frac{d \tilde{\omega}(t)}{d t} . \tag{1.2}
\end{equation*}
$$

An important case discussed in recent years in connection with quantum physics is the one where $\beta$ is a gradient field $\nabla u$, with $u$ built from the real and imaginary parts of a function satisfying the Schrödinger equation and $\tilde{\sigma}$ is a constant (proportional to Planck's constant), in which case one can look upon (1.2)' as the equation of stochastic mechanics ([Ne 1,2], [BlCZ]) (we remark that (1.1)' is, on the other hand, not of this type). Existence and uniquenes of solutions of this equation have been discussed (see e.g. [AHKS], [Car], [ Ne 2 ], $[\mathrm{Bl}-\mathrm{Go}]$ ), as well as other problems like asymptotics for $t \rightarrow \pm \infty$ ([Car]), attainability of singularities of $\beta$ (see above references and [AFKS], $[\mathrm{Fu}]$ ), and questions related to the physical interpretation of the equation (e.g. $[\mathrm{DZ}],[\mathrm{Tr}]$ ). Let us also mention that recently the inverse problem of finding $u$ (i.e. $\beta$ ) given observations outside a bounded region has also been discussed, for the equations of stochastic mechanics ([ABKS]). There has also been much work on stochastic mechanical equations for motions on Riemannian manifolds (see e.g. [Gue], [Bl-CZ], [AB/HK1] (and references therein)), as well as on the stochastic mechanical equations for the motion of a (quantum mechanical) particle in an electromagnetic field ( $[\mathrm{Mor}]$ ). We also mention a connection between stochastic mechanical models with potentials which are Fourier transforms of measures and models of filter theory [Are1].
Another important case discussed in connection with classical dynamical systems (e.g. engineering problems) is the one where the vector field $\beta$ in (1.2)'is such that zero is an equilibrium point
and there exists a Ljapunov function (see e.g. [Kh]). Also the discussion of the small $\tilde{\sigma}$-limit has been pursued intensively, this limit is connected in the case of the equations of stochastic mechanics with the semiclassical limit in quantum mechanics (cfr. [Ma-Fe], [A-HK1], [Re], [Jo-M-S], [A-Ar]). We also mention that explicit connection between solutions of (1.2)' and orbits of the corresponding classical dynamical system (1.2) has been studied in the case of motions on certain manifolds with symmetries (Lie groups, symmetric spaces) (see e.g. [E], [AAH], [Are2]). We also recall that the case of equation (1.1)' or (1.2)' where $\beta$ and $\tilde{\sigma}$ are linear is extensively discussed in the literature, see e.g. [H], [GQ] (and references therein) and serves as a standard reference case in many issues.
The case of infinite dimensional state space is obviously more complicated and less results are known. For the linear case see e.g., in connection with filter theory, [Ko-Lo] (and references therein) resp. in connection with quantum fields [AHK], [Rö], [Ko]. For the nonlinear case see e.g. [AHK2], [A-Rö], [A-Ku]. The stationary case is particularly well discussed in connection with infinite dimensional Dirichlet forms (see [ARö]). For a recent development concerning a four space-time dimensional model for quantum fields see [A-HK-I]. We also mention that the equations of hydrodynamics (Euler and Navier-Stokes) with stochastic perturbation i.e.

$$
\begin{equation*}
\frac{d u}{d t}=-(u \cdot \nabla) u+\nu \Delta u+f, \operatorname{div} u=0 \tag{1.3}
\end{equation*}
$$

$\nu=0$ resp. $\nu>0$ being the viscosity constant, (with suitable initial and boundary conditions), $f$ being a deterministic plus stochastic force, have also been discussed recently in particular in connection with the proof of the existence of invariant probability measures, see [A-C], [Cr], [Fuj]. Finally let us mention that whereas all above examples are of elliptic-parabolic type, also some hyperbolic stochastic (partial) differential equations can be handled, see e.g. [ARu] (and references therein).
In this paper we shall concentrate on the case of an Hamiltonian system with finite dimensional state space. We shall see that even in this simple case relatively little is known and many interesting problems are open. More concretely we study an Hamiltonian system of the form

$$
\begin{align*}
& d x(t)=v(t) d t \\
& d v(t)=K(x(t)) d t+d \omega_{t}, \tag{1.4}
\end{align*}
$$

with $\omega_{t}$ a Brownian motion in $\mathbb{R}^{d}$, started at time 0 at the origin, i.e. $\left\{\omega_{t}, t \geq 0\right\}$ is a Markov stochastic process, consisting of independent Gaussian distributed random variables with independent Gaussian increments $\omega(t)-\omega(s), s \leq t$, with mean zero and covariance $(t-s)$.
The inital data $x(0)=x_{0}, v(0)=v_{0}$ are given in $\mathbb{R}^{2 d}$ (e.g. independent of the point $\omega$ in the underlying probability space). We can rewrite (1.4) in the form

$$
\begin{align*}
y & \equiv\binom{x}{v}, \beta(y(t)) \equiv\binom{y(t)}{K(x(t))} \\
d y(t) & =\beta(y(t)) d t+\sigma d \tilde{\omega}_{t} \text { with } \tilde{\omega}_{t}=\binom{b_{t}}{\omega_{t}}, \tag{1.5}
\end{align*}
$$

$b_{t}$ a Brownian motion started at time 0 in 0 , independent of $\omega_{t}$.
Models of this type are frequently discussed in the literature mainly numerically, in connection with vibrations in mechanical systems, wave propagation and other problems (see e.g. [Li], [Kr-S], [Sch], [Ka]).
Hamiltonian systems of this type are also obviously important in celestial mechanics (see e.g. $[\mathrm{Mo}]$ ). One of the reasons for which they have not been so well studied on a mathematical basis is the quite intricated nature of the classical motion themselves (as compared to dissipative systems). Let us mention that a number of papers have been devoted to the study of the
corresponding deterministic problem (see e.g. [Di-Z] and references therein). For stochastic perturbations of other type (multiplicative ones) see e.g. [APW], [P]. Orbits of very different long time behavior cannot be separated in finite time intervals, stable and unstable behavior being mixed. In particular, the orbits are in general neither globally stable nor asymptotically stable, see e.g. [Ar], $[\mathrm{Mo}]$ and [ $\mathrm{Mo}-\mathrm{Ze}$ ]. The nature of the orbits depends in particular on the dimension of the system. For example in the case of more than three degrees of freedom the phenomenon of Arnold diffusion can met one. It is interesting to find out what one can say about perturbations in case the force is stochastic, hence typically non smooth, in particular, whether the complicated behaviour of classical orbits is enhanced or whether stochasticity is so strong as to change the picture radically. Potter [ Po ] (see also Mc Kean [ McK ]) analysed the case of a 1-dimensional nonlinear oscillator perturbed by a white noise force, described by the equations (1.4). Under assumptions on the force $K(x)=-V^{\prime}(x) \quad, V \in \mathbb{C}^{1}(\mathbb{R})$ being attracting towards the origin, i.e. $x \cdot K(x) \leq 0$, Potter proved the existence of global solutions and results about recurrence as well as the invariance of Lebesgue measure $d x d v$ under the flow given in (1.4). These results recently have been extended in [Mark-W] who studied in particular winding numbers around the origin associated with the solution process $(x, v)$.

Existence and uniqueness results for solutions of higher dimensional second order Ito equations, as the systems of the type given in (1.1) have been called by Borchers [Bo], have been deduced by Goldstein [Go] for systems with globally Lipschitz continuous force K, and Narita [ Na] in case there exists a function, decreasing along the paths analogously to a Ljapunov function in the deterministic theory.

In the first part of this paper we shall study equations of the form (1.4) in the case where x , and v run in $\mathbb{R}^{d}$. In Section 2 we establish existence and uniqueness results for strong solutions of the equations, under assumptions on K which are of the type $K(x)=-\nabla V(x)$ for some $V \in C^{1}\left(\mathbb{R}^{d}\right)$, with either a condition of the form $K$ is linear or $x \cdot K(x) \leq 0$ for $|x|$ sufficiently large or $V(x)$ sufficiently increasing at infinity. Then the solution process possesses the Markov property and continuous sample paths, furthermore it depends continuously on the initial conditions.
In section 3 we compare the solutions of the nonlinear system (1.4) with the ones of a corresponding linear system given by

$$
\begin{gather*}
d x=v d t \\
d v=-\gamma x d t+d w \tag{1.6}
\end{gather*}
$$

and $\gamma$ a constant dxd-matrix with positive eigenvalues. This is done by establishing a Comeron-Martin-Girsanov-Maruyama type formula for the Radon Nikodym derivative of the probability measures. We apply these results to prove some properties which hold with probability one for the nonlinear system by exploiting their validity for the associated linear system.
In section 4 we recover some features of the behavior of the solution process of the nonlinear system for large times. In particular, we give estimates for the energy functional of the process. We introduce the generator of the diffusion, solving (1.4), and show its hypoellipticity (in the sense that the coefficient functions span the tangent space to phase space). By a Hörmander's type theorem one obtains the absolute continuity of the transition probability of the transition probability w.r. to Lebesgue measure without further restrictions but continuity of the coefficient functions. We also show that for the solution process of (2.1) an existence and uniqueness theorem for a $\sigma$-finite invariant measure, which is the "normalized" Lebesgue measure, holds. For more details on the result presented here we refer to [AHZ], $[\mathrm{H}]$.

## 2. Existence and Uniqueness of Solution

We consider the Hamiltonian system with stochastic force given by the stochastic differential system

$$
\begin{align*}
& d x=v d t \\
& d v=K(x) d t \quad+\quad d w_{t} \tag{2.1}
\end{align*}
$$

where $t \in \mathbb{R}_{+}$is time, $x(t)$ is position in $\mathbb{R}^{d}$ at time $t, v(t)$ is velocity at time $t$.
$K(\cdot)$ is a deterministic force, $\left(w_{t}, \mathcal{F}_{t}, t \in \mathbb{R}_{+}\right)$is a Brownian motion in $\mathbb{R}^{d}$ started at the origin at time zero, see section 1 for motivations. The initial conditions are given.
It is useful to introduce the phase space variable $y=(x, v) \in \mathbb{R}^{2 d}$ and to write (2.1) in the form

$$
\begin{equation*}
d y=\beta(y) d t+\sigma d \tilde{w}_{t} \tag{2.2}
\end{equation*}
$$

with

$$
\beta(y) \equiv\binom{v}{K(x)}, \quad \sigma \equiv\left(\begin{array}{cc}
0 & 0 \\
0 & 1
\end{array}\right), \quad \tilde{w}_{t}=\binom{b_{t}}{w_{t}}
$$

where $\left(b_{t}, \mathcal{A}_{t}, t \in \mathbb{R}_{+}\right), \mathcal{A}_{t}=\sigma\left\{b_{s}, s \leq t\right\}$, is an $\left(\mathcal{F}_{t}\right)$ independent Brownian motion in $\mathbb{R}^{d}$ issued from 0 at time 0 . The initial condition $y(0)$ is given. For simplicity we state the theorem assuming $y(0)=0$.

## Theorem 2.1

Each of the following conditions is sufficient for the existence of pathwise solutions of (2.1), (2.2) for all $t \in \mathbb{R}_{+}$:
a) 1) $|K(\alpha)-K(\beta)| \leq C_{1}|\alpha-\beta| \quad \forall|\alpha|,|\beta|<R$, for some constants $R, C_{1}$.
2) $|K(a)| \leq C_{2}(1+|\alpha|) \quad \forall \alpha \in \mathbb{R}^{d}$, for some constant $C_{2}$.
b) 1) $\alpha \mapsto K(\alpha)$ is a locally Lipsschitz function from $\mathbb{R}^{d}$ into $\mathbb{R}^{d}$. Moreover
2) For $d \geq 1: K(\alpha)=-\nabla V(\alpha)$ for some $V \in C^{1}\left(\mathbb{R}^{d}\right)$
3) $\alpha \cdot K(\alpha) \leq 0 \quad$ for all $\alpha \in \mathbb{R}^{d}$.

## Remark

Corresponding statements holds for $t \geq t_{0}$ with initial condition $y\left(t_{0}\right)$ given, and $\left(\alpha-y\left(t_{0}\right)\right) \cdot K(\alpha) \leq 0$.

## Proof:

Statement a) can be proven by a stochastic version of Picard-Lindelof method of iteration, see e.g. [ Ar 1$]$ (Cor. 6.3.4) .

1) For $d=1$ the statement b) in a special case of a result of Potter [Po], see e.g. [McK], [ Na 1$]$.

We give a proof valid for $d \geq 1$ which uses a Ljapunov function for the solution of the stochastic differential equation. To this end we introduce the energy functional:

$$
\begin{equation*}
W(y)=\frac{1}{2}|v|^{2}+V(x)-V(0) . \tag{2.3}
\end{equation*}
$$

Let L be the differential operator (generator) associated with (2.2) i.e.

$$
\begin{equation*}
L=v \cdot \nabla_{x}+K(x) \cdot \nabla_{v}+\frac{1}{2} \Delta_{v} \tag{2.4}
\end{equation*}
$$

where $\nabla_{x}\left(\nabla_{v}\right)$ are the gradient w.r. to $x(v)$, respectively, and $\triangle_{v}$ is the Laplacian w.r. to $v$. All operators are acting on functions of $(x, v) \in \mathbb{R}^{2 d}$. Applying L to the energy function W we find

$$
\begin{equation*}
L W=\frac{1}{2} \tag{2.5}
\end{equation*}
$$

Furthermore from condition b)2)

$$
\begin{equation*}
\alpha \cdot K(\cdot)=-\alpha \cdot \nabla V(\alpha)=|\alpha| \frac{\partial V(\alpha)}{\partial|\alpha|} \tag{2.6}
\end{equation*}
$$

we conclude for $|\alpha| \neq 0$

$$
\begin{equation*}
V(\alpha)=V(0)-\int_{0}^{|\alpha|} \frac{1}{|\beta|}(\beta \cdot K(\beta)) d|\beta| \geq V(0) \tag{2.7}
\end{equation*}
$$

where we used assumption b)3).
For the energy functional in (2.3) this implies

$$
\begin{equation*}
W(y) \geq \frac{1}{2}|v|^{2} . \tag{2.8}
\end{equation*}
$$

and

$$
\left|\sigma \nabla_{y} W(y)\right|^{2}=|v|^{2} \leq 2 W(y) .
$$

Since $\beta$ in locally Lipschitz continuous, following [ Na 2 ] and [ $\mathrm{Ik}-\mathrm{Wa}$ ] (Def. 2.1) we can discuss local solutions.
In a first step we introduce stopping times

$$
\sigma_{n}(w) \equiv \inf \{t \geq 0| | y(t) \mid \geq n\}
$$

of the process $Y \equiv(y(t), t \geq 0)$ and define the explosion time $e(0)$ of $Y$ for given innitial condition $y(0)=0$, by

$$
\begin{equation*}
e(0)=\sup _{n \in N}\{\inf t| | y(t) \mid \geq n\} \tag{2.9}
\end{equation*}
$$

with inf replaced by $+\infty$ if the set is empty. For any $n \in \mathbb{N}$ the process $Y_{n} \equiv\left(y\left(t \wedge \sigma_{n}\right) \mid t \geq 0\right)$ defines a local solution of (2.1) on the ball $B_{n}(0)$, since there exists a uniform Lipschitz constant on every $B_{n}(0)$ which guarantees the existence and pathwise uniqueness of solutions. For $n \rightarrow \infty$ the local solutions $Y_{n}$ converge a.s. to the maximal solution of (2.1), corresponding to the martingale

$$
y^{i}(t)-y^{i}(0)-\left[\sum_{k=1}^{d} \int_{0}^{t} \sigma_{k}^{i}(y(s)) d \tilde{w}_{k}(s)+\int_{0}^{t} \beta^{i}(y(s)) d s\right] \quad i=1, \ldots, d
$$

with expectation 0 , for $t \in[0, e(0)]$ cfr. [ Na 3 ]. It is shown in [ Na 2 ] that the proof of existence of a global solution Y for initial condition $y(0)=0$ is equivalent to an infinite explosion time $e(0)$, i.e. to

$$
\begin{equation*}
N_{0} \equiv\left\{e(0)<\infty, \text { and } \lim _{t \uparrow(0)}|y(t)|=+\infty\right\} \tag{2.10}
\end{equation*}
$$

satisfying

$$
\begin{equation*}
\mathbb{P}\left(N_{0}=\{e(0)<\infty\}\right)=1 \tag{2.11}
\end{equation*}
$$

For times before an explosion occurs we can reexpress the energy functional $W$ given by (2.3) by applying Ito's formula to the differential

$$
d W=v(t) d w(t)+\frac{d}{2} d t+o\left(t^{\frac{3}{2}}\right)
$$

obtaining

$$
\begin{equation*}
W(y(t))=W(0)+\int_{0}^{t} v(s) d w(s)+d \frac{t}{2} \tag{2.12}
\end{equation*}
$$

The process $W(y(t))$ can be simplified by introducing analogously to [ McK ] a new Brownian motion

$$
a(\tau(t)) \equiv \int_{0}^{t} v(s) d w_{s}
$$

with a clock running according to the time

$$
\tau(t)=\int_{0}^{t}|v(s)|^{2} d s
$$

Under the assumption of Theorem 2.1 b) a global solution of (2.1) in established due to (2.11) by the following

## Lemma 2.2

Adopting the hypothesis of Theorem 2.1 b ) there holds

$$
\mathbb{P}(e(0))=1
$$

## Proof:

The proof of the higher dimensional statement can be reduced to the one for the one dimensional case in $[\mathrm{Po}]$ with $y$ being replaced by $|y|$. The proof is by contradiction, distinguishing the cases $\tau(e(0))<\infty$ and $\tau(e(0))=\infty$, and using the sample path properties of the Brownian motion $a(\cdot)$. Thus the a.s. finiteness of $|y(t)|$, where $0 \leq t \leq e(0)$, is deduced, which yields the contradiction.

## Theorem 2.3

Each of the conditions a) and b) of Theorem 2.1 is sufficient also for pathwise uniqueness of solutions of the equations (2.1), (2.2), i.e. if $y, y^{\prime}$ are two solutions of (2.2) on the same probability space with the same filtration s.t. $y(0)=y^{\prime}(0)$ a.s., then

$$
y(t)=y^{\prime}(t) \quad \forall t \geq 0 \text { a.s. }
$$

## Proof:

a) This case is covered e.g. by [Fr].
b) This case follows from [ Ik -Wa], (Theorem 3.1 p . 164) since the coefficients of the equations (2.1),(2.2) are in particular locally Lipschitz continuous. This yields uniqueness for times $t, 0 \leq$ $t \leq e(0)$, and this together with the fact (demonstrated in the proof of Theorem 2.1) that $e(0)=\infty$ a.s. yields uniqueness for all $t \geq 0$.
It also follows from [ $\mathrm{Ik}-\mathrm{Wa}$ ] that (2.1), (2.2) have unique strong solutions.

## 3. A Girsanov Formula

In this section we shall investigate whether the probability measure associated with the solution of the stochastic differential equation (2.1), i.e.

$$
d y(t)=\beta(y(t)) d t+\sigma d \tilde{w}_{t}
$$

with

$$
\beta(y)=\binom{v}{K(x)}, \quad \sigma=\left(\begin{array}{c}
0 \\
0 \\
0
\end{array}\right), \quad \tilde{w}=\binom{b_{t}}{w_{t}}
$$

is absolutely continuous with respect to the probability measure associated with the corresponding Gaussian process given by the stochastic differential equation

$$
\begin{equation*}
d \eta(t)=a(\eta) d t+\sigma d \tilde{v}_{t} \tag{3.1}
\end{equation*}
$$

with

$$
\begin{aligned}
\eta & =\binom{z}{u}, & & z, u \in \mathbb{R}^{2 d} \\
a(\eta) & =\binom{u}{-\gamma z}, & & \gamma \text { a constant matrix }
\end{aligned}
$$

and vice versa. If this holds, almost sure statements concerning the nonlinear system are equivalent to almost sure statements concerning the linear system.
In order to show the equivalence we shall derive a (Cameron-Martin-Maruyama-) Girsanov formula relating the probability measures.

## Lemma 3.1

Let $W(y(t)) \equiv \frac{1}{2}|v(t)|^{2}+V(x(t))-V\left(x_{0}\right)$ with $V$ s.t. $-\nabla V=K, Y=(y(t), t \geq 0)$ satisfying (2.1) or (2.2). Then, for all $t \geq 0$
i) $\mathbb{E}(W(y(t)))=\mathbb{E}(W(0))+d \frac{t}{2}$
and
ii) $\mathbb{E}\left(W^{2}(y(t))\right)=\mathbb{E}\left(W(0)^{2}\right)+\frac{d}{2} \int_{0}^{t} \mathbb{E}\left(\frac{|v(s)|^{2}}{2}+V(x(s))-V(0)\right) d s$ $+\frac{1}{2} \int_{0}^{t} \boldsymbol{E}\left(|v(s)|^{2}\right) d s$

## Proof:

Statement i) follows from 2.12 by taking expectation, and using that $\int_{0}^{t} v(s) \cdot d w(s)$ is a martingale with expectation zero.
ii) Given $F \in C^{2}(\mathbb{R})$, and a solution $Y=(y(t), t \geq 0)$ of (2.1) or (2.2) for $W(y(t))=$ $\frac{1}{2}|v(t)|^{2}+V(x(t))-V\left(x_{0}\right)$ we calculate, using Ito formula successively,

$$
\begin{aligned}
F(W(y(t)))=F(W(0)))+\int_{0}^{t} F^{\prime}(W(y(s))) v(s) \cdot d w(s)+ \\
\frac{1}{2} \int_{0}^{T}\left\{d F^{\prime}(W(y(s)))+|v(s)|^{2} F^{\prime \prime}(W(y(s)))\right\} d s
\end{aligned}
$$

Inserting $F(\lambda)=\lambda^{2}, \lambda \in \mathbb{R}$, we get the equation ii) of the lemma.

## Lemma3. 2

Let $Y$ be the solution of the stochastic differential equation (2.2). Then the configuration process $(x(t), t \geq 0)$ as well as the velocity process ( $v(t), t \geq 0)$ possess finite absolute moments of second order.

## Proof:

Starting from (2.8) we find using Lemma 3.1 i)

$$
\begin{align*}
\mathbb{E}\left(|v|^{2}\right) & \leq 2 \mathbb{E}(W(y(t)))+2 C_{k}  \tag{3.2}\\
& =2 W\left(y_{0}\right)+d t+2 C_{k}
\end{align*}
$$

Inserting equations (2.2) and (3.2) into the expression for the second moment of the configuration process then there holds:

$$
\begin{align*}
\mathbb{E}\left(|x|^{2}\right) & =\mathbb{E}\left(\left|\int_{0}^{t} v(s) d s\right|^{2}\right) \\
& \leq \int_{0}^{t} \mathbb{E}\left(|v(s)|^{2}\right) d s  \tag{3.3}\\
& \leq \frac{d}{2} t^{2}+2\left(W\left(y_{0}\right)+C_{k}\right) t
\end{align*}
$$

where $C_{k}$ is a positive constant.

## Lemma3. 3

Let $K, Y=((x(t), v(t)), o \leq t \leq T)$ and $\eta=((z(t), u(t)), u \leq t \leq T)$ be as in Theorem 2.1 and in (3.1) respectively. There holds

$$
\begin{gathered}
\mathbb{P}\left\{\int_{0}^{T}|K(x(t))+\gamma x|^{2} d s<\infty\right\}= \\
\mathbb{P}\left\{\int_{0}^{T}|K(z(t))+\gamma z(t)|^{2} d s<\infty\right\}=1
\end{gathered}
$$

## Proof:

As was shown in Lemma 3.2 the second absolute moment of the configuration process $X=$ $(x(t), t \geq 0)$ is finite (see (3.3)). Since $K$ is Lipschitz continuous and $X$ has continuous paths $\mathbb{P}$ a.e which are bounded on $[0, T]$ we have

$$
\begin{align*}
& \int_{0}^{t}|K(x(t))+\gamma x(t)|^{2} d t \leq 2 \int_{0}^{T}\left(|K(x(t))|^{2}+|\gamma x(t)|^{2}\right) d t \\
& \leq 2 \int_{0}^{T}\left(K_{T}^{2}\left|x(t)-x_{0}\right|^{2}+|\gamma|^{2}|x(t)|^{2}+\left|K\left(x_{o}\right)\right|^{2}\right) d t<\infty \tag{3.4}
\end{align*}
$$

where $K_{T}$, the Lipschitz constant of the sphere $B_{r_{0}}\left(x_{0}\right), r_{0}=\max _{t \in[0, t]}\left\{\left|x(t)-x_{0}\right|\right\}$, may depend on the specific path $\omega$. Since $\eta$ is a degenerate Gaussian process we find analogously $\mathbb{P}$ a.e

$$
\begin{align*}
& \int_{0}^{T}|K(z(t))+\gamma z(t)|^{2} d t \leq 2 \int_{0}^{T}\left(|K(z(t))|^{2}+|\gamma z(t)|^{2}\right) d t \\
& \leq 2 \int_{0}^{T}\left(\bar{K}_{T}^{2}\left|z(t)-z_{0}\right|^{2}+|\gamma|^{2}|z(t)|^{2}+\left|K\left(x_{o}\right)\right|^{2}\right) d t<\infty \tag{3.5}
\end{align*}
$$

where $\bar{K}_{T}^{2}$ the Lipschitz constant of the sphere $B_{s_{0}}\left(z_{0}\right), s_{0}=\max _{t \in[0, t]}\left\{\left|z(t)-z_{0}\right|\right\}$, may depend on $\omega$, and the second constant factor is defined as $|\gamma|^{2} \equiv \mid$ trace $\left.\gamma\right|^{2}$.

## Proposition 3.4

Let $T>0$, and ( $w_{t}, t \in[0, T]$ ) be a Wiener process with state space $\mathbb{R}^{n}$. Furthermore take $(\Omega, \mathcal{F}, \mathbb{P})$ as Wiener space, and $\mathcal{F}_{t} \equiv \sigma\left(w_{s}, s \in[0, T]\right)$ as the $\sigma$-algebra generated by the Wiener process up to time T. Let $\xi_{t}, \eta_{t}$ be stochastic processes on $\left(\Omega, \mathcal{F}, \mathcal{F}_{t}, \mathbb{P}\right)$ s.t. $\xi_{t}$ and $\eta_{t}$ are $\mathcal{F}_{t^{-}}$ measurable and assume $\xi_{t}, \eta_{t}$ satisfy the stochastic differential equations

$$
\begin{aligned}
& d \xi_{t}=A(t, \xi) d t+\sigma(t, \xi) d w_{t} \\
& d \eta_{t}=a(t, \eta) d t+\sigma(t, \eta) d w_{t},
\end{aligned}
$$

with $\eta_{0}=\xi_{0}$ being $\mathcal{F}_{0}$-measurable and

$$
\mathbb{P}\left\{\xi_{0}<\infty\right\}=1
$$

$A, a, \sigma$ should satisfy the following conditions:
i) $A(t, \xi), a(t, \eta), \sigma(t, \xi), \sigma(t, \eta)$ are $\mathcal{F}_{t}$-measurable and such that there exist unique strong solutions $\xi_{t}, \eta_{t}$ of the above stochastic differential equations(i.e. solutions on the probability space $(\Omega, \mathcal{F}, \mathbb{P})$ ).
ii) For arbitrary fixed $t \in[0, T]$ the system of algebraic equations in $x$ over $\mathbb{R}^{n}$ admits a solution $\alpha(t, x)$ :

$$
\sigma(t, x) \alpha(t, x)=A(t, x)-a(t, x) .
$$

The function $\alpha(\cdot, \cdot)$ should be measurable in $t \in[0, T], \quad x \in \mathbb{R}^{n}$, and satisfy
iii)

$$
\operatorname{Eexp}\left(\frac{1}{2} \int_{0}^{T}\left|\alpha\left(t, \xi_{t}\right)\right|^{2} d t\right)<\infty
$$

Let $\mu_{\xi}$ resp. $\mu_{\eta}$ be the probability measures on $(\Omega,, \mathcal{F}, \mathbb{P})$ associated with $\xi$ resp. $\eta$ in the sense that the probability of $\left\{\xi_{t_{1}} \in B_{1}, \ldots, \xi_{t_{n}} \in B_{n}\right\}$ is $\mu_{\xi}\left(\xi_{t_{1}} \in B_{1}, \ldots, \xi_{t_{n}} \in B_{n}\right)$ for all $B_{i} \in \mathcal{B}\left(\mathbb{R}^{n}\right)$, and all $t_{i} \in[0, T]$, and similarly for $\mu_{\eta}$ Then $\mu_{\xi}$ and $\mu_{\eta}$ are mutually absolutely continuous i.e. are equivalent and one has $\mathbb{P}$-almost surely

$$
\frac{d \mu_{\eta}}{d \mu_{\xi}}=e^{-\int_{0}^{T} \alpha(t, \xi) \cdot d w_{t}} \cdot e^{-\frac{1}{2} \int_{0}^{T}|\alpha(t, \xi)|^{2} d t}
$$

## Proof:

The proof is a multidimensional version of the statement made in [Lip] (Th. 5.4 p. 160) for $\mathbb{R}^{1}$.

## Theorem 3.5

Give the probability space $(\Omega, \mathcal{A} \otimes \mathcal{F}, \mathbb{P})$ as above. Let $Y=((x(t), v(t)), t \geq 0)$ with initial data $y_{0}$ be the global solution of the nonlinear stochastic differential equation (2.2) with $K$ satisfying the assumptions of the existence and uniqueness Theorem 2.1.
Let $\eta=((z(t), u(t)), t \geq 0)$ be the solution of the linear stochastic differential equation (3.1).
Then the process $\eta$ is equivalent to $Y$, in the sense that the probability measures $\mu_{\eta}$ and $\mu_{y}$ constructed on path space $\Omega$ are equivalent. The Radon-Nikodym derivatives are given by

$$
\frac{d \mu_{y}}{d \mu_{\eta}}(\eta)=\exp \left(+\int_{0}^{T}(K(z(t))+\gamma z(t)) \cdot d w(t)-\frac{1}{2} \int_{0}^{T}|K(z(t))+\gamma z(t)|^{2} d t\right)
$$

and

$$
\frac{d \mu_{\eta}}{d \mu_{y}}(y)=\exp \left(-\int_{0}^{T}(K(x(t))+\gamma x(t)) \cdot d w(t)-\frac{1}{2} \int_{0}^{T}|K(x(t))+\gamma x(t)|^{2} d t\right)
$$

## Proof:

For arbitrary fixed $t \in[0, T]$, and $a(y), \beta(y), \sigma$ as in Lemma 3.2 the system of algebraic equations in $x$ over $\mathbb{R}^{2 d}$ admits a measurable solution $\bar{\alpha}(x)$

$$
\begin{equation*}
\sigma \bar{\alpha}(y)=\beta(y)-a(y) . \tag{3.6}
\end{equation*}
$$

Denote by $\bar{\alpha}_{1}, \bar{\alpha}_{2}$ the first respectively last $d$ components of $\bar{\alpha}(x)$. $\bar{\alpha}_{1}$ is left undefined by (3.6), for convenience we choose $\vec{\alpha}_{1}(y)=0 . \bar{\alpha}_{2}$ is determined by (3.6) as

$$
\begin{equation*}
\bar{\alpha}_{2}(y) \equiv \alpha(x) \equiv K(x)+\gamma x . \tag{3.7}
\end{equation*}
$$

For any function $\kappa=\left(\kappa_{1}, \kappa_{2}\right), \kappa_{i} \in C\left([0, T], \mathbb{R}^{d}\right), i=1,2$, we define the exit time of $h(t, x):=$ $\int_{0}^{t}|\alpha(x(s))|^{2} d s$ in (3.5) from the sphere of radius $n$ by

$$
\vartheta_{n}\left(\kappa_{1}\right):= \begin{cases}\inf \{t \mid t \leq T\}, & h\left(t, \kappa_{1}\right) \geq n  \tag{3.8}\\ T & \text { otherwise }\end{cases}
$$

and set $\Theta_{n}(\kappa):=\vartheta_{n}\left(\kappa_{1}\right)$. The characteristic function $\chi_{n}(t, \kappa)=\chi_{\left\{\Theta_{n}(\kappa) \geq t\right\}}$ allows to define a truncated drift coefficient $A_{n}$ relating by a Girsanov transformation the solution of (3.12) below with (3.1). $A_{n}$ is defined by ( $3.9,3.10$ ):

$$
\begin{gather*}
K_{n}\left(\kappa_{1}\right)=-\left(\gamma \kappa_{1}\right)+\chi_{n}(t, \kappa)\left[K\left(\kappa_{1}\right)+\gamma \kappa_{1}\right],  \tag{3.9}\\
A_{n}(\kappa)=\binom{\kappa_{2}}{K_{n}\left(\kappa_{1}\right)} . \tag{3.10}
\end{gather*}
$$

Let us consider the process $Y^{(n)}(t)=\left(y_{t}^{(n)}, \mathcal{A}_{t} \otimes \mathcal{F}_{t} \mid 0 \leq t \leq T\right)$ defined by

$$
\begin{equation*}
y_{t}^{(n)}:=y_{t \wedge \Theta_{n}(y)}+\int_{0}^{t}\left[1-\chi_{n}\left(s, Y^{(n)}\right)\right]\left(-\gamma y^{(n)}(s)\right) d s+\int_{0}^{t}\left[1-\chi_{n}(s, Y)^{(n)}\right] \sigma d w_{s} \tag{3.11}
\end{equation*}
$$

By Proposition (2.8) equation (3.11) has a unique solution, with $y^{(n)}(t)=y(t)$ for $t \leq e(0)$. Applying Ito's calculus we find that $Y^{(n)}$ satisfies the following stochastic differential equation

$$
\begin{equation*}
d y^{(n)}(t)=A_{n}\left(y^{n}(t)\right) d t+\sigma d w_{t}, Y_{n, 0}=y_{0} . \tag{3.12}
\end{equation*}
$$

Using

$$
A_{n}(\kappa)-\binom{\kappa_{2}}{-\gamma \kappa_{1}}=\chi_{n}\left(t, \kappa_{1}\right)\left[K\left(\kappa_{1}\right)+\gamma \kappa_{1}\right]
$$

we deduce

$$
\left.\int_{0}^{t} \mid K_{n}\left(x^{(n)}\right)\right)+\left.\gamma x^{(n)}\right|^{2} \leq n \quad \mathbb{P} \text { a.e. }
$$

Moreover, the criterion for the Girsanov density to be a martingale given in Proposition (3.4) is fulfilled, i.e we have

$$
\begin{equation*}
\mathbb{E} \exp \left[-\int_{0}^{T}\left(K_{n}\left(x^{(n)}\right)+\gamma\left(x^{(n)}\right)\right) \cdot d w(t)-\frac{1}{2} \int_{0}^{T}\left|K_{n}\left(x^{(n)}(s)\right)+\gamma\left(x^{(n)}(s)\right)\right|^{2} d s\right]=1 \tag{3.13}
\end{equation*}
$$

and the following formula for the density holds:

$$
\begin{align*}
\frac{d \mu_{Y(n)}}{d \mu_{\eta}}(\eta) & =\exp \left[\int_{0}^{T}\left(K_{n}(\eta(s))+\gamma \eta(s)\right) \cdot d w(s) \cdot d w(s)-\frac{1}{2} \int_{0}^{T}\left|K_{n}(\eta(s))+\gamma \eta(s)\right|^{2} d s\right] \\
& =\exp \left[\int_{0}^{T \wedge \Theta_{n}(\eta)} \alpha(\eta(s)) \cdot d w(s)-\frac{1}{2} \int_{0}^{T \wedge \Theta_{n}(\eta)}|\alpha(\eta(s))|^{2} d s\right]  \tag{3.14}\\
& \equiv \zeta_{T \wedge \Theta_{n}(\eta)}(\eta) .
\end{align*}
$$

Let $\Gamma$ be a Baire set in the space $C\left([0, T], \mathbb{R}^{2 d}\right)$ Then we approximate $\mu_{Y}(\Gamma)$ by the density of the truncated process and rewrite the expression gained in this way by using (3.14)

$$
\begin{aligned}
\mu_{Y}(\Gamma) & =\lim _{n \rightarrow \infty} \mu_{Y^{(n)}}\left(\Gamma \cap\left\{\Theta_{n}(\kappa)=T\right\}\right) \\
& =\lim _{n \rightarrow \infty} \int_{\Gamma \cap\left\{\Theta_{n}(\kappa)=T\right\}} \zeta_{T \wedge \Theta_{n}(\kappa)}(\kappa) d \mu_{\eta}(\kappa) \\
& =\lim _{n \rightarrow \infty} \int_{\Gamma \cap\left\{\Theta_{n}(\kappa)=T\right\}} \zeta_{T}(\kappa) d \mu_{\eta}(\kappa) \\
& =\int_{\Gamma} \zeta_{T}(\kappa) d \mu_{\eta}(\kappa)
\end{aligned}
$$

This means that $\mu_{\xi}$ is absolutely continuous with respect to $\mu_{\eta}$, and the Radon-Nikodym derivative is given by $\zeta_{T}$.
Finally, let us point out that (3.5) also implies that the stochastic integral $\int_{0}^{T}(K(z(t))+\gamma z(t))$. $d w(t)$ exists and is finite ( $\mathbb{P}$ a.e), see e.g. [McKean], 2.3.6 (p. 25) or [Lip], Note 7 (pp. 104). Then

$$
\begin{equation*}
\mu_{\eta}\left\{\zeta_{T}(\kappa)=0 \mid \kappa \in \operatorname{supp} \mu_{\eta}\right\}=0, \tag{3.15}
\end{equation*}
$$

which ends the proof.

## Remark 3.6

All assumptions are satisfied e.g. in the following cases:
a) for some constants $\alpha_{0}, C_{1}, C_{2}>0$ :

$$
\begin{array}{ll}
d=1, & V(\alpha)=C_{1} \alpha^{2 q}+C_{2} \\
d>1, & V(\alpha)=C_{1}|\alpha|^{2 q}+C_{2}, \quad q \in \mathbb{N}, \forall|\alpha|^{2}>\alpha_{0}
\end{array}
$$

b) for $q \in \mathbb{N}$

$$
\begin{array}{ll}
d=1, & V(\alpha)=\alpha^{2 q} f(\alpha) \text { with } f \text { smooth and such that } \\
& \alpha f^{\prime}(\alpha)+2 q f(\alpha) \geq 0 \quad \forall|\alpha| \geq \alpha_{0}, \\
& V(\alpha) \text { Lipschitz for }|\alpha| \leq\left|\alpha_{0}\right|, \text { for some } \alpha_{0}>0 .
\end{array}
$$

## Remark 3.7

The restriction to initial condition $y_{0}=(0,0)$ is technical, and can be released since in the autonomous case all we need is a growth condition far away from the origin, see $[\mathrm{H}]$. Following [Mark-W] from the Girsanov-type stochastic equivalence result we can transfer some conclusions about the behaviour of the linear process in finite time to the nonlinear system. In particular we have:

Under the assumptions of Theorem 3.5

$$
x(t)^{2}+v(t)^{2}>0 \quad \forall t>0,
$$

provided

$$
x(0)^{2}+v(0)^{2}>0 \quad \text { almost surely.. }
$$

## 4. Some Additional Remarks

We can use the energy function defined in (2.15) to obtain some estimates on the behaviour of the solution process $Y$ of equation (2.2) with initial condition $y_{0}=0$ as $t \rightarrow \infty$. In fact we have

## Theorem 4.1

Under the assumptions of Theorem 2.1 the process $W(y(t))-d \frac{t}{2}, 0 \leq t<\infty$, is a martingale, and we have

$$
\left[W(0)+d \frac{t}{2}\right]^{n} \leq \mathbb{E}\left([W(y(t))]^{n}\right) \leq d^{n} \sum_{k=1}^{n} \prod_{l=1}^{k}(2 l-1)\binom{2 n}{2 k}[W(0)]^{n-k}\left(d \frac{t}{2}\right)^{k}
$$

## Proof:

As was pointed out in section 2 see (2.12), $W(y(t))-W(0)$ is the sum of a martingale and the nonrandom function $d \frac{t}{2}$.
a) Applying Ito's formula to $(W(y(t)))^{n}$, repeatedly, and using that the martingale $\int_{0}^{t} v(s)(W(y(s)))^{n-1} \cdot \mathrm{~d} w_{s}$ has expectation zero, we find taking expectation

$$
\begin{align*}
& \mathbb{E}\left(W(y(t))^{n}\right)-\mathbb{E}\left(W\left(y_{0}\right)^{n}\right)= \\
& \quad=\int_{0}^{t}\left(d \frac{n}{2} \mathbb{E}\left(W(y(s))^{n-1}\right)+\frac{n(n-1)}{2} \mathbb{E}\left(|v|^{2} W(y(s))^{n-2}\right)\right) \mathrm{d} s, \quad n \geq 2 \tag{4.1}
\end{align*}
$$

From (4.1) and

$$
0 \leq \mathbb{E}\left(|v(s)|^{2} W(y(s))^{n-2}\right) \leq 2 \mathbb{E}\left(W(y(s))^{n-1}\right)
$$

we have

$$
\begin{align*}
d \frac{n}{2} \int_{0}^{t} \mathbb{E}\left(W(y(s))^{n-1}\right) \mathrm{d} s & \leq \mathbb{E}\left(W(y(t))^{n}\right)-\mathbb{E}\left(W(0)^{n}\right)  \tag{4.2}\\
& \leq n\left(n-1+\frac{d}{2}\right) \int_{0}^{t} \mathbb{E}\left(W(y(s))^{n-1}\right) \mathrm{d} s \tag{4.3}
\end{align*}
$$

Moreover, one can easily see by induction that

$$
\begin{equation*}
\mathbb{E}\left(W(y(t))^{n}\right) \leq d^{n} \mathbb{E}\left((\sqrt{t d} z+\sqrt{W(0)})^{2 n}\right) \tag{4.4}
\end{equation*}
$$

where the expectation on the r.h.s. is with respect to standard Gaussian measure $z$ (with mean 0 and covariance $\frac{1}{2}$ ). Computing the expectation on the r.h. side in (4.4) yields the right estimate given in this theorem. The proof of the left inequality goes by induction. We integrate the inequality given by the assumption of the induction, and insert the resulting estimate into (4.2).

The process of Theorem 4.1 is a Markov diffusion process, since it solves the stochastic equation (2.2). The Markov kernel $P(t, a, \mathrm{~d} b), a, b \in \mathbb{R}^{2 d}$, defined by the transition probability is then well defined. Since $K(x)$ is continuous by our assumptions, $P(t, a, \mathrm{~d} b)$ defines a (Feller) Markov semigroup on $C_{b}\left(\mathbb{R}^{2 d}\right)$. Let $L$ be its infinitesimal generator s.t. for $f \in C_{0}^{\infty}\left(\mathbb{R}^{2 d}\right)$

$$
(L f)(a)=\left.\frac{d}{d t} \mathbb{E}^{a}(f(y(t)))\right|_{t=0}
$$

Using Ito's formula, see e.g. [Fr], [Si]:

$$
\begin{equation*}
(L f)(a)=\left(\Delta_{v}+K(x) \cdot \nabla_{v}+v \cdot \nabla_{x}\right) f(a) \tag{4.5}
\end{equation*}
$$

with $a \equiv(x, v)$. Following [ Po ] one can show that $P(t, a, \mathrm{~d} b)$ is absolutely continuous w.r. to Lebesgue-measure $\mathrm{d} b$ for fixed $t, a$. This is seen by looking at the transition probability kernels $P_{n}(t, a, d b)$ for the approximation of (2.2) obtained by replacing $K(x)$ by

$$
K_{n}(x)= \begin{cases}K(x) & \text { if }|x| \leq n \\ K(n) & \text { if }|x|>n\end{cases}
$$

By known results on the fundamental solution of degenerate parabolic equations with globally Lipshitz coefficients we have that $P^{n}(t, a, \mathrm{~d} b)=p^{n}(t, a, b) \mathrm{d} b$, with $p^{n}(t, a, b) \in$ $\in L_{\text {loc }}^{1}(d b)$. A dominated convergence argument shows that $P^{n}(t, a, A) \rightarrow P(t, a, A)$ and from $P^{n}(t, a, A)=0$ for $|A|=0$ follows that $P(t, a, A)=0$, hence the absolute continuity of $P(t, a, \mathrm{~d} b)$.

Let us regard $P(t, a, \mathrm{~d} b)$ as defining a Markov semigroup $T_{t}$ in the space $\mathcal{M}$ of signed measures with finite total variation, by defining for $\mu \in \mathcal{M}$

$$
\begin{equation*}
T_{t} \mu(\cdot) \equiv \int_{\mathbb{R}^{2 d}} P(t, a, \cdot) \mu(\mathrm{d} a) \tag{4.6}
\end{equation*}
$$

We call $\mu$ an invariant measure for the Markov semigroup $T_{t}$, or an invariant measure for the process $Y$ if

$$
\begin{equation*}
T_{t} \mu(A)=\mu(A) \tag{4.7}
\end{equation*}
$$

for any Borel subset $A$ of $\mathbb{R}^{2 d}$ and all $t \geq 0$. We shall see that under the above assumptions $P$ has a density $p$ w.r. to Lebesgue measure. We have for (4.4) by Fubini's Theorem

$$
\begin{equation*}
\mu(A)=\int_{\mathbb{R}^{2 d}} P(t, a, A) \mu(\mathrm{d} a)=\int_{A}\left(\int_{\mathbb{R}^{2 d}} p(t, a, y) \mu(\mathrm{d} a)\right) \mathrm{d} y \tag{4.8}
\end{equation*}
$$

where we used the absolute continuity of $P(t, a, \cdot)$ w.r. to Lebesgue measure. From (4.8) absolute continuity of the invariant measure follows.
At this point, we insert the following

## Remark 4.2

Let $\tilde{L}$ be the infinitesimal generator of the semigroup $T_{t}$, then any invariant measure $\mu$ of the process $Y$ of Theorem 2.1 satisfies

$$
\begin{equation*}
\tilde{L} \mu(\mathrm{~d} b)=0 \tag{4.9}
\end{equation*}
$$

and conversely.

## Lemma 4.3

The Lebesgue measure on $\mathbb{R}^{2 d}$ is an invariant measure for $Y$.

## Proof:

Let $\tilde{L}^{*}$ be the formal dual of the operator $\tilde{L}$ with $\mathcal{D}(L) \subset C_{b}\left(\mathbb{R}^{2 d}\right)$. For $P_{t}$ the dual of $T_{t}$ given by:

$$
\int_{\mathbb{R}^{2 d}}\left(P_{t} f\right)(b) \mu(\mathrm{d} b)=\int_{\mathbb{R}^{2 d}} f(b) T_{t} \mu(\mathrm{~d} b) \quad \forall f \in C_{0}^{2}\left(\mathbb{R}^{2 d}\right)
$$

we get by differentiating w.r. to $t$ :

$$
\int_{\mathbb{R}^{2 d}} f(b) \tilde{L} \mu(\mathrm{~d} b)=\int_{\mathbb{R}^{2 d}} L f(b) \mu(\mathrm{d} b)
$$

hence, on $C_{0}^{2}\left(\mathbb{R}^{2 d}\right)$, there holds

$$
L=(\tilde{L})^{*}
$$

This implies using the special form (4.5) of $L$ :

$$
\begin{equation*}
\tilde{L}=\Delta_{v}-K(x) \cdot \nabla_{v}-v \cdot \nabla_{x} . \tag{4.14}
\end{equation*}
$$

In particular, applied to Lebesgue measure $\lambda$ this yields

$$
\begin{equation*}
\tilde{L} \lambda=0 . \tag{4.11}
\end{equation*}
$$

By Remark $4.2 \lambda$ is then an invariant measure of the process $Y$ solving (2.1) (or (2.2)).

## Remark 4.4

One verifies easily that, if $K$ is $C^{\infty}, L$ is hypoelliptic in the sense that it has the form

$$
\begin{equation*}
L=X_{1}^{2}+X_{0} \tag{4.12}
\end{equation*}
$$

with $X_{i_{1}}=\frac{\partial}{\partial v_{i}}, X_{i_{0}}=K_{i}(x) \cdot \frac{\partial}{\partial v_{i}}+v_{i} \cdot \frac{\partial}{\partial x_{i}}, 1 \leq i \leq d$, so that $\left\{X_{i_{1}},\left[X_{i_{1}}, X i_{0}\right] \mid \quad 1 \leq i \leq d\right\}$ span the smooth vector fields over $\mathbb{R}^{2 d}$.
For additional results see [AHZ], [H].
Obviously many open problems remain, to name only one of them: is the solution process null recurrent (cfr. [Bat])?

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# STABILITTY OF GRONN STAIES FOR NONIINEAR CLASSICAL FIELD THEORIES 

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## O. Introduction

In this contribution we want to describe a stability theory for solitary waves of Hamiltonian systems which arise from many models in classical field theories. Most classical field theoretic models have common properties, e.g. scaling behaviour, existence of a ground state etc. [4]. Our abstract version will contain these ingredients. This will enable us to extend previous results by the authors $[2,3]$ and to present an alternative approach to the theory of Grillakis, Shatah and Strauss [7]. To illustrate the general context let us start with an well known example namely the nonlinear logarithmic schrödinger equation. on $\mathbb{R}^{n} n \geq 3$ we consider the equation
(NLS)

$$
i \phi_{t}+\Delta \phi+f(\phi)=0
$$

where $f(\phi)=\phi \log |\phi|^{2}$. The associated energy is given by

$$
\begin{equation*}
E(\phi)=K(\phi)-V(\phi) \tag{0.1}
\end{equation*}
$$

with
(0.2a)

$$
\begin{align*}
& K(\phi)=\frac{1}{2} \int_{\mathbb{R}^{N}}|\nabla \phi|^{2} d x \\
& V(\phi)=\frac{1}{2} \int_{\mathbb{R}^{N}}|\phi|^{2}\left(-1+\log |\phi|^{2}\right) d x \tag{0.2b}
\end{align*}
$$

$E(\phi)$ is, at least formally, conserved for solutions of (NLS) . $\mathrm{E}(\phi)$ is well-defined in the class

$$
\begin{equation*}
W \equiv H^{1}\left(\mathbb{R}^{n}\right) \cap\left\{\left.\phi \in L_{l o c}^{1}| | \phi\right|^{2} \log |\phi|^{2} \in L^{1}\right\} \tag{0.3}
\end{equation*}
$$

It was proven by Cazenave [5] that $W$ is a reflexive Banach space (so-called Sobolev-Orlicz space) and that $E$ is of class $C^{1}$ on $W$. See also [13].

For the moment we restrict ourselves to the subspace of radial functions in W , i.e. we consider only functions in
(0.4)

$$
X \equiv W_{r}=\{\phi \in W, \phi \text { radial }\}
$$

We can write equation (NLS) also in the standard Hamiltonian form

$$
\begin{equation*}
\frac{d \phi}{d t}=J E^{\prime}(\phi) \quad \text { on } \quad X \tag{H}
\end{equation*}
$$

where $J$ is simply the multiplication by (-i) . This Hamiltonian system is invariant under gauge transformations

$$
\begin{equation*}
U(s) \phi=e^{i s} \phi \quad, s \in \mathbb{R} \tag{0.5}
\end{equation*}
$$

The invariance implies the conservation of the charge

$$
\begin{equation*}
Q(\phi)=-\frac{1}{2} \int_{\mathbb{R}^{n}}|\phi|^{2} d x \tag{0.6}
\end{equation*}
$$

which is well-defined on $X$.

We are interested in solitary wave solutions of (H) of the form
(0.7) $\phi(t, x)=U(\omega t) \phi_{i \omega}(x) \quad \omega$ fixed.

They satisfy the "stationary" equation
(0.8) $-\Delta \phi_{\omega}=\phi_{\omega} \log \left|\phi_{\omega}\right|^{2}-\omega \phi_{\omega}$
which can be equivalently written in the form $E^{\prime}(\phi)=\omega Q^{\prime}(\phi)$.

The existence of nontrivial solutions of (0.8) can be proved by variational methods as follows:

Associated to the stationary problem we define the "action" functional

$$
\begin{align*}
I_{\omega}(\phi) & \equiv E(\phi)-\omega Q(\phi)  \tag{0.9}\\
& =\mathrm{K}(\phi)-\mathrm{V}_{\omega}(\phi)
\end{align*}
$$

where $V_{\omega}=V+\omega Q$. In particular we are interested in solutions $\phi_{\omega}$ which have least action among all possible nontrivial solutions of the stationary equation. Such solutions are called ground states.

To obtain ground state solutions one solves the following constrained minimization problem (see e.g. [1,9]):
(0.10)

$$
I(\omega)=\inf \left\{K(\phi), \phi \in X, V_{\omega}(\phi)=1\right\}
$$

The fact that a solution of this problem can be transformed into a solution of the field equation relies essentially on the nice behaviour of $K$ and $V_{\omega}$ under the action of a scaling group $T_{\sigma} \cdot$

Indeed, let

$$
\begin{equation*}
T_{\sigma} \phi(x)=\phi(x / \sigma) \tag{0.11}
\end{equation*}
$$

i.e. $T_{\sigma}$ is a representation of the group of dilations, then we have
(0.12a)

$$
\begin{aligned}
\mathrm{K}\left(\mathrm{~T}_{\sigma} \phi\right) & =\sigma^{\mathrm{n}-2} \mathrm{~K}(\phi) \\
\mathrm{V}_{\omega}\left(\mathrm{T}_{\sigma} \phi\right) & =\sigma^{\mathrm{n}} \mathrm{~V}_{\omega}(\phi)
\end{aligned}
$$

This property of $K$ and $V_{\omega}$ will be called "scale covariance".
In the case of the logarithmic Schrödinger equation the ground state is explicitly known and given by
(0.13a) $\quad \phi_{\omega}(x)=\phi_{o} \exp (-\omega / 2)$
and
(0.13b)

$$
\phi_{0}(x)=\exp \left(N / 2-1 / 2 x^{2}\right)
$$

Now we are interested in the Liapunov stability of the $\phi_{\omega}$-orbit $\left\{U(s) \phi_{\omega}, s \in \mathbb{R}\right\}$ in $X$.

For the logarithmic schrödinger equation the result is well-known: The first proof of orbital stability was given by Cazenave [5] using compactness methods. Recently the same result was obtained by Blanchard, Stubbe and Vazquez [2 ] who extended the methods of two earlier papers by Shatah and Strauss [11,12].

The most general approach to the stability problem has been given very recently by Grillakis, Shatah and Strauss [7] who studied the stability of solitary waves of Hamiltonian systems in a real Banach space. The main tool of their theory is the linearized Hamiltonian

$$
\begin{equation*}
H_{\omega} \equiv E^{\prime \prime}\left(\phi_{\omega}\right)-\omega Q^{\prime \prime}\left(\phi_{\omega}\right) \tag{0.14}
\end{equation*}
$$

The idea is to show that the Liapunov stability of the $\phi_{\omega}$-orbit is equivalent to the fact that $\phi_{\omega}$ minimizes locally the energy $E$ subject to constant charge $\phi$. Moreover these properties are equivalent to the condition that the action of the solitary wave $L_{\omega}\left(\phi_{\omega}\right)$ considered as a function of $\omega$ is convex. Unfortunately this theory is not applicable to the logarithmic schrödinger equation since the energy functional is not twice differentiable as pointed out in [ 2 ] and thus the linearized operator $H_{\omega}$ does not exist. Our main goal is to present a stability theory for ground state solitary waves of Hamiltonian systems for which the linearized Hamiltonian does not exist.

In Section 1 we describe our abstract framework and state the main results. In Section 2 we give the outline's of the proofs. Technical details are omitted. They will be published elsewhere. In Section 3 we present some examples and in Section 4 we discuss some extensions, in particular to systems invariant under more than one symmetry.

1. The abstract model

Let $X$ be a real Banach space. We consider the following Hamiltonian system on $X$

$$
\begin{equation*}
\frac{d u}{d t}=J^{\prime} E^{\prime}(u) \tag{H}
\end{equation*}
$$

where $J: X{ }^{*} \xrightarrow{\text { onto }} X$ is a skew-symmetric linear operator from $X^{*}$ onto $X$ and $E: X \rightarrow \mathbb{R}$ is a $C^{1}$ functional on $X$.

We are now in a position to list our main assumptions:

### 1.1 Properties of the energy $E$

$E$ can be splitted into two scale covariant parts $K$ and $V$ i.e. there exists a continuous mapping

$$
\begin{equation*}
\sigma \longmapsto \mathrm{T}_{\sigma} \in L(X, X) \quad, \quad \sigma \in \mathbb{R}^{+} \tag{1.1}
\end{equation*}
$$

such that

$$
\begin{equation*}
T_{\sigma_{1}} T_{\sigma_{2}}=T_{\sigma_{1} \sigma_{2}}, \quad T_{1}=I d_{X} \tag{1.2}
\end{equation*}
$$

and there exist $C^{1}$ functionals $K$ and $V$ such that

$$
\begin{equation*}
E(u)=K(u)-V(u) \tag{1.3}
\end{equation*}
$$

and
(1.4a)

$$
K\left(T_{\sigma} u\right)=\sigma^{r} K(u)
$$

$$
\begin{equation*}
V\left(T_{\sigma} u\right)=\sigma^{s} V(u) \tag{1.4b}
\end{equation*}
$$

with $0<r<s$, i.e. $\gamma \equiv r / s \in(0,1)$.
$K$ and $V$ (and hence $E$ ) are invariant under a strongly continuous one parameter group of isometries $U(s): X \longrightarrow X, U(s+r)=$ $=U(s) U(r)$, i.e.
(1.5a)

$$
K(U(s) u)=K(u)
$$

(1.5b)

$$
V(U(s) u)=V(u)
$$

Associated to this invariance there is another conserved quantity

$$
\begin{equation*}
Q(u)=\frac{1}{2}\langle B u, u\rangle \tag{1.6}
\end{equation*}
$$

where $B: X \rightarrow X^{*}$ is symmetric and $J B$ is an extension of the infinitesimal generator $U^{\prime}(0)$. <, > denotes the dual pairing between $X$ and $\mathrm{x}^{*}$.

### 1.2 Property of the 'charge' $Q$

We assume that $Q$ has the following behaviour under the action of the scaling group $\mathrm{T}_{\sigma}$.

$$
\begin{equation*}
Q\left(\mathbb{T}_{\sigma} u\right)=\sigma^{q} Q(u) \quad \text { with } q=r \text { or } q=s \tag{1.7}
\end{equation*}
$$

We define

$$
\begin{align*}
& \mathrm{K}_{\omega}(\mathrm{u}) \equiv \mathrm{K}(\mathrm{u})-\omega \delta_{r q} Q(u)  \tag{1.8a}\\
& \mathrm{v}_{\omega}(\mathrm{u}) \equiv \mathrm{V}(\mathrm{u})+\omega \delta_{q s} Q(\mathrm{u}) \tag{1.8b}
\end{align*}
$$

and assume $K_{\omega}(u)>0$ if $u \neq 0$ and $K_{\omega}(0)=V_{\omega}(0)=0$.
Furthermore we define

$$
\begin{align*}
L_{\omega}(u) & \equiv E(u)-\omega Q(u)  \tag{1.9}\\
& =K_{\omega}(u)-V_{\omega}(u) .
\end{align*}
$$

1.3 The minimization principle

We assume that there exist $\omega_{1}<\omega_{2}$ such that for any $\omega \in\left(\omega_{1}, \omega_{2}\right)$ the problem

$$
\begin{equation*}
I(\omega)=\inf \left\{K_{\omega}(u) \mid u \in x, V_{\omega}(u)=1\right\} \tag{1.10}
\end{equation*}
$$

has a unique solution $\tilde{\phi}_{\omega}$, i.e.

$$
\mathrm{K}_{\omega}\left(\tilde{\Phi}_{\omega}\right)=\mathrm{I}(\omega) \quad, \quad \mathrm{V}_{\omega}\left(\tilde{\phi}_{\omega}\right)=1
$$

in the sense that for any minimizing sequence $\left(u_{n}\right)_{n}$ (i.e. $V\left(u_{n}\right)=1$ and $\left.\frac{\lim _{n \rightarrow \infty}}{n} K_{\omega}\left(u_{n}\right)=I(\omega)\right)$ there exists a subsequence $\left(u_{n_{j}}\right)$ and a sequence of real numbers $\left(s_{n_{j}}\right)$ such that

$$
\begin{equation*}
U\left(s_{n_{j}}\right) u_{n_{j}} \longrightarrow \tilde{\phi}_{\omega} \text { in } x \text { as } n_{j} \longrightarrow \infty \tag{1.11}
\end{equation*}
$$

This means that any minimization sequence for (1.10) tends to the orbit $\left\{U(s) \tilde{\Phi}_{\omega}, s \in \mathbb{R}\right\}$.

If we define $\phi_{\omega} \equiv \mathrm{T} \tilde{\sigma}(\omega) \tilde{\phi}_{\omega}$ with $\tilde{\sigma}(\omega)=(\gamma I(\omega))^{1 / \mathrm{s}}$ then $\phi_{\omega}$ is a ground state of the stationary equation

$$
\begin{equation*}
L_{\omega}^{\prime}\left(\phi_{\omega}\right)=0 \tag{1.12}
\end{equation*}
$$

and $u(t)=U(\omega t) \phi_{\omega}$ is a solitary wave solution of (H).
A consequence of the scaling property is the following virial theorem for $\phi_{\omega}$.

$$
\begin{equation*}
\gamma K_{\omega}\left(\phi_{\omega}\right)=V_{\omega}\left(\phi_{\omega}\right) . \tag{1.13}
\end{equation*}
$$

We define

$$
\begin{equation*}
d(\omega) \equiv I_{\omega}\left(\phi_{\omega}\right) \tag{1.14}
\end{equation*}
$$

then $d(\omega)$ satisfies $d(\omega)=(1-\gamma) K_{\omega}\left(\phi_{\omega}\right)>0$ by (1.13) and $d^{\prime}(\omega)=-Q\left(\phi_{\omega}\right)$.

Furhtermore we have the following alternative characterizations of d(w):

$$
\begin{align*}
& d(\omega)=\inf \left\{L_{\omega}(v) \mid v \neq 0, \gamma K_{\omega}(v)-v_{\omega}(v) \leq 0\right\}  \tag{1.15a}\\
& d(\omega)=\inf \left\{L_{\omega}(v) \mid K_{\omega}(v)=K_{\omega}\left(\phi_{\omega}\right)\right\} .
\end{align*}
$$

(1.15b)

Definition: The $\phi_{\omega}$-orbit $\left\{U(\omega t) \phi_{\omega}, t \in \mathbb{R}\right\}$ is stable if for all $\varepsilon>0$ there exists a $\delta>0$ such that if $\left\|u_{o}-\phi_{\omega}\right\|<\delta$ and $u(t)$ is a solution of (H) with $u(0)=u_{0}$ existing for any $t \geq 0$ and

$$
\sup _{t>0} \inf _{s \in \mathbb{R}}\left\|u(t)-U(s) \phi_{\omega}\right\|<\varepsilon .
$$

Otherwise we call the $\phi_{\omega}$-orbit unstable.
Before describing our results it will be convenient first to consider a one-dimensional mechanical analogue of our abstract model (see also [10] ).

Qualitatively $\mathrm{I}_{\omega}(\omega)$ has the same form as the energy for the one-dimensional motion in a potential well. In our context $\mathrm{L}_{\omega}$ turns out to be the energy of the 'modulated' Hamiltonian system
$\left(\mathrm{H}_{\text {mod }}\right) \quad \frac{\mathrm{dv}}{\mathrm{dt}}=\mathrm{J} \mathrm{L}_{\omega}^{\prime}(\mathrm{v})$
which we obtain from (H) by the transformation $u(t)=U(\omega t) v(t)$ where $u(t)$ is a solution of (H).

For arbitrary fixed nonzero $u \in X$ we consider the function

$$
\begin{aligned}
l(\sigma) & =L_{\omega}\left(T_{\sigma} u\right) \\
& =\sigma^{r} K_{\omega}(u)-\sigma^{s} V_{\omega}(u) .
\end{aligned}
$$

If $\mathrm{V}_{\omega}(\mathrm{u}) \leq 0$ then $l(\sigma)$ is increasing and goes to infinity as $\sigma \rightarrow \infty$.

If $V_{\omega}(u)>0$ then $1(\sigma)$ as a unique maximum $\sigma^{*}=\sigma^{*}(u)=\left(\gamma K_{\omega}(u) /\right.$ $\left.\mathrm{v}_{\omega}(\mathrm{u})\right)^{1 / s-r}$ and $\quad 1(\sigma) \rightarrow \infty$ as $\sigma \rightarrow \infty$.

We calculate the height

$$
\begin{equation*}
h_{\omega}(u) \equiv I_{\omega}\left(T_{\sigma^{*}} u\right)=(1-\gamma) \sigma^{*} K_{\omega}(u)>0 . \tag{1.16}
\end{equation*}
$$

We now define the lowest height to be passed by

$$
d(\omega)=\inf _{u \neq 0} h_{\omega}(u)
$$

Equivalently, if we always normalize $u$ so that $\sigma^{*}=1$, i.e. $\gamma K_{\omega}(u)-v_{\omega}(u)=0$, we recover the variational characterization (1.15a).

Hence $\phi_{\omega}$ lies on the 'mountain pass' of the energy mountains when travelling from zero to regions far away. The height of the mountain pass is exactly $d(\omega)$. This situation is sketched in the figure below.


Having in mind this picture one has the most important informations necessary for the proofs of our main results:

Fix $\omega_{0} \in\left(\omega_{1}, \omega_{2}\right)$ and let $M_{\omega_{0}} \equiv\left\{u \in X \mid Q(u)=Q\left(\phi_{\omega_{0}}\right)\right\}$.
Theorem 1: Let $d^{\prime \prime}\left(\omega_{0}\right)>0$. Then the $\phi_{\omega_{0}}$-orbit is stable. In particular we have $Q\left(\phi_{\omega_{0}}\right) \neq 0$ and $\phi_{\omega_{0}}$ is the local minimum of $\left.E\right|_{\mu_{\omega_{0}}}$. Theorem 2: Let $d^{\prime \prime}\left(\omega_{0}\right)<0$. Then $\left.E\right|_{M}$ is not locally minimized at $\phi_{\omega_{0}}$ and the $\phi_{\omega_{0}}$-orbit is unstable ${ }^{\omega} .0$

Furthermore we have the following particular cases:
Theorem 3: If $Q\left(\phi_{\omega_{0}}\right)=0$ then the $\phi_{\omega_{0}}$-orbit is unstable and $\phi_{\omega_{0}}$ is not a local minimum of $\left.E\right|_{\omega_{0}}$.

Theorem 4: If no symmetry $U$ exists then any ground state $\phi_{o}$ of $E^{\prime}(\phi)=0$ is unstable. $\phi_{o}$ is not a local minimum of the energy. Theorems 3/4 can be interpreted as abstract versions of Derrick's theorem [14] .

## 2. Outlines of the proofs

First of all we prove the following intermediate results.

## Theorem 2.1:

a) Let $\phi_{\omega}$ be a charged ground state $\left(Q\left(\phi_{\omega}\right) \neq 0\right)$. Then the energy $E$ has a local minimum in $M_{\omega_{0}}$ at $\phi_{\omega_{0}}$ if and only if $d(\omega)$ is convex at $\omega_{0}$.
b) Let $Q\left(\phi_{\omega_{0}}\right)=0$. Then $\phi_{\omega_{0}}$ is not a local minimum of $\left.E\right|_{\mu_{\omega}}$.
c) A ground state $\phi_{o}$ of $E^{\prime}(\phi)=0$ is not a local minimum of the energy.

## Proof:

a) Let $d$ be convex at $\omega_{0}$. For $u \in M_{\omega_{0}}$ sufficiently close to $\phi_{\omega}$ there exists $\omega$ such that $K_{\omega}(u)^{\circ}=K_{\omega}\left(\phi_{\omega}\right)=(1-\gamma) d(\omega)$ since $d^{\prime}\left(\omega_{0}\right)=-Q\left(\phi_{\omega_{0}}\right) \neq 0$. Thus

$$
\begin{aligned}
E(u) & =L_{\omega}(u)+\omega Q(u) \\
& \geq d(\omega)-\omega d^{\prime}\left(\omega_{0}\right) \\
& \geq d\left(\omega_{0}\right)-\omega_{0} d^{\prime}\left(\omega_{0}\right) \quad \text { by the convexity of } d \\
& =E\left(\phi_{\omega_{0}}\right) \cdot
\end{aligned}
$$

For the opposite direction we define $\psi_{\omega} \equiv T_{\sigma(\omega)} \phi_{\omega}$ such that $\psi_{\omega} \in M_{\omega_{0}}$. Then

$$
\begin{aligned}
d\left(\omega_{0}\right)-\omega_{0} d^{\prime}\left(\omega_{0}\right) & =E\left(\phi_{\omega_{0}}\right) \leq E\left(\psi_{\omega}\right) \\
& =\sigma^{r} K_{\omega}\left(\phi_{\omega}\right)-\sigma^{s} V_{\omega}\left(\phi_{\omega}\right)+\omega Q\left(\psi_{\omega}\right) \\
& =\left(\sigma^{r}-\gamma \sigma^{s}\right) K_{\omega}\left(\phi_{\omega}\right)-\omega d^{\prime}\left(\omega_{o}\right) \text { by the } \\
& \leq(1-\gamma) K_{\omega}\left(\phi_{\omega}\right)-\omega d^{\prime}\left(\omega_{o}\right) \text { theorem (1.13) } \\
& =d(\omega)-\omega d^{\prime}\left(\omega_{0}\right) \quad \text { by }(1.14) .
\end{aligned}
$$

For b) and c) we choose the curve $\psi_{\sigma}=T_{\sigma} \phi_{\omega_{o}}$ resp. $\psi_{\sigma}=T_{\sigma} \phi_{\sigma}$.
Now we sketch the stability proof, which is a generalisation of the method presented by Shatah [11] for nonlinear Klein-Gordon equations:

Let $d "\left(\omega_{0}\right)>0$. For arbitrary fixed $\omega$ we consider the modulated system
$\left(H_{\bmod }\right) \quad \frac{d v}{d t}=J L_{\omega}^{\prime}(v) \quad$.
Then it can be shown that

$$
\begin{align*}
& \mathbb{R}_{\omega}^{1} \equiv\left\{u \in X \mid L_{\omega}(u)<d(\omega), \gamma K_{\omega}(u)-V_{\omega}(u)>0\right\} U\{0\}  \tag{2.1.1}\\
& \mathbb{R}_{\omega}^{2}=\left\{u \in X \mid L_{\omega}(u)<d(\omega), \gamma K_{\omega}(u)-V_{\omega}(u)<0\right\} \tag{2.1.2}
\end{align*}
$$

are invariant regions under the flow of ( $\mathrm{H}_{\text {mod }}$ ).
This fact can be understood if one has in mind the one-dimensional analogues presented in Section 1. If the energy $L_{\omega}$ is less then the height of the mountain pass one can never cross it. Here $\mathbb{R}_{\omega}^{1}$ represents the region inside the well and $\mathbb{R}_{\omega}^{2}$ represents the exterior region.


Furthermore we can characterize these regions as follows:

$$
\begin{equation*}
R_{w}^{1 / 2}=\left\{u \in X\left|L_{\omega}(u)<d(w)\right|(1-\gamma) K_{\omega}(u)<d(w)\right\} . \tag{2.2}
\end{equation*}
$$

Now for every $\varepsilon>0$ small enough there exists a $t>0$ such that if $\left\|u-\phi_{\omega_{0}}\right\|<\delta$ then

$$
\begin{equation*}
d\left(\omega_{+}\right)<(1-\gamma) k_{\omega_{0}}(u)<d\left(\omega_{-}\right) \text {if } Q\left(\phi_{\omega_{0}}\right)>0 \tag{2.3}
\end{equation*}
$$ or the reversed inequality if $Q\left(\phi_{\omega_{0}}\right)<0$, where $\omega_{ \pm}=\omega \neq \varepsilon$ In addition $v_{ \pm}$defined by $u=U\left(\omega_{ \pm} t\right) v_{ \pm}$satisfies

$$
\begin{equation*}
\mathrm{L}_{\omega_{ \pm}}\left(\mathrm{v}_{ \pm}\right)<d\left(\omega_{ \pm}\right) \tag{2.4}
\end{equation*}
$$

which follows from the strict convexity of $d$ in a neighborhood of $\omega_{0}$.

Assume that $\phi_{\omega}$ is unstable. Then there exists a sequence of initial data $u_{n}(0) \xrightarrow{\omega_{0}} \phi_{\omega_{0}}$ in $x$ and $\delta>0$ and $t_{n}>0$ such that

$$
\begin{equation*}
\left\|u_{n}\left(t_{n}\right)-\phi_{\omega_{0}}\right\| \geq \delta \tag{2.5}
\end{equation*}
$$

Applying (2.3) we conclude

$$
(1-\gamma) K_{\omega_{0}}\left(u_{n}\left(t_{n}\right)\right) \xrightarrow{n \rightarrow \infty} d\left(\omega_{0}\right)
$$

and (2.4) then implies

$$
L_{\omega_{0}}\left(u_{n}\left(t_{n}\right)\right) \xrightarrow{n \rightarrow \infty} d^{1} \leq d\left(\omega_{0}\right)
$$

But now the assumptions on the minimization problem imply that there exists a subsequence of $\left(u_{n}\left(t_{n}^{*}\right)\right)$ which tends to the $\phi_{\omega_{0}}$-orbit contradicting (2.5).

The following figure illustrates the proof given above:


The instability proof first of all will be sketched for the case where we have no symmetry:

By Theorem 2.1 we know that $\phi_{o}$ (the ground state of $E^{\prime}(u)=0$ ) is a local maximum along the curve $\psi_{\sigma}=T_{\sigma} \phi_{\sigma}$.

We define $y=\left.\frac{d}{d \sigma}\right|_{\sigma=1} T_{\sigma} \phi_{O}$. There exists $Y \in X^{*}$ such that $J Y=Y$. We define a linear functional on $X$ by

$$
\begin{equation*}
A(u) \equiv-\langle Y, u\rangle \tag{2.6}
\end{equation*}
$$

Let $u_{o} \in U_{\varepsilon}\left(\phi_{o}\right)$. It suffices to show that $\frac{d A(u(t))}{d t}$ is bounded away from zero as long as $u(t) \in U_{\varepsilon}\left(\phi_{o}\right)$.

The trajectories along the vector field - J A'(u) are given by

$$
\begin{equation*}
\mathbb{R}(\lambda, v)=v+\lambda y \tag{2.7}
\end{equation*}
$$

Then there exists a functional $\Lambda: U_{\varepsilon} \longrightarrow \mathbb{R}$ such that

$$
\begin{equation*}
K(R(\Lambda(v), v))=K\left(\phi_{o}\right) \tag{2.8}
\end{equation*}
$$

Doing a Taylor expension for the energy and using the minimum property of $\phi_{o}$ we obtain for any $v \in U_{\varepsilon}\left(\phi_{o}\right)\left(v \neq \phi_{o}\right)$

$$
\begin{equation*}
E\left(\phi_{0}\right)<E(v)+\Lambda(v) P(v) \tag{2.9}
\end{equation*}
$$

where $P(v)=\left\langle E^{\prime}(v), y\right\rangle$. Using $\frac{d A(u(t))}{d t}=P(u(t))$ as long as $u(t) \in U_{\varepsilon}$ we easily conclude.

In the presence of a symmetry we proceed in a similar way. Let $d^{\prime \prime}\left(\omega_{0}\right)<0 \quad\left(\right.$ or $\left.Q\left(\phi_{\omega_{0}}\right)=0\right)$.
The neighborhood $U_{\varepsilon}$ is replaced by the tube

$$
u_{\varepsilon} \equiv\left\{u \in x \mid \inf _{S \in \mathbb{R}}\left\|u-U(s) \phi_{\omega_{0}}\right\|<\varepsilon\right\}
$$

We decompose

$$
X=\left\langle U^{\prime}(0) \phi_{\omega_{0}}\right\rangle \oplus \tilde{X}
$$

There exists a functional $s(u): U_{\varepsilon} \longrightarrow \mathbb{R}$ such that $U(s(u)) u-$ $-\phi_{\omega_{0}} \in \widetilde{X}$. Then we define

$$
\begin{equation*}
A(u)=-\left\langle Y_{\omega_{0}}, \mathrm{~J}(s(u)) u\right\rangle \tag{2.10}
\end{equation*}
$$

for any $u \in u_{\varepsilon} \cap M_{\omega_{0}}$ where $Y_{\omega_{0}}=\left.J^{-1} \frac{d}{d \omega}\right|_{\omega=\omega_{0}} \psi_{\omega}$ with $\psi_{\omega}$ the
curve given in the proof of theorem 2.1. Up to simple modification the proof of instability will now run along the same lines as above .

## 3. Applications

3.1. We start with the logarithmic Schrödinger equation

$$
\begin{equation*}
i \phi_{t}+\Delta \phi+\phi \log |\phi|^{2}=0 \quad \text { on } \mathbb{R}^{n}, n \geq 3 . \tag{3.1}
\end{equation*}
$$

As mentioned in the introduction the ground state with frequency $\omega$ is of the form

$$
\phi_{\omega}(x)=\phi_{0}(x) \exp (-\omega / 2)
$$

Hence

$$
\begin{equation*}
d(\omega)=\frac{1}{N} \exp (-\omega) K\left(\phi_{0}\right) \tag{3.2}
\end{equation*}
$$

with $K$ given by ( 0.2 a ) and therefore all ground states.are stable.
3.2. Consider the following nonlinear Schrödinger equation with a nonlocal nonlinearity

$$
\begin{equation*}
i \phi_{t}+\Delta \phi+\left(\int_{\mathbb{R}^{N}} v(x-y)|\phi(y)|^{2} d y\right) \phi=0 \tag{3.3}
\end{equation*}
$$

As scaling group we take $T_{\sigma} \phi=\sigma \phi$. We define

$$
\begin{align*}
& \mathrm{K}_{\omega}(\phi)=\frac{1}{2} \int_{\mathbb{R}^{\mathbb{N}}}|\nabla \phi|^{2}+\omega|\phi|^{2}, \quad \omega>0  \tag{3.4a}\\
& \mathrm{~V}_{\omega}(\phi)=\frac{1}{4} \int_{\mathbb{R}^{\mathbb{N}} \mathbb{R}^{\mathbb{N}}} \mathrm{V}(\mathrm{x}-\mathrm{y})|\phi(\mathrm{y})|^{2}|\phi(\mathrm{x})|^{2} \mathrm{dxdy} \tag{3.4b}
\end{align*}
$$

under suitable assumptions on $V$ the assumptions of Section 1 are valid.

Consider, e.g. equation (3.3) on $\mathbb{R}^{3}$ with $V(x)=\frac{1}{|x|}$. This is the so-called Pekard-Choquard equation. The minimization problem was solved by Lieb [8]. Existence of solutions for the cauchy problem in $H_{r}^{1}\left(\mathbb{R}^{3}\right)$ was proved by Ginibre and Velo [6].

Using the scaling properties of the stationary Pekard-Choquard we see that

$$
\psi(x)=\omega^{-1} \phi_{\omega}\left(\omega^{-1 / 2} x\right)
$$

is independent of $\omega$. Hence

$$
\begin{equation*}
a(\omega)=\omega^{3 / 2} d(1) \tag{3.5}
\end{equation*}
$$

and therefore all ground states are stable. The same stability result was obtained by Cazenave and Lions [ 15] using the concentration compactness principle.

### 3.3. Consider the logarithmic Klein-Gordon equation

$$
\begin{equation*}
\phi_{t t}-\Delta \phi=\phi \log (\phi)^{2} \quad \text { on } \mathbb{R}^{N}, \quad n \geq 3 \tag{3.6}
\end{equation*}
$$

We study this problem in the space

$$
X=W_{r} \oplus L_{r}^{2}\left(\mathbb{R}^{N}\right)
$$

regarded as a real Banach space and $W_{r}$ given by (0.3) resp. (0.4). The energy and the charge are well-defined on $X$ and given by

$$
\begin{align*}
& E\left(\phi_{1}, \phi_{2}\right)=\frac{1}{2} \int_{\mathbb{R}^{n}}\left|\phi_{1}\right|^{2}+\left|\nabla \phi_{2}\right|^{2}-\frac{1}{2} \int\left|\phi_{2}\right|^{2}\left(-1+\log \left|\psi_{2}\right|^{2}\right)  \tag{3.7}\\
& Q\left(\phi_{1}, \phi_{2}\right)=\operatorname{Im} \int_{\mathbb{R}^{n}} \bar{\phi}_{2} \phi_{1} . \tag{3.8}
\end{align*}
$$

As scaling group we take $T_{\sigma} u(x)=u(\gamma / \sigma)$ where $u(x)=\left(\phi_{1}(x), \phi_{2}(x)^{t}\right.$ and define

$$
\begin{align*}
& \mathrm{K}_{\omega}\left(\phi_{1}, \phi_{2}\right)=\frac{1}{2} \int_{\mathbb{R}^{\mathrm{n}}}\left|\nabla_{2}\right|^{2}  \tag{3.9a}\\
& \mathrm{~V}_{\omega}\left(\phi_{1}, \phi_{2}\right)=\frac{1}{2} \int_{\mathbb{R}^{\mathrm{n}}}\left|\phi_{1}\right|^{2}-\left|\phi_{2}\right|^{2}\left(-1+\log \left|\phi_{2}\right|^{2}\right)+\omega \int_{\mathbb{R}^{\mathrm{n}}} \phi_{1} \bar{\phi}_{2} . \tag{3.9b}
\end{align*}
$$

Now for any fixed $\omega \in \mathbb{R}$ we have the inequality

$$
\begin{equation*}
\mathrm{V}_{\omega}\left(\phi_{1}, \phi_{2}\right) \leq \mathrm{V}_{\omega}\left(\mathrm{i} \omega \phi_{2}, \phi_{2}\right) \tag{3.10}
\end{equation*}
$$

with equality if and only if $\phi_{1}=i \omega \phi_{2}$ and therefore the minimization problem yields indeed a solitary wave solution. For the logarithmic Klein Gordon equation the ground state is given by

$$
\begin{equation*}
\phi_{\omega}(x)=\phi_{0}(x) \exp \left(-\omega^{2} / 2\right) \tag{3.11}
\end{equation*}
$$

Hence $d(\omega)=\frac{1}{N} \exp \left(-\omega^{2}\right) \int_{\mathbb{R}^{n}}\left|\nabla \phi_{o}\right|^{2}$ and we have the following result: If $\omega^{2} \leq 1 / 2$ then the ground state $\phi_{\omega}$ is unstable.
If $\omega^{2}>1 / 2$ it is stable.
3.4. We study the nonlinear wave equation

$$
\begin{equation*}
u_{t t}-\Delta u=g(u) \quad \text { on } \quad \mathbb{R}^{n} \quad, n \geq 3 \tag{3.12}
\end{equation*}
$$

where for real-valued functions $u$ on $\mathbb{R}^{n}$. Under suitable conditions on $g(u)$ (see e.g. [1,9]) (3.12) possesses a nontrivial ground state $U_{0}(x), U_{0}$ is always unstable by Theorem 4.

## 4. Extensions and Comments

4.1. First of all it is natural to ask about the relation between the model presented here and the one of Grillakis, Shatah and Strauss [7]:

We have the following:
If E" exists then $H_{\omega}=E^{\prime \prime}\left(\phi_{\omega}\right)-\omega Q^{\prime \prime}\left(\phi_{\omega}\right)$ where $\phi_{\omega}$ is a ground state satisfies the spectral assumptions of [7], i.e. $H_{\omega}$ has exactly one negative eigenvalue, a Kernel spanned by $U^{\prime}(0) \phi_{\omega}$ and the rest of the spectrum positive and bounded away from zero.

Indeed $H_{\omega}$ has at least some negative spectrum since $\left.\frac{d^{2}}{d \sigma^{2}}\right|_{\sigma=1}$, $I_{\omega}\left(T_{\sigma} \phi_{\omega}\right)<0$. On the other hand it has at most one negative eigenvalue since $\phi_{\omega}$ is a local minimum of $I_{\omega}$ on a $C^{1}$-hypersurface by (1.15). The assumption for solution of the minimization principle (1.10).
4.2. An extension to non-scale covariant models is also possible. Let $L_{\omega}$ satisfy the condition
(4.1)

$$
<L_{\omega}^{\prime}(\lambda u)-\lambda L_{\omega}^{\prime}(\cdot u),(\lambda-1) u><0 \quad \begin{aligned}
& u \neq 0 \\
& \lambda \neq 1
\end{aligned}
$$

and define

$$
\begin{equation*}
R_{\omega}(u)=\left\langle L_{\omega}^{\prime}(u), u\right\rangle \tag{4,2}
\end{equation*}
$$

We can define $d(w)$ by the minimization principle

$$
\begin{equation*}
d(\omega)=\inf \left\{I_{\omega}(u), u \in X, u \neq 0, R_{\omega}(u)=0\right\} \tag{4.3}
\end{equation*}
$$

Note that 0 is isolated in the set $\left\{u \in X, R_{\omega}(0)=0\right.$ in view of condition (4.1).

Under these assumptions one obtains the same criteria for stability/ instability in terms of $d(\omega)$ as for scale-covariant models.
4.3. The extension to more symmetries in the context of scale covariant functionals is very difficult. But there is one simple extension:

Let $K(u), V(u)$ be invariant under $N$ one parameter groups $\left(U_{j}(s)\right)_{i \leq j \leq N}$ which satisfy the 'commutation law'

$$
\begin{equation*}
U_{j}(s) U_{k}(t)=U_{k}(t) U_{j}(s) \quad \forall j_{k} \quad \forall s_{t} \in \mathbb{R} \tag{4.4}
\end{equation*}
$$

The associated conserved quantities $Q_{j}$ are required to satisfy the following scaling properties:
$Q_{1}$ satisfies (1.7) while $Q_{2}, \ldots, Q_{N}$ may have different scaling behaviour. We look for ground states

$$
\begin{equation*}
\mathrm{U}_{1}\left(\omega_{1} t\right) \mathrm{U}_{2}\left(\omega_{2} t\right) \ldots \mathrm{U}_{\mathrm{N}}\left(\omega_{\mathrm{N}} t\right) \quad \phi \vec{\omega} \tag{4.5}
\end{equation*}
$$

of (H) which satisfy $Q_{2}=\ldots=Q_{N}=0$.
In this case the stability of the orbit

$$
\left\{U_{1}\left(s_{1}\right) \ldots U_{N}\left(s_{N}\right) \phi_{\vec{\omega}} \quad, \quad s_{j} \in \mathbb{R}\right\}
$$

is determined by the behaviour of the function

$$
\begin{equation*}
d(\vec{u})=E\left(\phi_{\vec{w}}\right)-\omega_{1} Q\left(\phi_{\vec{w}}\right) \tag{4.6}
\end{equation*}
$$

in the variable $\omega_{1}$.
This assertion enables us to extend the stability results of section 3 to stability under perturbations which are not necessarily radial. Consider e.g. the logarithmic Schrödinger equation (3.1). Now we choose $X=W$ given by (0.3) instead of $W_{r}$. Then equation (3.1) is also invariant under translations. Gauge invariance and the of translation invariance satisfy the commutation law (4.4).

The momentum is given by

$$
\begin{equation*}
(u)=-\frac{1}{2} \int_{\mathbb{R}^{n}} \bar{u} \vec{\nabla} u \tag{4.7}
\end{equation*}
$$

which satisfies the scale relation

$$
\vec{P}\left(T_{\sigma} u\right)=\sigma^{n-1} \vec{P}(u) .
$$

Now $\phi_{\omega}$ is not only a ground state in $W_{r}$, it is also a ground state in $W$. Furthermore it has zero momentum.

Hence by the result above its stability under perturbations in $W$ is determined only by the frequency and there $\phi_{\omega}$ is also stable in $W$.

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A NOTE ON SOLUTIONS OF TWO-DIMENSIONAL SEMILINEAR ELLIPTIC VECTOR-FIELD EQUATIONS WITH STRONG NONLINEARITY

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Abstract: We prove the existence of solutions $u$ for some semilinear elliptic vector-field equations on $\operatorname{IR}^{2}$ with a nonlinearity which is allowed to grow at infinity 'nearly like a linear exponential' in lul. In some cases a growth like $\exp \left(a|u|^{\gamma}\right)$, $a>0,0<\gamma<2$, is allowed.
This is achieved by introducing an appropriate space $E$ of 'a priorisolutions' for which some important continuous imbeddings are proven replacing the well known Sobolev-imbeddings for the $d \geq 3$-dimensional case. Then the standard variational method is applied.

## I. Introduction

A first important step in the existence proof of a solution of a semilinear elliptic equation of the type

$$
\begin{align*}
& -\Delta u=g(u) \\
& u: \mathbb{R}^{d} \rightarrow \mathbb{R}^{n}, g: \mathbb{R}^{n} \rightarrow \mathbb{R}^{n}, n \geq 1 \tag{1.1}
\end{align*}
$$

is always to fix a space $E$ of à priori solutions. Motivated by the context where such equations occur one usually decides to look for solutions of finite kinetic energy

$$
\begin{equation*}
k(u)=\frac{1}{2}\|\nabla u\|_{2}^{2}=\left.\frac{1}{2} \int \sum_{j=1}^{n} \sum_{i=1}^{d} \frac{\partial u_{j_{j}}}{\left\lvert\, \frac{1}{\partial x_{i}}\right.}\right|^{2} d x . \tag{1.2}
\end{equation*}
$$

Another natural contraint is that the functions in $E$ should vanish at infinity in some sense. A way to express this is to define $E$ to be the completion $E^{1,2}\left(\mathbb{R}^{d} ; \mathbb{R}^{n}\right)$ of $C_{0}^{\infty}\left(\mathbb{R}^{d} ; \mathbb{R}^{n}\right)$ with respect to the norm $f \rightarrow\|\nabla f\|_{2}$. For $d \geq 3$ this definition has proved to be rather useful (see for instance [3-5] and references there) because then by Sobolev's inequality

$$
\begin{equation*}
\|f\|_{2 *} \leq s\|\nabla f\|_{2}, \quad 2^{*}=\frac{2 d}{d-2} \tag{1.3}
\end{equation*}
$$

this completion can be supposed to be realized as a subspace of $L^{2 *}\left(\mathbb{R}^{d} ; \mathbb{R}^{n}\right)$. However for $d=2$ this inequality breaks down and as a consequence this completion even is not a space of distributions as has been shown by J.L. Lions long ago.

So when treating the 'positive mass case' for scalar fields (i.e. $n=1)$ Berestycki et al. [2] decided to look for solutions in the subspace of radial functions of the Sobolev space $H^{1}\left(R^{2}\right)$. For the case of vector fields (i.e. $n>1$ ) and mass zero the problem is considerably more complicated as is well known. This case has been treated by Brezis and Lieb [3]. They admit arbitrary polynomial bounds for $g$ at infinity. The restriction on the behaviour of $g$ near zero is expressed by the condition that the potential $G$ of $g$ is negative near zero (i.e. $G(0)=0$ and $G(y)<0$ for $0<|y|<\varepsilon$ for some $\varepsilon>0)$. Their choice of the space of a priori solutions is well adapted to this situation but at the expense of not being a Banach space of functions on $\mathbb{R}^{2}$.

Furthermore by the context where such equations occur and also from a systematic point of view one should admit also a stronger than polynomial growth of $g$ at infinity as it was done in [2]. Here we want to show that this is indeed possible even if a more general kind of behaviour of $g$ near $y=0$ is allowed than that treated in [2].

Basically we will follow the general strategy of [3] respectively of [5] but with a different space $E$ of à priori solutions. Correspondingly the main step will be to introduce an appropriate space $E$ and to study its properties, in particular properties of continuous imbeddings and continuity properties of certain classes of Niemytski operators on it. This is done in section II and finally applied in section III to solve equation (1.1) under appropriate conditions admitting also a growth of $g$ at infinity nearly like a linear exponential in |u| and in the 'positive mass case' like an exponential in $\mid u l^{\gamma}$, $0<\gamma<2$.
II. The Sobolev-type spaces $E_{q}\left(\mathbb{R}^{d} ; \mathbb{R}^{n}\right), 2 \leq d \leq q<\infty, n \geq 1$. The spaces $E_{q}=E_{q}\left(\mathbb{R}^{d} ; \mathbb{R}^{n}\right)$ are defined as the completion of $C_{o}^{\infty}\left(\mathbb{R}^{d} ; \mathbb{R}^{n}\right)$ with respect to the norm

$$
\begin{equation*}
f \rightarrow\|f\|=\|f\|_{q}+\|\nabla f\|_{d} \tag{2.1}
\end{equation*}
$$

We consider this completion to be realized as a subspace of the Lebesgue space $L^{q}\left(\mathbb{R}^{d} ; \mathbb{R}^{n}\right)$ with norm $\|\cdot\|_{q}$.

If in section III for $d=2$ such a space $E_{q}$ is considered to be a space of à priori solutions for equation (1.1) the exponent $q$ will be choosen according to the behaviour of the nonlinearity $g$ near $y=0$. Note that for $q=d=2 E_{q}$ is just the Sobolev space $H^{1}\left(\mathbb{R}^{2} ; \mathbb{R}^{n}\right)$.
In this section we prove some properties of the spaces $E_{q}$ and some important inequalities for elements in $E_{q}$ which are fundamental for a variational approach to solutions of equation (1.1) (just as Sobolev's inequality is for the case $d \geq 3$ ).

## Lemma 2.1

For any $d, n, q, \quad 2 \leq d \leq q<\infty, n \geq 1$, the spaces $E_{q}\left(\mathbb{R}^{d} ; \mathbb{R}^{n}\right)$ are real separable reflexive Banach spaces.

Proof: With minor changes the proofs of the corresponding statements for Sobolev spaces apply [6,7].

The main information about the spaces $E_{q}$ which we will use is contained in the following imbedding theorem.

Theorem 2.2
Suppose $d \leq q<\infty$ and denote $\frac{1}{p}=1-\frac{1}{d}$. Then we have
a) $E_{q}\left(\mathbb{R}^{d}\right) \subset L^{r}\left(\mathbb{R}^{d}\right)$ for all $r \geq q$ with continuous injection, expressed by the inequality

$$
\begin{equation*}
\|u\|_{r}^{r} \leq c_{r}\|u\|_{q}^{q}\left(\frac{1}{d}\|\nabla u\|_{d}\right)^{r-q} \tag{2.2}
\end{equation*}
$$

for all $u \in E_{q}$.
b) In particular the following inequality holds for all $u \in E_{q}$, all $k=0,1,2, \ldots$,

$$
\begin{equation*}
\|u\|_{r(k)}^{r(k)} \leq(\alpha)_{k}^{p}\|u\|_{q}^{q}\left(\frac{1}{d}\|\nabla u\|_{d}\right)^{k \cdot p} \tag{2.3}
\end{equation*}
$$

where $r(k)=q+k \cdot p, \quad(\alpha)_{k}=\alpha(\alpha+1) \ldots(\alpha+k-1), \alpha=1+\frac{q}{p}$.

## Remark 2.1

Theorem 2.2 has some immediate consequences:
a) For $d=2 \leq q^{\prime} \leq q$ we have
$H^{1}\left(\mathbb{R}^{2} ; \mathbb{R}^{n}\right) \equiv E_{2} \subseteq E_{q^{\prime}} \subseteq E_{q}$
and by examples one sees that
$E_{q} \xlongequal{f} E_{q}$
whenever $q^{\prime}<q$.
b) For $q=d=2$ Theorem 2.2 provides a simple proof of the wellknown Sobolev imbeddings
$H^{1} \Leftrightarrow L^{r}\left(\operatorname{RR}^{2}\right) \quad$ for all $r \geq 2$
with explicitly known imbedding constants. At the end of this secion there is a comment on an impovement of these constants.

Proof: 1. Denote by $D^{1,1}$ the completion of $D=D\left(\mathbb{R}^{d}\right)=C_{o}^{\infty}\left(\mathbb{R}^{d}\right)$ with respect to the norm $\phi \rightarrow\|\nabla \phi\|_{1}$ on $D$. By continuous extension the Gagliardo-Nirenberg inequality ([1,6] and references)

$$
\begin{equation*}
\|V\|_{p} \leq \frac{1}{d}\|\nabla V\|_{1}, \quad \frac{1}{p}=1-\frac{1}{d} \tag{2.4}
\end{equation*}
$$

still holds on all of $D^{1,1}$. By induction on $k=0,1,2, \ldots$ we will show that for any fixed $u \in E_{q}$

$$
\begin{equation*}
v_{k}=|u|^{\alpha+k}, \quad \alpha=1+\frac{q}{p} \tag{2.5}
\end{equation*}
$$

belongs to $D^{1,1}$ so that the above inequality applies. Then the theorem easily follows.

In order to prepare the induction proof we recall some facts about smooth approximations of Lebesgue-integrable functions:

Choose some functions $x, n \in D$ such that

$$
0 \leq n \leq 1 \quad \text { supp } x, \eta \subseteq\{x:|x| \leq 1\}, \int x d x=1
$$

and $n(x)=1$ for all $|x| \leq \frac{1}{2}$.
Then define for $m=1,2, \ldots$

$$
x_{m}(x)=m^{-d} x(m x) \text { and } \eta_{m}(x)=\eta\left(\frac{1}{m} x\right)
$$

Clearly $x_{m}, \eta_{m} \in D, \int x_{m} d x=1$, and $\eta_{m}(x)=1$ on $|x| \leq \frac{1}{2} m$. Now for any $V \in L^{s}\left(\mathbb{R}^{d}\right)$ it is known $[6,7]$ :

$$
\begin{align*}
& V * X_{m} \in C^{\infty}\left(\mathbb{R}^{d}\right) \text { and } \quad V * X_{m} \rightarrow V \text { in } L^{s} \text { a) } \\
& \left.\Psi_{m}=n_{m}\left(V * \chi_{m}\right) \in D\left(I^{d}\right) \text { and } \quad \Psi_{m} \rightarrow V \quad \text { in } L^{s} \quad b\right) . \tag{2.6}
\end{align*}
$$

${ }^{2} \dot{\underline{q}}$ Now suppose $u \in E_{q}$ to be given. Clearly $V_{0}=|u|^{\alpha}$ belongs to $L^{\frac{q}{\alpha}}\left(\mathbb{R}^{d}\right)$ and according to the rules for weak derivatives we know

$$
\nabla v_{0}=\alpha|u|^{\alpha-1} \nabla|u|, \quad\|\nabla|u|\|_{d} \leq\|\nabla u\|_{d} .
$$

Hence Hölder's inequality implies by choice of $\alpha$

$$
\begin{equation*}
\left\|\nabla V_{0}\right\|_{1} \leq \alpha\left\||u|^{\alpha-1}\right\|_{p}\|\nabla|u|\|_{d} \leq \alpha\|u\|_{q}^{\alpha-1}\|\nabla u\|_{d} \tag{2.7}
\end{equation*}
$$

thus $\nabla V_{0} \in L^{1}$.
Next we prove that $V_{0}$ can be approximated in the $\|\nabla \cdot\|_{1}$-norm by the functions

$$
\begin{equation*}
\Psi_{m}^{0}=n_{m}\left(V_{0} * x_{m}\right) \tag{2.8}
\end{equation*}
$$

Let us write $\nabla V_{0}-\nabla \Psi_{m}^{0}=\left(1-\eta_{m}\right) \nabla V_{0}+\eta_{m} \cdot\left\{\nabla V_{0}+\nabla V_{0} * X_{m}\right\}-(\nabla n) \cdot\left(V_{0} * X_{m}\right)$. Then by Holder's inequality and the fact $\left\|\eta_{m}\right\|_{\infty} \leq 1$ the following estimate is available

$$
\left\|\nabla V_{0}-\nabla \Psi_{m}^{0}\right\|_{1}^{\prime} \leq\left\|\left(1-n_{m}\right) \nabla V_{0}\right\|_{1}+\left\|\nabla V_{0}-\nabla V_{0} * x_{m}\right\|_{1}+\left\|\nabla n_{m}\right\|_{\beta}\left\|V_{0}^{*} x_{m}\right\|_{\frac{q}{\alpha}}
$$

where $1=\frac{1}{\bar{\beta}}+\frac{\alpha}{q}$.
Since $\nabla V_{0} \in L^{1}$, the first term tends to zero for $m \rightarrow \infty$ by dominated convergence. The limit $m \rightarrow \infty$ of the second term also vanishes by statement 2.6.a).
Since $\left\|V_{0}{ }^{*} X_{m}\right\|_{\frac{q}{\alpha}}$ is bounded and since

$$
\left\|\nabla n_{m}^{\alpha}\right\|_{B}=m^{\frac{d}{B}-1}\|\nabla n\|_{B}, \quad \frac{d}{\beta}-1=-\frac{d}{q}
$$

the limit $m \rightarrow \infty$ of the third term also vanishes.

This proves $\nabla \Psi_{m}^{0} \rightarrow \nabla V_{0}$ in $L^{1}$ and thus $V_{0} \in D^{1,1}$.
Hence inequality (2.4) applies

$$
\left\|V_{0}\right\|_{p} \leq \frac{1}{d}\left\|F V_{o}\right\|_{1}
$$

that is by inequality (2.7)

$$
\|u\|_{\alpha \cdot p}^{\alpha} \leq \frac{\alpha}{d}\|u\|_{q}^{\frac{q}{p}}\|\nabla u\|_{d} .
$$

Now we have $\alpha p=q+p=r(1)$. Therefore inequality (2.3) holds for $k=1$.
3. Our induction hypothesis is that for some $k \geq 1$

$$
v_{j}=|u|^{\alpha+j} \in D^{1,1}
$$

and $\left.\quad\left\|v_{j}\right\|_{p}=\|u\|_{p(\alpha+j)}^{\alpha+j} \leq(\alpha)_{j+1}^{p}\|u\|_{q}^{\frac{q}{p}}\left(\frac{1}{d}\|\nabla u\|_{d}\right)^{j+1} \quad b\right)$
for $\mathrm{j}=0,1, \ldots \mathrm{k}-1$.
For $V_{k}$ it follows $V_{k} \in L^{q_{k}}\left(I^{d}\right)$ with

$$
q_{k}=\frac{p(\alpha+k-1)}{\alpha+k}>1
$$

and for $\nabla V_{k}$ we get

$$
\nabla V_{k}=(\alpha+k)|u|^{\alpha+k-1} \nabla|u|=(\alpha+k) V_{k-1} \nabla|u|
$$

hence by induction hypothesis $\nabla V_{k} \in L^{1}$ follows from Hölder's inequality.

Now we proceed as for $k=0$. This time the smooth approximation are

$$
\Psi_{m}^{k}=\eta_{m} \cdot\left(V_{k} * \chi_{m}\right), \quad m=1,2, \ldots
$$

and the relevant estimate is

$$
\left\|\nabla V_{k}-\nabla \Psi_{m}^{k}\right\|_{1} \leq\left\|\left(1-\eta_{m}\right) \nabla V_{k}\right\|_{1}+\left\|\nabla V_{k}-x_{m}^{*} \nabla V_{k}\right\|_{1}+\left\|\nabla \eta_{m}\right\| q_{k}^{\prime}\left\|V_{k} * x_{m}\right\| q_{q_{k}}
$$

and we conclude as above since

$$
\left\|\nabla \eta_{m}\right\|_{q_{k}^{\prime}}=m^{\frac{d}{q_{k}^{\prime}}-1}\|\nabla n\|_{q_{k}^{\prime}}
$$

with $\frac{d}{q_{k}^{1}}-1=d\left\{\frac{1}{q_{k}^{\prime}}-\frac{1}{d}\right\}=d \cdot\left\{\frac{1}{p}-\frac{1}{q_{k}}\right\}=-\frac{d}{r(k)}<0$.
This proves $V_{k} \in D^{1,1}$ so that by the Gagliardo-Nirenberg inequality

$$
\left\|v_{k}\right\|_{p} \leq \frac{1}{d}\left\|\nabla v_{k}\right\|_{1}=\frac{\alpha+k}{d}\left\|v_{k-1} \nabla|u|\right\|_{1} \leq \frac{\alpha+k}{d}\left\|v_{k-1}\right\|_{p}\|\nabla u\|_{d}
$$

follows. Thus inequality (2.9.b) holds also for $j=k$.
Hence this inequality and the statement $V_{k} \in D^{1,1}$ hold for all $k=0,1,2, \ldots$. Therefore the inequality in part b) of the theorem follows. The inequality in part a) follows from this using the interpolation inequality where the constant $C_{r}$ is explicitly known.
The following two corollaries provide some information about the definition and some continuity properties of certain Niemytski operators on $E_{q}$.

## Corollary 2.3

Suppose $F$ is a continuous function $I^{n} \rightarrow I R$ such that

$$
\begin{equation*}
|F(y)| \leq C|y|^{q} \sum_{k=0}^{\infty} a_{k}|y|^{\frac{p}{s} k} \quad \text { for } a \mid l \quad y \in I R^{n} \tag{2.10}
\end{equation*}
$$

with some exponents $d \leq q<\infty, 1 \leq s<\infty$, and some coefficients $a_{k}$ satisfying

$$
\begin{equation*}
\limsup _{k \rightarrow \infty}\left|a_{k}(\alpha)_{k}^{\frac{p}{s}}\right|^{1 / k}=\frac{1}{R}<\infty \tag{2.11}
\end{equation*}
$$

Then the Niemytski-operator $\hat{F}_{\underline{s}}$ associated with $F$ maps the subset $\left\{u \in E_{q}\left(\mathbb{R}^{d} ; \mathbb{R}^{n}\right) \left\lvert\, \frac{1}{d}\|\nabla u\|_{d}<R^{\frac{s}{p}}\right.\right\}=Z_{R}$ into $L^{1}\left(\mathbb{R}^{d}\right)$ and satisfies

$$
\begin{align*}
& \|\hat{F}(u)\|_{1} \leq \frac{c}{\frac{p}{p}}\|u\|_{q}^{q} \sum_{k=0}^{\infty} b_{k}\left[\frac{1}{d}\|\nabla u\|_{d}\right]^{\frac{p}{s} k}<\infty,  \tag{2.12}\\
& b_{k}=a_{k}(\alpha)_{k}^{s} . \tag{2.13}
\end{align*}
$$

Proof: For $u \in Z_{R}$ the following estimate is available by Theorem 2.2

$$
\begin{aligned}
\|\hat{F}(u)\|_{1} & =\int|F(u(x))| d x \leq C \int \sum_{k=0}^{\infty} a_{k}|u|^{q+\frac{p}{s} k} d x \leq \\
& \leq C \sum_{k=0}^{\infty} a_{k}\left\||u|^{\frac{q}{s^{+}}} \cdot|u|^{\frac{r(k)}{s}\left\|_{1} \leq c\right\| u \|_{q}^{\frac{q}{s}} \sum_{k=0}^{\infty} a_{k} \frac{r(k)}{s}} \begin{array}{l}
\|u\|_{r(k)} \\
\end{array} \quad \leq\right\| u \|_{q}^{q} C \sum_{k=0}^{\infty} b_{k}\left[\frac{1}{d}\|\nabla u\|_{d}\right]^{\frac{p}{s} k} .
\end{aligned}
$$

Since by assumption $\left[\frac{1}{d}\|\nabla u\|_{d}\right]^{\frac{p}{s}}<R$ this series converges, this poves the collary.

## Corollary 2.4

Under the assumptions of Corollary 2.4 with $R=+\infty$ the Niemytski operator $\hat{F}$ is sequential continuous as a map from $E_{q}$ equipped with weak topology into $L_{l o c}^{1}$, i.e. if $u_{i} \rightarrow u$ weakly in $E_{q}$ then $\hat{F}\left(u_{i}\right) \rightarrow \hat{F}(u)$ in $L_{l o c}^{1}\left(\mathbb{R}^{d}\right)$.

Proof: If a sequence $\left(u_{i}\right)_{i \in N}$ converges weakly in $E_{q}$ to some element $u$ then it is (strongly) bounded

$$
\sup _{i}\left\|u_{i}\right\|=c<\infty
$$

and we may assume also $u_{i} \rightarrow u$ almost everywhere on $\mathbb{R}^{d}$. By continuity of $F$ it follows $\hat{F}\left(u_{j}\right) \rightarrow \hat{F}(u)$ for $i \rightarrow \infty$ for almost every $x \in \mathbb{R}^{d}$. Thus the statement follows from Vitali's convergence theorem if we can show

$$
\sup _{i}\left\|\pi_{A} \hat{F}\left(u_{i}\right)\right\|_{1}=H(|A|) \rightarrow 0 \text { for }|A| \rightarrow 0
$$

for any measurable subset $A \subset \mathbb{R}^{d},|A|<\infty, \pi_{A}$ the operator of multiplication by $1_{A}$.
Similarly as in the proof of Corollary 2.3 we have for such a set $A$

$$
\left\|\pi_{A} \hat{F}\left(u_{i}\right)\right\|_{1} \leq\left\|\pi_{A} u_{i}\right\|_{q}^{\frac{q}{s}} c \sum_{k=0}^{\infty} a_{k}\left\|u_{i}\right\|_{r(k)}^{\frac{r(k)}{s}} .
$$

If we observe now that with $\frac{1}{\beta}=\frac{1}{q}-\frac{1}{q+p}>0$

$$
\left\|\pi_{A} u_{i}\right\|_{q} \leq|A|^{1 / \beta}\left\|u_{i}\right\|_{r(1)}
$$

we get

$$
\left\|\pi_{A} \hat{F}\left(u_{i}\right)\right\|_{1} \leq|A|^{\frac{q}{s} \beta}\left\|u_{i}\right\|_{\dot{r}(1)}^{\frac{q}{s y}}\left\|u_{i}\right\|_{c^{\prime}}^{\frac{q}{s}} \sum_{k=0}^{\infty} b_{k}\left(\frac{1}{d}\left\|u_{i}\right\|_{d}\right)^{\frac{p}{s} k}
$$

and thus by boundedness of $\left(u_{i}\right)_{i \in \mathbb{N}}$ in $E_{q}$ and by Theorem 2.2

$$
\sup _{i}\left\|\pi_{A} \hat{F}\left(u_{i}\right)\right\|_{1} \leq|A|^{\frac{q}{s^{\top} \beta}} c \text {, }
$$

since the series converges everywhere. Thus we conclude.
Some examples will illustrate the growth restrictions used in the above corollaries.

## Example 1

Suppose $F: \mathbb{R}^{n} \rightarrow \mathbb{R}$ to be continuous satisfying

$$
|F(y)| \leq\left(a|y|^{\gamma}\right)^{n_{\gamma}} \sum_{k=0}^{\infty} \frac{a^{k}|y|^{\gamma k}}{\left(n_{\gamma}+k\right)!}=e^{a|y|^{\gamma}}-\sum_{l=0}^{n_{\gamma}^{-1}} \frac{\left(a|y|^{\gamma}\right)^{1}}{1!}
$$

with $0<\gamma \leq 1, a>0$ and $q=\gamma n_{\gamma} \geq d$.

## Example 2

The continuous function $F: \mathbb{R}^{n} \rightarrow \mathbb{R}$ is supposed to satisfy for some constants $\mathrm{a}, \mathrm{b}>0$ and some exponents $0<\gamma \leq 1, \mathrm{~d} \leq \mathrm{q}<\infty$

$$
|F(y)| \leq b|y|^{q} c h\left(a|y|^{\gamma}\right) .
$$

In both cases we get for the radius $R_{\gamma}$ of convergence for the series defined by the coefficients $b_{k}$ according to eq.(2.13):

$$
R_{\gamma}=\left\{\begin{array}{llr}
\infty & \text { for } & 0<\gamma<1 \\
\text { finite } & \text { for } & \gamma=1 \\
0 & \text { for } & \gamma>1
\end{array}\right.
$$

Hence these corollaries are conviently applied for $0<\gamma<1$, i.e. if $F$ is bounded at infinite by some exponential of the form

$$
e^{a|y|^{\gamma}}
$$

For the case of a linear exponential (i.e. $\gamma=1$ ) some more complications arise. This case is not yet worked out in detail.

If we compare the estimates given in Theorem 2.2 with the corresponding estimates for the case $q=2$ obtained by other methods we see that at least for $q=2$ our estimates are not the best possible one's. We comment on this point.

For $q=2$ the space $E_{q}$ coincides with the usual Sobolev space $H^{1}=H^{1}\left(\mathbb{R}^{2} ; \mathbb{R}^{n}\right)$. It is well known that the Fouriertransform can be used to characterize conveniently elements of $L^{\mathrm{p}}$-spaces by integrability conditions on their Fouriertransforms only for $p=2$. Relying on Theorem 5.3 of reference [9] this fact can be used to give a simple proof of the following inequalities

$$
\begin{equation*}
\|u\|_{r} \leq \alpha(r)\|u\|_{1,2} \quad u \in H^{1}, \quad\|u\|_{1,2}^{2}=\|u\|_{2}^{2}+\|\nabla u\|_{2}^{2} \tag{2.14}
\end{equation*}
$$

for all $r \geq 2$ with

$$
\begin{equation*}
\alpha(r)=2^{-3\left(\frac{1}{2}-\frac{1}{r}\right)} \pi^{-\frac{1}{2}+\frac{3}{r}}(r-2)^{\left(\frac{1}{2}-\frac{1}{r}\right)}, \tag{2.15}
\end{equation*}
$$

hence $\alpha(r) \leq c r^{1 / 2}$ with some constant $c>0$ while inequality (2.3) says

$$
\begin{equation*}
\|u\|_{r} \leq \text { const } r\|u\|_{1,2} \tag{2.16}
\end{equation*}
$$

which is a weaker estimate.
Therefore one expects that the estimates of Theorem 2.2 can be improved also for the cases $q>2$. However for $q>2$ the convenient Fouriertransform method is not available. So a more general approach has to be used, may be along the lines of the method used for the Moser-Trudinger inequility [1,6,8].

We consider some examples. For $q=2$, $a>0$, and $0<\gamma \leq 2$ choose

$$
\begin{equation*}
F(y)=e^{a|y|^{\gamma}}-\sum_{k=0}^{n_{\gamma}^{-1}} \frac{\left(a|y|^{\gamma}\right)^{k}}{k!} \tag{2.17}
\end{equation*}
$$

with $\gamma \cdot n_{\gamma} \geq 2$. Relying on (2.14)-(2.15) we get for all $u \in H^{1}$

$$
\|\hat{F}(u)\|_{1} \leq \sum_{k=n_{\gamma}}^{\infty} \frac{a^{k}}{k!} \alpha(\gamma k)^{\gamma k}\|u\|_{1,2}^{\gamma k}
$$

and we see, using definition 2.11 , that the radius $R$ of convergence satisfies

$$
R=R_{\gamma}=\left\{\begin{array}{llr}
\infty & \text { for } & 0<\gamma<2  \tag{2.18}\\
\text { finite } & \text { for } & \gamma=2 \\
0 & \text { for } & \gamma>2
\end{array} .\right.
$$

We mention another example:

$$
\begin{equation*}
F(y)=|y|^{q} \text { ch } a|y|^{\gamma}, \quad a>0, \quad q \geq 2, \quad 0<\gamma \leq 2 . \tag{2.19}
\end{equation*}
$$

For $q=2$ we can proceed as above to obtain (2.18) also in this case. For $q>2$ however when we rely on Corollary 2.3 we get according to our example 2 from above for $\gamma=1$ and all $\beta>0$ with $A=\beta \frac{a}{2}\left(1+\frac{q}{2}\right)<1, u \in E_{q}, u \neq 0:$

while for $0<\gamma<1$ the integral is finite for all $\beta>0$ since then $R_{\gamma}=\infty$.

## III. Solution of some two dimensional vector field equations

As we will see the class of spaces $E_{q}$ introduced in the last section is well adapted to look for solution of equation (1.1) in them for $d=2$. According to the behaviour of the nonlinearity near $y=0$ the exponent $q \geq 2$ has to be fixed.

So first we list our hypotheses on the nonlinearity $g$ :
$\left(H_{0}\right) \quad g: \mathbb{R}^{n} \rightarrow \mathbb{R}^{n}$ is continuous, $g(0)=0$, such that $g(y)=\operatorname{grad} G(y)$ for $y \neq 0$ for some potential $G$.
$\left(H_{1}\right) \quad G: \mathbb{R}^{n} \rightarrow \mathbb{R}$ is continuous, of class $C^{1}$ in $\mathbb{R}^{n} \backslash\{0\}$, $G(0)=0, G(y)>0$ somewhere.
$\left(H_{2}\right) \quad G$ admits a decomposition $G=G_{+}-G_{-}$with continuous nonnegative functions $G_{ \pm}, G_{ \pm}(0)=0$, such that there exists an exponent $q \geq 2$ and $a$ constant $b>0$ with
(i) $b|y|^{q} \leq G_{-}(y)$ for all $y \in \mathbb{R}^{n}$
(ii) $G_{+}(y) \leq o\left(|y|^{q}\right)$ for $|y| \rightarrow 0$.
$\left(H_{3}\right) \quad y \quad|\rightarrow| y|l g(y)|$ satisfies the growth restrictions as expressed by statements (2.10) and (2.11) with $R=+\infty$.
$\left(\mathrm{H}_{3}\right) \quad \mathrm{G}(\cdot)$ satisfies the growth restrictions of assumption $\left(\mathrm{H}_{3}\right)$.

## Remarks

a) According to $\left(\mathrm{H}_{2}\right)$ the potential $G$ is negative near $y=0$, more precisely $G(y) \leq-c|y|^{q}$ for all sufficiently small $|y|$ with some constant $0<c<b$. Hence this assumption is a particular case of the corresponding assumption in [3].
b) By the examples mentioned in section II we know that assumption $\left(H_{3}^{\prime}\right)$ admits potentials $G$ which grow nearly like a linear exponential. More precisely ( $\mathrm{H}_{3}^{\prime}$ ) allows

$$
|G(y)| \leq c|y|^{q} e^{a|y|^{\gamma}} \text { for }|y| \rightarrow \infty
$$

with some constants $a, c>0$, and some exponent $0<\gamma<1$. For $q=2$ we can allow $0<\gamma<2$ according to our remarks in the previous section. Hence this assumption is considerably more general then the corresponding one of a polynomial bound.

According to these assumptions we decide to look for solutions in the Banach space $E_{q}=E_{q}\left(\mathbb{R}^{2} ; \mathbb{R}^{n}\right)$ where the exponent $q \geq 2$ is given by assumption ( $\mathrm{H}_{2}$ ).

When looking for weak solutions of equation (1.1) one clearly has to ensure that the 'potential' is differentiable in some sense. It turns out that essentially the weakest notion of differentiability which is natural in this context is sufficient for our purposes. So a first lemma states the existence of linear continuous Gâteaux-derivatives in all directions $v \in C_{0}^{\infty}\left(\mathbb{R}^{2} ; \mathbb{R}^{n}\right)$ for all points where it is defined.

## Lemma 3.1

Under the assumptions $\left(H_{0}\right),\left(H_{1}\right)$, and $\left(H_{3}\right)$ the functional

$$
\begin{align*}
& V(u):=\int G(u(x)) d x, \quad u \in D(V) \\
& D(V)=\left\{v \in E_{q} \mid G(u(\cdot)) \in L^{1}\left(\mathbb{R}^{2}\right)\right\} \tag{3.1}
\end{align*}
$$

has linear continuous Gâteaux-derivatives $V^{\prime}(u ; v)$ at every $u \in D(V)$ in all directions $v \in C_{o}^{\infty}\left(\mathbb{R}^{2} ; \mathbb{R}^{n}\right)$ given by

$$
\begin{equation*}
v^{\prime}(u ; v)=\int g(u) \cdot v d x . \tag{3.2}
\end{equation*}
$$

Proof: The standard versions of such a differentiability result assume some polynomial bound for $g$ and $G$ (see $[4,5]$ and references there). So they don't apply directly. However if we take into account the basic inequalities from Theorem 2.2 it is not hard to show that the proof of Proposition 2.5 in [4] can be extended to the present case. The estimates for the term

$$
\|\hat{g}(u) \cdot v[\lambda \leq|u|]\|_{1}
$$

([ $\lambda \leq|u|]=$ characteristic function of the set $\left\{x \in \mathbb{R}^{2}|\lambda \leq|u(x)|\}\right)$ used there can be done in analogy with the estimate for

$$
\left\|\hat{G}_{+}\left(f_{j}\right)\left[\lambda<\| f_{j} \mid\right]\right\|_{1}
$$

for $\lambda \rightarrow \infty$ which will be given explicitly later in the proof of Theorem 3.3.

Lemma 3.2
Suppose $g$ satisfies $\left(H_{0}\right),\left(H_{1}\right),\left(H_{3}\right)$, and

$$
\begin{equation*}
g^{-1}(0) \subseteq\{0\} u\left\{y \in \operatorname{IR}^{n}|\delta \leq|y|\}\right. \tag{3.3}
\end{equation*}
$$

for some $\delta>0$. If for $u \in E_{q}$

$$
\hat{g}(u)=0
$$

holds in the sense of distributions then the function $u$ vanishes : $u=0$.

Proof: This results follows from the observation that weakly differentiable functions have no finite jumps [3]. Relying on [3'] a complete proof is given in [5]. By our assumptions on $g$ we know $\hat{g}(u) \in L_{\text {loc }}^{1}$ for every $u \in E_{q}$, hence $\hat{g}(u)=0$ almost everywhere on $I^{2}$. By the assumption about the zeros of $g$ this equation can hold only if lu| had a jump of height at least $\delta$. Thus $u=0$ follows.

Remark 3.1
a) Assumptions $\left(H_{0}\right),\left(H_{1}\right)$, and $\left(H_{3}\right)$ imply assumption ( $H_{3}^{\prime}$ ).
b) Hypotheses $\left(H_{0}\right)-\left(H_{2}\right)$ imply assumption (3.3).

## Theorem 3.3

If the assumption $\left(H_{0}\right)-\left(H_{2}\right)$ and $\left(H_{3}\right)$ are satisfied then the equation (1.1) has a nontrivial solution in $E_{q}$.

Proof: Step 1: Construction of an appropriate minimizing sequence: By our assumptions and definition (3.1) the following minimization problem is well defined:

$$
I=\inf \left\{\left.K(v)=\frac{1}{2}\|\nabla v\|_{2}^{2} \right\rvert\, v \in D(v), v \neq 0, V(v) \geq 0\right\}
$$

and there is a minimizing sequence

$$
\begin{equation*}
f_{j} \in D(V), f_{j} \neq 0, V\left(f_{j}\right) \geq 0, I=\lim _{j \rightarrow \infty} K\left(f_{j}\right) \tag{3.5}
\end{equation*}
$$

Because of the covariance properties

$$
\begin{equation*}
K\left(v_{\sigma}\right)=K(v), V\left(v_{\sigma}\right)=\sigma^{-2} V(v), v_{\sigma}(\cdot)=v(\sigma \cdot) \tag{3.6}
\end{equation*}
$$

we may assume in addition for all $j \in \mathbb{N}$

$$
\begin{equation*}
\left\|f_{j}\right\|_{q}=1 \tag{3.7}
\end{equation*}
$$

Hence we get the following chain of inequalities using first part i) and then part ii) of assumption $\left(\mathrm{H}_{2}\right)$ :

$$
\begin{align*}
b & \leq \int \hat{G}_{-}\left(f_{j}\right) d x \leq \int \hat{G}_{+}\left(f_{j}\right) d x \leq \\
& \left.\leq \int\left[\left|f_{j}\right|<\varepsilon\right] \hat{G}_{+}\left(f_{j}\right) d x+\int\left[\varepsilon \leq\left|f_{j}\right| \leq \lambda\right] \hat{G}_{+}\left(f_{j}\right) d x+\int\left[\lambda<\mid f_{j}\right]\right] \hat{G}_{+}\left(f_{j}\right) d x \\
& \leq a(\varepsilon)+C_{\varepsilon, \lambda}\left|\left[\varepsilon \leq\left|f_{j}\right| \leq \lambda\right]\right|+J_{\lambda}\left(f_{j}\right) \tag{3.8}
\end{align*}
$$

where by part (ii)

$$
G_{+}(y) \leq a(\varepsilon)|y|^{q} \quad \text { for } \quad|y|<\varepsilon
$$

with $a(\varepsilon) \rightarrow 0 \quad$ for $\varepsilon \rightarrow 0 \quad$ and where

$$
C_{\varepsilon, \lambda}=\sup \left\{G_{+}(y)|\cdot \varepsilon \leq|y| \leq \lambda\}<\infty .\right.
$$

The Lebesgue measure of a measurable set $M$ is denoted by $|M|$. The third term in (3.8) is controlled by assumption $\left(\mathrm{H}_{3}\right)$ :

$$
\begin{aligned}
& \int\left[\lambda<\mid f_{j} \|\right] \hat{G}_{+}\left(f_{j}\right) d x \leq \int\left(\left[\lambda<\left|f_{j}\right|\right] \sum_{k=0}^{\infty} a_{k}\left(f_{j}\right)^{q+\frac{2}{s} k}\right) d x \\
& \leq \sum_{k=0}^{\infty} a_{k}\left\|\left[\lambda<\left|f_{j}\right|\right]\left|f_{j}\right|^{\frac{q}{s^{+}}}\left|f_{j}\right|^{\frac{r(k)}{s}}\right\|_{1} \leq \\
& \leq\left\|\left[\lambda<\mid f_{j} \|\right] f_{j}\right\|\left\|_{k=0}^{\frac{q}{s^{+}}} \cdot \sum_{k}^{\infty} a_{k}\right\| f_{j} \|_{r(k)}^{\frac{r(k)}{s}} .
\end{aligned}
$$

By Theorem 2.2 we have

$$
\begin{equation*}
\left\|f_{j}\right\| \underset{r(k)}{r(k)} \leq(\alpha)_{k}^{2}\left(\frac{1}{2} k\left(f_{j}\right)\right)^{k} \tag{3.9}
\end{equation*}
$$

where $\alpha=1+q / 2$, hence

$$
\begin{equation*}
J_{\lambda}\left(f_{j}\right) \leq\left\|\left[\lambda<\left|f_{j}\right|\right] f_{j}\right\|_{q}^{\frac{q}{s^{+}}} \sum_{k=0}^{\infty} b_{k}\left(\frac{1}{2} k\left(f_{j}\right)\right)^{\frac{2}{k}} \tag{3.10}
\end{equation*}
$$

with $b_{k}$ according to definition (2.13).
Since $\left(K\left(f_{j}\right)\right)_{j \in \mathbb{N}}$ is bounded Theorem 2.2 implies that $\left(\left\|f f_{j}\right\|_{j \in \mathbb{N}}\right.$ is bounded for every $r \geq q$.
Therefore for fixed $r>q, \frac{1}{q}=\frac{1}{r}+\frac{1}{\tau}$, we know

$$
\left\|\left[\lambda<\left|f_{j}\right|\right] f_{j}\right\|_{q} \leq\left|\left[\lambda<\left|f_{j}\right|\right]\right|^{1 / \tau}\left\|f_{j}\right\|_{r} \leq C\left|\left[\lambda<\left|f_{j}\right|\right]\right|^{1 / \tau} \leq C^{\prime} \lambda^{-\frac{q}{\tau}}
$$

since $1=\left\|f_{j}\right\|_{q}^{q} \geq\left\|\left[\lambda<\left|f_{j}\right|\right]\left|f_{j}\right|\right\|_{1} \|_{1} \geq \lambda^{q}\left|\left[\lambda<\left|f_{j}\right|\right]\right|$.
By assumption ( $\mathrm{H}_{3}$ ) the series in inequality (3.10) converges everywhere on $\mathbb{C}$. Thus we get

$$
\begin{equation*}
\sup _{j} J_{\lambda}\left(f_{j}\right) \rightarrow 0 \quad \text { for } \quad \lambda \rightarrow \infty \tag{3.11}
\end{equation*}
$$

Hence there are $0<\varepsilon<\lambda<\infty$ such that by inequality (3.8)

$$
\frac{b}{2} \leq C_{\varepsilon, \lambda}\left|\left[\varepsilon \leq\left|f_{j}\right| \leq \lambda\right]\right|
$$

holds for every $j \in \mathbb{N}$. Now we can use the lemma of concentration by translation [3,5] to conclude as in [3] that there is a minimizing sequence $\left(u_{j}\right)_{j} \in \mathbb{N}$ and an element $u \in E_{q}$ such that

$$
u_{j} \rightarrow u \quad \text { for } \quad j \rightarrow \infty
$$

a) weakly in $E_{q}$
b) almost everywhere on $\operatorname{IR}^{2}$
c) $u \neq 0$.

Step 2: We are going to show that the limit function $u$ according to statement (3.12) actually is a solution of the differential equation (1.1).

Inequality (3.8) implies in particular

$$
\sup _{i \in N} \int \hat{G}_{+}\left(u_{i}\right) d x<\infty .
$$

Thus by (3.12.b) and Fatou's lemma we get $\hat{G}_{+}(u) \in L^{1}\left(\mathbb{R}^{2}\right)$. By $V\left(u_{j}\right)=V_{+}\left(u_{j}\right)-V_{-}\left(u_{j}\right) \geq 0$ we get in the same way $\hat{G}_{-}(u) \in L^{1}\left(R^{2}\right)$. This proves $\hat{G}(u) \in L^{1}$, i.e. $u \in D(V)$. If also $V(u)=\int \hat{G}(u) d x \geq 0$ were known the limit function $u$ would be a non trivial minimizer. In the case of scalar fields ( $n=1$ ) it is easy to prove $V(u) \geq 0$ (see later remark). Here instead it is shown directly that a suitably scaled version of $u$ is a weak solution of our equation (without showing first $V(u) \geq 0$ ).

Take a fixed $v \in C_{o}^{\infty}\left(\mathbb{R}^{2} ; \mathbb{R}^{n}\right)$ with compact support $K$ and apply Corollary 2.4 to the sequences $\left(\pi_{K} u_{i}\right)_{i \in \mathbb{N}}$ and $\left(\pi_{K} u_{i}+v\right)_{i \in \mathbb{N}}$.

Since $\pi_{K} \hat{G}\left(u_{i}\right)=\hat{G}\left(\pi_{K} u_{i}\right)$ this corollary implies
that $\hat{G}\left(\pi_{K} u_{i}\right) \rightarrow \hat{G}\left(\pi_{K} u\right)$ in $L^{1}\left(R^{2}\right)$ and similarly

$$
\hat{G}\left(\pi_{K} u_{i}+v\right) \rightarrow \hat{G}\left(\pi_{K} u+v\right), \quad \text { for } \quad i \rightarrow \infty ;
$$

hence

$$
\begin{equation*}
\hat{G}\left(\pi_{K} u_{i}+v\right)-\hat{G}\left(\pi_{K} u_{i}\right) \underset{i \rightarrow \infty}{ } \hat{G}\left(\pi_{K} u+v\right)-\hat{G}\left(\pi_{K} u\right) \quad \text { in } L^{1} \tag{3.13}
\end{equation*}
$$

but $\hat{G}\left(\pi_{K} u+v\right)-\hat{G}\left(\pi_{K} u\right)=\hat{G}(u+v)-\hat{G}(v)$,
and therefore (3.13) proves

$$
\begin{equation*}
V\left(u_{i}+v\right)-V\left(u_{i}\right) \underset{i \rightarrow \infty}{ } V(u+v)-V(u) . \tag{3.14}
\end{equation*}
$$

Now using the differentiability result of Lemma 3.1 one can show just as in [3] that there is some $\lambda \geq 0$ such that

$$
\begin{equation*}
V^{\prime}(u ; v)=\lambda K^{\prime}(u ; v) \text { for for all } v \in C_{o}^{\infty}\left(\mathbb{R}^{2} ; \mathbb{R}^{n}\right) \tag{3.15}
\end{equation*}
$$

If $\lambda=0$ then Remark 3.1 b and Lemma 3.2 imply $u=0$.
This contradiction proves $\lambda>0$ and we may rescale $u$ according to

$$
\bar{u}=u_{\sigma}, \quad \sigma=\sqrt{\lambda}
$$

and obtain a nontrivial weak solution $\bar{u}$ :

$$
\begin{equation*}
V^{\prime}(\bar{u} ; v)=K^{\prime}(\bar{u} ; v) \quad \text { for all } \quad v \in C_{0}^{\infty} . \tag{3.16}
\end{equation*}
$$

## Remark 3.2

a) In the case of scalar fields $(n=1)$ and an even potential $G$ spherically symmetric rearrangement of functions can be used. Then it suffices to restrict to the subspace of spherically symmetric (nonincreasing) functions in $E_{q}$.
In this case for $q=2$ (the 'positive mass' case) the existence of infinitely many solutions has been stated in [2] for a class of potentials which are bounded at infinity by

$$
\begin{equation*}
e^{a|y|^{\gamma}}, \quad a>0, \quad \gamma=2 \tag{3.17}
\end{equation*}
$$

Taking into account the comments on the case $q=2$ at the end of tion 2 it is not hard to see by (2.14) - (2.15) that Theorem 3.3 can easily be extended in the vector field case for $q=2$ to cover a nonlinearity satisfying (3.17) with $0<\gamma<2$. For $\gamma=2$ the radius of convergence of the relevant power series is finite according to (2.18); this allows only to treat those cases for which it is possible to find a minimizing sequence of elements with sufficiently small $H^{1}$-norms. This point is stillunder investigation.
b) Recall from [6], Theorem 8.8 that the basic step of the elliptic regularity theory applies if $\hat{g}(u) \in L_{\text {loc }}^{2}$ is known. The following Lemma provides this information.

## Lemma

If the continuous function $g$ satisfies hypothesis $\left(\mathrm{H}_{3}\right)$ then for any $u \in E_{q}\left(I R^{2}\right): \hat{g}(u) \in L^{2}\left(I R^{2}\right)$.

Proof: By $\left(\mathrm{H}_{3}\right)$ one has

If $q \geq 2$ then $q^{\prime}=2(q-1) \geq q$ so that we can apply the estimates of Theorem 2.2 for $r_{k}^{\prime}=q^{\prime}+2 k$, i.e. $r_{2 k}^{\prime}=2\left(r_{k}-1\right)$, where $r_{k}=r(k)=q+2 k$. Hence with $\alpha^{\prime}=1+\frac{q^{\prime}}{2}=q$ we have

$$
\|u\|_{r}^{r_{2 k}^{\prime}}{ }_{2 k}^{\prime} \leq\left(\alpha^{\prime}\right)_{2 k}^{2}\|u\|_{q}^{q^{\prime}}\left(\frac{1}{2}\|\nabla u\|_{2}\right)^{2(2 k)}
$$

and thus by Hölder's inequality and Theorem 2.2 (observe $E_{q} \subseteq E_{q}$ for $q \leq q^{\prime}$ )

$$
\begin{aligned}
&\left\||u|^{2(q-1)}+\frac{2 k}{s}+\frac{21}{s}\right\|_{1} \leq\left\|\left|u\left\|^{q-1+\frac{2}{s} k}\right\|_{2}\left\||u|^{q-1+\frac{21}{s}}\right\|_{2}\right.\right. \\
& \leq\|u\|_{q^{\prime}}\|u\|_{r_{2 k}^{\prime}}^{r_{2 k}^{\prime} / s}\|u\|_{r_{2}^{\prime} 1}^{r_{21}^{\prime} / s} \leq\|u\|_{q}^{q^{\prime}}\left(\alpha^{\prime}\right)_{2 k}^{1 / s}\left(\alpha^{\prime}\right)_{21}^{1 / s} A^{k+1}
\end{aligned}
$$

where $A=\left(\frac{1}{2}\|\nabla u\|_{2}\right)^{2 / s}$.

This proves finally

$$
\|\hat{g}(u)\|_{2} \leq\|u\|_{2(q-1)}^{q-1} \cdot \sum_{k=0}^{\infty} a_{k}(q)_{2 k}^{1 / s} A^{k}<\infty
$$

since
implies

$\lim _{k \rightarrow \infty} \sup _{k \rightarrow}\left|a_{k}(q)_{2 k}^{1 / s}\right|^{1 / k}=0$
so that the above series converges for every $A \in \mathbb{C}$.

Therefore by Theorem 8.8 of [6] it follows that a weak solution $u$ of equation (1.1) in $E_{q}$ actually belongs to $W_{l o c}^{2,2} \cap E_{q}$.

Hence such a solution solves this equation in the sense of equality almost everywhere on $\mathrm{IR}^{2}$.
c) In analogy with method explained in [4] for the $d \geq 3$-dimensional case equations of the form

$$
\begin{aligned}
& -\Delta u(x)=g(x, u(x)) \\
& u: \mathbb{I}^{2} \rightarrow \mathbb{R}^{n}, g: \mathbb{R}^{2} x \mathbb{R}^{n} \rightarrow \mathbb{R}^{n}, n \geq 1
\end{aligned}
$$

can also be treated in the spaces $E_{q}\left(\mathbb{R}^{2} ; \mathbb{R}^{n}\right)$ and for some classes of examples the existence of a (weak) solution can be proven.

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# SOME REMARKS ON THE NONLINEAR SCHRÖDINGER 

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## 1 INTRODUCTION

We consider the Cauchy problem (initial value problem) for nonlinear Schrödinger equations in $R^{n}$, of the form

$$
\begin{equation*}
\mathrm{iu}_{\mathrm{t}}+\Delta \mathrm{u}=\mathrm{g}(\mathrm{u}) \quad, \quad \mathrm{u}(0, \cdot)=\varphi(\cdot) . \tag{NLS}
\end{equation*}
$$

Here $u$ is a complex-valued function defined on $[0, \mathbf{T}) \times \mathbf{R}^{\mathrm{n}}$ for some $\mathrm{T}>0, \varphi$ is some initial condition defined on $\mathbf{R}^{\mathrm{n}}$ and g is some nonlinear (local or non-local) mapping. In most of the examples that have been considered, g has some symmetry properties and is also the gradient of some functional G . Thus, at least formally, we have both conservation of charge and conservation of energy, that is

$$
\begin{aligned}
& \int_{\mathbf{R}^{\mathrm{n}}}|u(\mathrm{t}, \mathrm{x})|^{2} \mathrm{dx}=\int_{\mathbf{R}^{\mathrm{n}}}|\varphi(\mathrm{x})|^{2} \mathrm{dx}, \\
& \frac{1}{2} \int_{\mathbf{R}^{\mathrm{n}}}|\nabla \mathrm{u}(\mathrm{t}, \mathrm{x})|^{2} \mathrm{dx}+\mathbf{G}(\mathrm{u}(\mathrm{t} \cdot \cdot))=\frac{1}{2} \int_{\mathbf{R}^{\mathrm{n}}}|\nabla \varphi(\mathrm{x})|^{2} \mathrm{dx}+\mathrm{G}(\varphi(\cdot)) .
\end{aligned}
$$

Clearly, the charge and energy involve the $\mathrm{H}^{1}$-norm of the solution and therefore it is important to be able to solve the local Cauchy problem in the space $\mathrm{H}^{1}\left(\mathbf{R}^{\mathrm{n}}\right)$. Indeed, when this is possible, then global existence results follow easily from the above conservation laws and some conditions on $G$ (for example $\mathrm{G} \geq 0$ ). Obviously, in order to be able to do so, there are some necessary requirements on g ; for example, g and G need to be well defined on $\mathrm{H}^{1}$. In the applications, this will impose some "growth" conditions on g.

The initial value problem for the nonlinear Schrödinger equation in $\mathrm{H}^{1}$ has been studied in the past few years, essentially by J Ginibre and G Velo [5,6,7,8] and by T Kato [9]. In the model case where $g(u)=|u|^{-1} u$, the Cauchy problem is well posed in $H^{1}\left(R^{n}\right)$ for $1 \leq p<(n+2) /(n-2)$. The methods are of a perturbative nature and rely basically on sharp dispersive properties of the linear equation. All the previous proofs require at some stage (for obtaining local estimates of the solution in $\mathrm{H}^{1}\left(\mathbf{R}^{\mathrm{n}}\right)$ ) differentiation of the equation with respect to x , and so they don't apply to nonlinearities for which the x -dependence is not smooth enough. For example, the results mentioned above do not cover the case where $g(u)=|u|^{-1} u+V u, V$ being a non-smooth potential.

We present here a result that covers most of the previously known cases and that holds without any smoothness assumption on $g(u)$ with respect to $x$. The proof proceeds by an approximation argument followed by a passage to the limit. Uniform estimates on the approximating solutions are obtained from the conservation of the energy, and the passage to the limit (as well as uniqueness) relies on the dispersive properties of the linear Schrödinger equation. Let us remark that we do not need the conservation of charge ( $g$ does not have to satisfy the corresponding symmetry properties) while we definitely need the conservation of the energy (g must be the gradient of some potential G). This is in contrast with the result of [9], which applies to local nonlinearities for which there is possibly no energy (but that are sufficiently smooth with respect to x ).

In section 2 we state the main result and we give some examples of applications, and in section 3 we give a sketch of the proof. The reader is referred to [3] for the complete proof and to [2,4] for some related results in the critical case where the present method just fails.

## 2 THE MAIN RESULT

We begin by introducing some notation. We denote by $H^{k}$ the Sobolev space $H^{k}\left(\mathbf{R}^{n}, C\right)$ for any integer $k$, equipped with its usual norm and scalar product (always considered as a real Hilbert or Banach space) and by $L^{p}$ the space $L^{p}\left(\mathbf{R}^{n}, C\right)$ for any $p \in[1, \infty]$, also equipped with its usual norm. We denote by $\left\|\|_{H^{k}}\right.$ (respectively $\| \|_{L^{p}}$ ) the norm in $H^{k}$ (respectively $L^{p}$ ), and by $<,>$ the duality pairing between $\mathrm{H}^{-1}$ and $\mathrm{H}^{1}$. $\mathrm{p}^{\prime}$ is the conjugate exponent of $p$, given by $1 / \mathrm{p}+1 / \mathrm{p}=1$. For a given $\varphi \in H^{1}$, we are interested in the initial value problem (NLS).

We now state the assumptions on the nonlinear interaction g . We assume that g is of the form $g=\sum_{k=1}^{N} g_{k}$, where $g_{k} \in C\left(H^{1}, H^{-1}\right)$. For each of the $g_{k^{\prime}}$, we assume the following. There exists a function $\mathrm{C}_{\mathrm{k}} \in \mathrm{C}\left(\mathbf{R}_{+}, \mathbf{R}_{+}\right)$, two numbers $\mathrm{r}_{\mathrm{k}}, \rho_{\mathrm{K}} \in\left[2,2 \mathrm{n} /(\mathrm{n}-2)\right.$ ) ( $\mathrm{r}_{\mathrm{k}}, \rho_{\mathrm{K}} \in[2, \infty)$ if $\left.\mathrm{n}=1,2\right)$, and a sequence $\mathrm{g}_{\mathrm{k}, \mathrm{m}} \in \mathrm{C}\left(\mathrm{L}^{2}, \mathrm{~L}^{2}\right)$ such that
$g_{k, m}(0)=0$ and $g_{k, m}$ is Lipschitz continuous from bounded sets of $L^{2}$ to $L^{2}$.
$\mathrm{g}_{\mathrm{x}, \mathrm{m}} \rightarrow \mathrm{g}_{\mathrm{k}}$ in $\mathrm{L}^{(\rho \mathbf{k})^{\prime}}$ as $\mathrm{m} \rightarrow \infty$, uniformly on bounded sets of $\mathrm{H}^{1}$.

There exists $\mathrm{G}_{\mathrm{k}, \mathrm{m}} \in \mathrm{C}^{1}\left(\mathrm{~L}^{2}, \mathbf{R}\right)$ such that $\mathrm{G}_{\mathrm{k}, \mathrm{m}}(0)=0$ and $\mathrm{g}_{\mathrm{k}, \mathrm{m}}=\left(\mathrm{G}_{\mathrm{k}, \mathrm{m}}\right)^{\prime}$.

$$
\begin{equation*}
\left\|\mathrm{g}_{\mathrm{k}, \mathrm{~m}}(\mathrm{v})-\mathrm{g}_{\mathrm{k}, \mathrm{~m}}(\mathrm{u})\right\|_{\mathrm{L}}(\mathrm{pk})^{\prime} \leq \mathrm{C}_{\mathbf{k}}(\mathrm{M})\|\mathrm{v}-\mathrm{u}\|_{\mathrm{L}^{\mathrm{rk}}} \text { for } u, \mathrm{v} \in \mathrm{H}^{1} \text {, with }\|\mathrm{u}\|_{\mathbf{H}^{1}} \leq \mathrm{M} \text { and }\|\mathrm{v}\|_{\mathbf{H}^{1}} \leq \mathrm{M} . \tag{4}
\end{equation*}
$$

From (3) we have $G_{k, m}(u)=\int_{0}^{1}\left\langle g_{k, m}(s u), u\right\rangle d s$, for every $u \in H^{1}$. Thus, if we set

$$
\begin{equation*}
\mathrm{G}_{\mathrm{k}}(\mathrm{u})=\int_{0}^{1}\left\langle\mathrm{~g}_{\mathrm{k}}(\mathrm{su}), \mathrm{u}\right\rangle \mathrm{ds}, \tag{5}
\end{equation*}
$$

then from (2), (5) and the embeddings $\mathrm{H}^{1} \subset \mathrm{~L}^{\mathrm{rk}}, \mathrm{L}^{(\mathrm{pk})^{\prime}} \subset \mathrm{H}^{-1}$, we get

$$
\begin{equation*}
\mathrm{G}_{\mathrm{k}, \mathrm{~m}} \rightarrow \mathrm{G}_{\mathrm{k}} \text { as } \mathrm{m} \rightarrow \infty \text {, uniformly on bounded sets of } \mathrm{H}^{1} . \tag{6}
\end{equation*}
$$

Finally, let us define the functionals $G, G^{m}, E$ and $E^{m}$ by

$$
\begin{align*}
& G^{m}(u)=\sum_{k=1}^{N} G_{k, m^{\prime}}(u) \text { for } u \in H^{1},  \tag{7}\\
& G(u)=\sum_{k=1}^{N} G_{k}(u) \text { for } u \in H^{1},  \tag{8}\\
& E^{m}(u)=\frac{1}{2} \int_{\mathbf{R}^{n}}|\nabla u(x)|^{2} d x+G^{m}(u) \text { for } u \in H^{1},  \tag{9}\\
& E(u)=\frac{1}{2} \int_{\mathbf{R}^{n}}|\nabla u(x)|^{2} d x+G(u) \text { for } u \in H^{1} . \tag{10}
\end{align*}
$$

We can now state our main result.

THEOREM 1. Assume that $g$ satisfies the above hypotheses. Then for any $\varphi \in \mathrm{H}^{1}$, there exists $\mathrm{T}^{*}>0$ and a solution $u \in \mathrm{C}\left(\left[0, \mathrm{~T}^{*}\right), \mathrm{H}^{1}\right) \cap \mathrm{C}^{1}\left(\left[0, \mathrm{~T}^{*}\right), \mathrm{H}^{-1}\right)$ of (NLS). In addition, we have the following properties.
(i) $u$ is unique in $C\left(\left[0, \mathrm{~T}^{\prime}\right), \mathrm{H}^{1}\right)$ for any $\mathrm{T}^{\prime}>0$,
(ii) either $\mathrm{T}^{*}=\infty$ or else $\mathrm{T}^{*}<\infty$ and $\|u(t)\|_{H^{1}} \rightarrow \infty$ as $\mathrm{t} \uparrow \mathrm{T}^{*}$,
(iii) $E(u(t))=E(\varphi)$, for every $t \in\left[0, T^{*}\right)$.

Several remarks and comments are in order, concerning both the hypotheses on g and the statement of Theorem 1.

REMARK 1. It follows from conditions (1) and (2) that $g(0)=0$. We assume this only for the sake of simplicity.Allowing $\mathrm{g}(0) \neq 0$ would result in adding a constant term $\varphi \in \mathrm{H}^{-1}$ to the right hand side of (NLS), which would not be too difficult to handle.

REMARK 2. We assume that g is split into N terms $\mathrm{g}_{\mathrm{k}}$ satisfying different conditions. This is rather natural since in the applications (see below) the nonlinearity can be the sum of several terms having properties that are actually different.

REMARK 3. When applying Theorem 1 to some particular example, what is given in general is $g$ as the sum of several terms $\mathrm{g}_{\mathrm{k}}$. One would expect to need only assumptions on the $\mathrm{g}_{\mathrm{k}}$ 's. However, our assumptions ((1) to (4)) are on some approximating sequence $g_{k, m}$, which seems somewhat unnatural.

This phenomenon comes from the following technical difficulty. In the proof, we need to approximate $g_{\mathrm{k}}$ by some sequence $\mathrm{g}_{\mathrm{k}, \mathrm{m}}$ satisfying (1) and (2). So if we assume only that $\mathrm{g}_{\mathrm{k}}$ satisfies (4) and is the gradient of some functional $G_{\mathbf{k}} \in \mathrm{C}^{1}\left(\mathrm{H}^{1}, \mathbf{R}\right)$, we have to find some approximate sequence satisfying (1) and (2), and this is not obvious. For example if we think of $g_{k, m}(u)=\rho_{m} * g_{k}\left(\rho_{m} * u\right)$ where $\rho_{m}$ is a sequence of mollifiers, then $g_{k, m}$ will satisfy (1) to (4) except that the convergence in (2) will not be uniform on bounded sets of $\mathrm{H}^{1}$. However, let us point out that the approximating sequence is easily found in the important examples (see below).

REMARK 4. The solution of (NLS) does not need to satisfy the conservation of charge. However, it we assume in addition to the other hypotheses that $<g(u), i u>=0$ for every $u \in H^{1}$, then we get conservation of charge. This is easily seen by multiplying the equation by iu, in the sense of the duality between $\mathrm{H}^{-1}$ and $\mathrm{H}^{1}$.

REMARK 5. Some global existence results are easily obtained from property (ii) of Theorem 1, conservation of energy and conservation of charge (if any). See [5,6,7,8], [9] and the examples given below.

Let us now give some examples of nonlinearities that satisfy the hypotheses of Theorem 1.

EXAMPLE 1 (external potential). Let $V$ be a real-valued function on $\mathbf{R}^{n}$. Assume that $V \in L^{\sigma}+L^{\infty}$, with $\sigma \geq 1, \sigma>n / 2$. Let $g$ be given by $g(u)=V u$ for $u \in H^{1}$ and let $V=V_{1}+V_{2}$, with $V_{1} \in L^{\sigma}$ and $V_{2} \in L^{\infty}$. Then $g=g_{1}+g_{2}$ with $g_{1}(u)=V_{1} u$ and $g_{2}(u)=V_{2} u$. Now choose $r_{1}=p_{1}=2 \sigma /(\sigma-1), r_{2}=\rho_{2}=2$, and set $\mathrm{V}_{1, \mathrm{~m}}(\mathrm{x})=\operatorname{Min}\left\{\mathrm{m}, \operatorname{Max}\left\{-\mathrm{m}, \mathrm{V}_{1}(\mathrm{x})\right\}\right\}, \mathrm{g}_{1, \mathrm{~m}}(\mathrm{u})=\mathrm{V}_{1, \mathrm{~m}} \mathrm{u}$ and $\mathrm{g}_{2, \mathrm{~m}}=\mathrm{g}_{2}$. It is easily verified that g satisfies the hypotheses of Theorem 1 with

$$
G_{k, m}(\mathrm{u})=\frac{1}{2} \int_{\mathbf{R}^{\mathrm{n}}} \mathrm{~V}_{\mathrm{k}, \mathrm{~m}}(\mathrm{x})|\mathrm{u}(\mathrm{x})|^{2} \mathrm{dx}
$$

In this case, we always have global existence and conservation of charge.

EXAMPLE 2 (Hartree-type nonlinearity, see also [7]). Let $W$ be a real valued even function on $R^{\text {n }}$. Assume that $W \in L^{\delta}+L^{\infty}$ with $\delta \geq 1, \delta>n / 4$. Let $g$ be given by $g(u)=\left(W *|u|^{2}\right) u$ and let $W=W_{1}+W_{2}$, with $W_{1} \in L^{\delta}$ and $W_{2} \in L^{\infty}$. Then $g=g_{1}+g_{2}$ with $g_{1}(u)=\left(W_{1} *|u|^{2}\right) u$ and $g_{2}(u)=\left(W_{2} *|u|^{2}\right) u$. Now choose
$\mathrm{r}_{1}=\rho_{1}=4 \delta /(2 \delta-1), \mathrm{r}_{2}=\rho_{2}=2$, and set $\mathrm{W}_{1, \mathrm{~m}}(\mathrm{x})=\operatorname{Min}\left\{\mathrm{m}, \operatorname{Max}\left\{-\mathrm{m}, \mathrm{W}_{1}(\mathrm{x})\right\}\right\}, \mathrm{g}_{1, \mathrm{~m}}(\mathrm{u})=\left(\left.\mathrm{W}_{1, \mathrm{~m}^{*} \mid \mathrm{u}}\right|^{2}\right) \mathrm{u}$ and $\mathrm{g}_{2, \mathrm{~m}}=\mathrm{g}_{2}$. Applying Young's inequality, it is easily seen that the hypotheses of Theorem 1 are fulfilled with

$$
\mathrm{G}_{\mathrm{k}, \mathrm{~m}}(\mathrm{u})=\frac{1}{4} \int_{\mathbf{R}^{\mathrm{n}}}\left(\mathrm{~W}_{\mathrm{k}, \mathrm{~m}} *|\mathrm{u}|^{2}\right)|\mathbf{u}|^{2} \mathrm{dx}
$$

In this case we always have conservation of charge. All the solutiuons are global if, for example, the negative part of $W$ belongs to $L^{v}+L^{\infty}$, with $v=1$ if $n=1, v>1$ if $n=2$ and $v=n / 2$ if $n \geq 3$.

EXAMPLE 3 (local nonlinearity, see also [5,6,8] and [9]). Let $\mathrm{f}: \mathbf{R}^{\mathrm{n}} \times \mathbf{C} \rightarrow \mathbf{C}$ be a measurable function. Assume that $\mathrm{f}(\mathrm{x}, 0)=0$ almost everywhere and that there exists $\mathrm{M} \geq 0$ and $\alpha \in[0,4 /(\mathrm{n}-2))(\alpha \in[0, \infty)$ if $n=1,2)$ such that $\left|f\left(x, z_{2}\right)-f\left(x, z_{1}\right)\right| \leq M\left(1+\left|z_{1}\right|^{\alpha}+\left|z_{2}\right|^{\alpha}\right)\left|z_{2}-z_{1}\right|$ for almost all $x \in R^{n}$ and all $z_{1}, z_{2} \in C$. Assume that $f(x, z)=(z|z|) f(x,|z|)$. Let $g$ be given by $g(u)(x)=f(x, u(x))$. Then $g$ satisfies the hypotheses of Theorem 1. Indeed, a family $g_{k, m}$ is easily found. For example, let $f_{1}$ and $f_{2}$ be given by $f_{1}(x, z)=f(x, z)$ if $|z| \leq 1, f_{1}(x, z)=z f(x, 1)$ if $|z| \geq 1, f_{2}=f-f_{1}$. Let $f_{1, m}=f_{1}$ and let $f_{2, m}$ be defined by $\mathrm{f}_{2, \mathrm{~m}}(\mathrm{x}, \mathrm{z})=\mathrm{f}_{2}(\mathrm{x}, \mathrm{z})$ if $|\mathrm{z}| \leq \mathrm{m}, \mathrm{f}_{2, \mathrm{~m}}(\mathrm{x}, \mathrm{z})=(\mathrm{z} / \mathrm{m}) \mathrm{f}_{2}(\mathrm{x}, \mathrm{m})$ if $|\mathrm{z}| \geq \mathrm{m}$. Let the function $\mathrm{F}_{\mathrm{k}, \mathrm{m}}$ be defined by

$$
\mathrm{F}_{\mathrm{k}, \mathrm{~m}}(\mathrm{x}, \mathrm{z})=\int_{0}^{|\mathrm{z}|} \mathrm{f}_{\mathrm{k}, \mathrm{~m}}(\mathrm{x}, \mathrm{~s}) \mathrm{ds}
$$

Then a suitable sequence $g_{k, m}$ is given by $g_{k, m}(u)(x)=f_{k, m}(x, u(x))$ and

$$
\mathrm{G}_{\mathrm{k}, \mathrm{~m}}(\mathrm{u})=\int_{\mathbf{R}^{\mathrm{n}}} \mathrm{~F}_{\mathrm{k}, \mathrm{~m}}(\mathrm{x}, \mathrm{u}(\mathrm{x})) \mathrm{dx}
$$

Here also we have conservation of charge. The solutions are global for all the initial data if for example

$$
\int_{0}^{s} f(x, \sigma) d \sigma \geq-C-C s^{\delta} \text { for all } s \geq 0 \text { and for some } \delta \in[0,1+4 / n) .
$$

EXAMPLE 4. It is quite clear that if $\mathrm{g}^{1}, \mathrm{~g}^{2}, \ldots, \mathrm{~g}^{j}$ satisfy the hypotheses of the Theorem, then so does $\mathrm{g}=\mathrm{g}^{1}+\mathrm{g}^{2}+\ldots+\mathrm{g}^{\mathrm{j}}$. Therefore Theorem 1 applies when the nonlinearity is any finite sum of the nonlinearities considered in the examples 1,2 and 3 .

## 3 SKETCH OF THE PROOF.

The dispersive properties of the Schrödinger equation that we need are described in the following Lemma.

LEMMA 1. Let $r, p \in[2,2 n /(n-2)$ ) ( $r, \rho \in[2, \infty]$ if $n=1$ and $r, \rho \in[2, \infty)$ if $n=2$ ) and $q, \gamma \in(2, \infty]$ $\left(\mathrm{q}, \gamma \in[2, \infty]\right.$ if $\mathrm{n}=1$ ) with $2 / \mathrm{q}=\mathrm{n}(1 / 2-1 / \mathrm{r})$ and $2 / \gamma=\mathrm{n}(1 / 2-1 / \rho)$. Let $T>0, u_{0} \in \mathrm{~L}^{2}$ and $f \in \mathrm{~L}^{\gamma}\left(0, T, \mathrm{~L}^{\rho}\left(\mathbf{R}^{\mathrm{n}}\right)\right.$ ). Then there exists $C$ depending only on $n, r, \rho$ such that the solution $u$ of

$$
\begin{equation*}
\mathrm{iu}_{\mathrm{t}}+\Delta \mathrm{u}=\mathrm{f}, \mathrm{u}(0)=\varphi, \tag{LS}
\end{equation*}
$$

satisfies

Lemma 1 is proved in [8] for $f=0$ and in [10] (see also [9]) in the special cases $r=\rho, r=2, p=2$. The general case follows by interpolating between two of these three cases, depending on whether $r>\rho$ or $r<\rho$ (see [2] for a suitable interpolation theorem).

REMARK 6. It is immediate from Lemma 1 that if $f=f_{1}+\ldots+f_{M}$, where each of the $f_{j}$ satisfy the assumption of the lemma with exponents $\left(\gamma_{j}, \rho_{j}\right)$, then

Now we can proceed to prove Theorem 1. We assume that g satisfies the hypotheses and we consider $\varphi \in H^{1}$. For $\mathrm{m} \in \mathbf{N}$ we consider the solution $\mathbf{u}^{\mathrm{m}}$ of the problem

$$
\mathrm{i}\left(\mathrm{u}^{\mathrm{m}}\right)_{\mathrm{t}}+\Delta \mathrm{u}^{\mathrm{m}}=\mathrm{g}^{\mathrm{m}}\left(\mathrm{u}^{\mathrm{m}}\right) \quad, \quad \mathbf{u}^{\mathrm{m}}(0 ;)=\varphi(\cdot),
$$

(NLS,m)
where $g^{m}=g_{1, m}+\ldots+g_{N, m}$. It is not too difficult to prove (see [3]) that
$\mathrm{u}^{\mathrm{m}} \in \mathrm{C}\left(\left[0, \mathrm{~T}^{\mathrm{m}}\right) ; \mathrm{H}^{1}\right) \cap \mathrm{C}^{1}\left(\left[0, \mathrm{~T}^{\mathrm{m}}\right), \mathrm{H}^{-1}\right)$ for some $\mathrm{T}^{\mathrm{m}}>0$ and that we have

$$
\begin{equation*}
E^{\mathrm{m}}\left(\mathbf{u}^{\mathrm{m}}(\mathrm{t})\right)=\mathrm{E}^{\mathrm{m}}(\varphi), \text { for every } \mathrm{t} \in\left[0, \mathrm{~T}^{\mathrm{m}}\right) \tag{12}
\end{equation*}
$$

The next step is the following.

LEMMA 2. There exists $T_{1}>0$ depending on $\|\varphi\|_{H^{1}}$ such that $\left\|u^{m}\right\|_{L^{\infty}\left(0, T 1, H^{1}\right)} \leq 2\|\varphi\|_{H^{1}}$.

PROOF. Let $\left[0, \mathrm{~T}_{\mathrm{m}}\right]$ be the maximal interval on which $\left.\| u^{m}()_{i}\right)\left\|_{\mathrm{H}^{1}} \leq 2\right\| \varphi \|_{\mathrm{H}^{1}}$. All we need is a positive lower bound on $T_{m}$. Now it follows from (4), (1) and the equation that $\left\|\left(u^{m}\right)_{t}\right\|_{H^{-1}}$ is bounded on $\left[0, T_{m}\right]$ by some $K$ independent of $m$. Therefore, there exists a constant $K$ ' such that

$$
\begin{equation*}
\left\|u^{m}(t)-\varphi\right\|_{L^{2}} \leq K^{\prime} t^{1 / 2}, \text { for } t \in\left[0, T_{m}\right] \tag{13}
\end{equation*}
$$

From (12) we obtain

$$
\begin{equation*}
\left(\left\|u^{m}(t)\right\|_{\mathbf{H}^{1}}\right)^{2}=\left(\|\varphi\|_{\mathbf{H}^{1}}\right)^{2}+\mathrm{G}^{m}(\varphi)-\mathrm{G}^{\mathrm{m}}\left(\mathrm{u}^{\mathrm{m}}(\mathrm{t})\right)+\left(\left\|u^{m}(\mathrm{t})\right\|_{\mathrm{L}^{2}}\right)^{2}-\left(\|\varphi\|_{L^{2}}\right)^{2} . \tag{14}
\end{equation*}
$$

In (14) now, we estimate the $\mathrm{L}^{2}$ terms using (13) and the $\mathrm{G}^{\mathrm{m}}$ terms using (4). Together with Gagliardo-Nirenberg and Sobolev's inequalities, we get for some K" independent of m, some $\delta>0$, and for, say, $\mathrm{t} \leq 1$,

$$
\begin{equation*}
\left(\left\|\mathbf{u}^{\mathrm{m}}(\mathrm{t})\right\|_{\mathrm{H}^{1}}\right)^{2} \leq\left(\|\varphi\|_{\mathbf{H}^{1}}\right)^{2}+\mathrm{K}^{\prime \prime} \mathrm{t}^{\delta} . \tag{15}
\end{equation*}
$$

(See [3] for the details of this calculation.) For tless than some $\mathrm{T}_{0}>0$, the right-hand side of (15) is less than $4\left(\|\phi\|_{H^{1}}\right)^{2}$, and so $T_{m}$ is bounded from below by $T_{0}$.

The next Lemma is crucial for the passage to the limit.

LEMMA 3. $\mathrm{u}^{\mathrm{m}}$ is a Cauchy sequence in $\mathrm{C}\left(\left[0, \mathrm{~T}_{2}\right], \mathrm{L}^{2}\right)$ for some $\mathrm{T}_{2}>0$ depending on $\|\varphi\|_{\mathrm{H}^{1}}$.

PROOF. Let $j, m \in N$. From ( $N L S, j$ ) and ( $N L S, m$ ) we obtain $\left(u^{j}-u^{m}\right)(0)=0$ and

$$
i\left(u^{j}-u^{m}\right)_{t}+\Delta\left(u^{j}-u^{m}\right)=\sum_{k=1}^{N}\left(g_{k, j}\left(u^{j}\right)-g_{k, j}\left(u^{m}\right)\right)+\sum_{k=1}^{N}\left(g_{k, j}\left(u^{m}\right)-g_{k, m}\left(u^{m}\right)\right)
$$

Let $r \in[2,2 n /(n-2))(r \in[2, \infty)$ if $n=1,2)$ and $q \in(2, \infty]$ with $2 / q=n(1 / 2-1 / r)$. We apply Lemma 1 and Remark 6 to estimate ( $u^{\mathrm{j}}-u^{\mathrm{m}}$ ) in $\mathrm{L}^{\mathrm{q}}\left(0, \mathrm{~T}, \mathrm{~L}^{\mathrm{r}}\left(\mathrm{R}^{\mathrm{n}}\right)\right.$ ), for some $\mathrm{T} \leq \mathrm{T}_{1}$. To this end we introduce the exponents $\gamma_{k}$ given by $2 / \gamma_{k}=n\left(1 / 2-1 / \rho_{k}\right)$, and we estimate the terms $\left(g_{k, j}\left(u^{j}\right)-g_{k, j}\left(u^{m}\right)\right)$ and $\left(\mathrm{g}_{\mathrm{k}, \mathrm{j}}\left(\mathrm{u}^{\mathrm{m}}\right)-\mathrm{g}_{\mathrm{k}, \mathrm{m}}\left(\mathrm{u}^{\mathrm{m}}\right)\right)$ in $\mathrm{L}^{\left(\gamma_{k}\right)}\left(0, \mathrm{~T}, \mathrm{~L}^{(\mathrm{pk})^{\prime}}\left(\mathbf{R}^{\mathrm{n}}\right)\right)$. By Lemma 2 and (2), we obtain

$$
\sum_{\mathrm{k}=1}^{\mathrm{N}} \| \mathrm{g}_{\mathrm{k}, \mathrm{j}}\left(\mathrm{u}^{\mathrm{m}}\right)-\mathrm{g}_{\mathrm{k}, \mathrm{~m}}\left(\mathrm{u}^{\mathrm{m})} \|_{\mathrm{L}^{(\mathrm{kNj})}}\left(0, \mathrm{~T}, \mathrm{~L}^{(\mathrm{pk})}\right) \rightarrow 0 \text { as } \mathrm{j}, \mathrm{~m} \rightarrow \infty .\right.
$$

We estimate the terms $\left(\mathrm{g}_{\mathrm{k}, \mathrm{j}}\left(\mathrm{u}^{\mathrm{j}}\right)-\mathrm{g}_{\mathrm{k}, \mathrm{j}}\left(\mathrm{u}^{\mathrm{m}}\right)\right)$ by using (4) together with Hölder's inequality on $(0, \mathrm{~T})$. This yields

$$
\sum_{k=1}^{N}\left\|g_{k_{j, j}}\left(u^{j}\right)-g_{k, j}\left(u^{m}\right)_{L^{(\mu k)}\left(0, T, L^{(p k)}\right)} \leq C \sum_{k=1}^{N} T^{\alpha k}\right\| u^{j}-u^{m} \|_{L^{2 k}}\left(0, T, L^{i k}\right),
$$

where $C$ is some constant depending on $\|\varphi\|_{H^{1}}, q_{\mathrm{F}_{\mathrm{k}}}$ is given by $2 / \mathrm{q}_{\mathrm{k}}=\mathrm{n}\left(1 / 2-1 / \mathrm{r}_{\mathrm{k}}\right)$, and $\alpha_{\mathrm{k}}=\left(\mathrm{q}_{\mathrm{k}}-2\right) / \mathrm{q}_{\mathrm{k}}>0$. The choice of ( $q, r$ ) is arbitrary, so we choose successively $(q, r)=(\infty, 2)$, and $(q, r)=\left(q_{k}, r_{k}\right)$ for $\mathrm{k}=1, \ldots, \mathrm{~N}$. Adding the resulting inequalities, and choosing T small enough (depending on $\|\varphi\|_{\mathrm{H}^{1}}$ ), we find that

$$
M(T) \leq \varepsilon(j, m)+(1 / 2) M(T)
$$

where $\varepsilon(\mathrm{j}, \mathrm{m}) \rightarrow 0$ as $\mathrm{j}, \mathrm{m} \rightarrow \infty$ and

$$
\mathbf{M}(\mathbf{T})=\left\|u^{j}-u^{m}\right\|_{L^{\infty}\left(0, T, L^{2}\right)}+\sum_{k=1}^{N}\left\|u^{j}-u^{m}\right\|_{L^{\mathrm{qk}}\left(0, T, L^{\mathrm{rk}}\right)} .
$$

Hence the result.

We are now in position to complete the proof of Theorem 1. First, uniqueness is obtained with the same technique as in the proof of Lemma 3 since the $\mathrm{g}_{\mathbf{k}}$ 's satisfy the same estimates as the $\mathrm{g}_{\mathrm{k}, \mathrm{m}}$ 's do. Now we consider the sequence $\mathrm{u}^{\mathrm{m}}$ defined above, and we let $\mathrm{T}=\operatorname{Min}\left(\mathrm{T}_{1}, \mathrm{~T}_{2}\right)$ where $\mathrm{T}_{1}$ and $\mathrm{T}_{2}$ are given by Lemmas 2 and 3 . Let $u$ be the limit of $u^{m}$ in $\mathrm{C}\left(0, T, \mathrm{~L}^{2}\right)$. From the uniform $\mathrm{H}^{1}$ bound on the $u^{m}$, we get also $u \in L^{\infty}\left(0, T, H^{1}\right)$ and by Sobolev's inequality, $u^{m} \rightarrow u$ in $C\left(0, T, L^{\eta}\right)$ for any $r \in[2,2 n /(n-2))(r \in[2, \infty)$ if $n=1,2)$. Therefore, it follows from (2) and (4) that $g^{m}\left(u^{m}\right) \rightarrow g(u)$ in $\mathrm{C}\left(0, \mathrm{~T}, \mathrm{H}^{-1}\right)$. Thus u solves (NLS) in $\mathrm{L}^{\infty}\left(0, \mathrm{~T}, \mathrm{H}^{-1}\right)$. Now, from (4), (6), (12) and the weak lower semicontinuity of the $H^{1}$-norm we obtain that $E(u(t)) \leq E(\varphi)$ for $t \in[0, T]$. Reversing the sense of time and using uniqueness we obtain the same property for $w(s)=u(t-s), s \in[0, t]$, and in particular we get conservation of the energy (property (iii)). Therefore, the map $t \rightarrow\|u(t)\|_{H^{1}}$ is continuous. Since $u$ is weakly continuous in $\mathrm{H}^{1}$, it follows that in fact $\mathbf{u} \in \mathrm{C}\left(0, T, \mathrm{H}^{1}\right)$ and then also that $\mathbf{u} \in \mathrm{C}^{1}\left(0, T, \mathrm{H}^{-1}\right)$.

Thus we have established the existence of the solution described in Theorem 1 (all properties except (ii)) on an interval $[0, T]$ where $T$ depends only on $\|\varphi\|_{H^{1}}$. We now extend $u$ to be a maximal solution on $\left[0, \mathrm{~T}^{*}\right.$ ) and property (ii) follows easily. This proves Theorem 1.

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# The Cauchy problem for the Dirac equation with 

 cubic nonlinearity in three space dimensionsJoão-Paulo Dias and Mário Figueira
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1. Introduction

Let us consider the nonlinear Dirac equation in $\mathbb{R} \times \mathbb{R}^{3}$
(1.1) $i \frac{\partial \psi}{\partial t}=-i \alpha . \nabla \psi+m \beta \psi+k\left(\psi^{+} \beta \psi\right) \beta \psi$,

$$
\psi=\psi(t, x), m \geq 0, k \in \mathbb{R},
$$

where $\alpha \cdot \nabla \psi=\sum_{j=1}^{3} \alpha_{j} \frac{\partial \psi}{\partial x_{j}}$,
$\alpha_{1}=\left[\begin{array}{llll}0 & 0 & 0 & 1 \\ 0 & 0 & 1 & 0 \\ 0 & 1 & 0 & 0 \\ 1 & 0 & 0 & 0\end{array}\right], \alpha_{2}=\left[\begin{array}{rrrr}0 & 0 & 0 & -i \\ 0 & 0 & i & 0 \\ 0 & -i & 0 & 0 \\ i & 0 & 0 & 0\end{array}\right], \alpha_{3}=\left[\begin{array}{rrrr}0 & 0 & 1 & 0 \\ 0 & 0 & 0 & -1 \\ 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0\end{array}\right], \alpha_{4}=\beta=\left[\begin{array}{rrrr}1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1\end{array}\right]$
$\left(\alpha_{j}^{\dagger}=\bar{\alpha}_{j}^{\top}=\alpha_{j}, \alpha_{j}^{2}=I, \alpha_{j} \alpha_{m}=-\alpha_{m} \alpha_{j}\right.$ for $\left.j \neq m\right)$
and $\psi: \mathbb{R} \times \mathbb{R}^{3} \rightarrow \mathbb{C}^{4}$ is a column vector $\left(\psi_{1}, \psi_{2}, \psi_{3}, \psi_{4}\right), \psi^{\dagger}=\bar{\psi}^{\top}$. We put
$|\psi|^{2}=\psi^{\dagger} \psi$. The operator $A=-i \alpha \cdot \nabla+m \beta$ is self-adjoint in $\mathbb{R}^{2}=\left(L^{2}\left(\mathbb{R}^{3}\right)\right)^{4}$ with domain $\mathbb{H}^{1}$. We can consider the unitary group $S(t)=e^{-i t A}(t \in \mathbb{R})$ in $\mathbb{I H}^{s}, \$ \geq 0$, and so we write the Cauchy problem for the equation (1.1) in the following integral form
(1.2) $\psi(t)=S(t) \psi_{0}-i \int_{0}^{t} S(t-\tau) J(\psi(\tau)) d \tau$,
where $\psi_{0} \in \mathbb{H}^{s}, s \geq 2$, and $J \psi=k\left(\psi^{\dagger} \beta \psi\right) \beta \psi$. The function $w(t, x)=S(t) \psi_{0}$ veryfies the Klein-Gordon equation (wave equation if $m=0$ ) with

$$
\frac{\partial}{\partial t}\left[S(t) \psi_{0}\right]_{t=0}=-\alpha . \nabla \psi_{0}-i m \beta \psi_{0} .
$$

We recall that a solution $\psi \in \mathbb{C}\left(\mathbb{R} ; \mathbb{H}^{\prime}\right)$ of (1.1) satisfies the conservation laws
(1.3) $|\psi(t)|_{2}=\left|\psi_{0}\right|_{2}, t \in \mathbb{R}$.
(1.4) $\operatorname{Im} \int \psi^{+}(t) \alpha . \nabla \psi(t) d x+m \int \psi^{\dagger}(t) \beta \psi(t) d x+$

$$
+\frac{k}{2} \int\left(\psi^{+}(t) \beta \psi(t)\right)^{2} d x=(\text { energy), } t \in \mathbb{R},
$$

where $\int=\int_{\mathbb{R}} 3$. The energy will not be useful in our estimates. A local existence and uniqueness theorem for the cauchy problem for equation (1.1) is easy to prove if $\psi_{0} \in \mathbb{H}^{2}$. In [7], M. Reed has obtained a global existence result for equation (1.1) with higher nonlinearity, if $m>0, \psi_{0} \in H^{3}$ and $\left\|\psi_{0}\right\|=\left|\psi_{0}\right|_{H^{3}}+\sup _{t \in \mathbb{R}}\left[(1+|t|)^{3 / 2}\left|S(t) \psi_{0}\right|_{\infty}\right]$ is smal 1 enough (cf. [7], theorem 2.2). The cubic case (if m>0) can also be included in his result (cf. [3]). Furthermore, if m>0 and $\psi_{0} \in \mathbb{H}^{2+s}$, $0<s<1$, we can modify the work of $M$. Reed in order to obtain an existen ce result of the same kind, with $\frac{3}{2}$ replaced by $\left(1+\frac{1}{2} s\right)(1-\varepsilon), \varepsilon>0$ such that $\left(1+\frac{1}{2} s\right)(1-\varepsilon)^{2}>1(c f$. [3]). In $\S 3$ of this paper we extend this result to $\psi_{0} \in \mathbb{H}^{2}$, replacing the $\mathbb{L}^{\infty}$ norm by an $\mathbb{N} \mathbb{N}, \mathrm{p}$ norm, $3<p \leq \frac{10}{3}$, and applying the estimates of Ph. Brenner for the 1 inear Klein--Gordon equation (cf. [2]). In §2 we study the null mass case (m=0). In this case we need $\psi_{0} \in \mathbb{H}^{3}$ but, as $J$. Ginibre and $G$. Velo pointed out to us, this condition can be weakned in the framework of Besov spaces. Then, replacing the $\mathbb{I}^{\infty}$ norm by an $\mathbb{W}^{1}, p$ norm, $4<p<+\infty$, and applying the estimates of $H$. Pecher for the linear wave equation (cf. [5] and [61) we obtain a global existence result for the cauchy problem if

$$
\left.\|\psi\|_{\Sigma}=\left|\psi_{0}\right|_{\mathbb{H}} 3+\sup _{t \in \mathbb{R}}\left[(1+|t|)^{1-\frac{2}{p}}\left|S(t) \psi_{0}\right|_{\mathbb{W}}\right], p\right]
$$

is small enough (cf. [4]).
2. The null mass case $(m=0)$.

Assume $m=0, \phi \in \mathbb{H}^{3}$; Let $\dot{H}^{s}, p^{\prime}\left(\mathbb{R}^{3}\right), s \in \mathbb{R}, 4<p<+\infty, \frac{1}{p}+\frac{1}{p^{\prime}}=1$, be the completion of $D\left(\mathbb{R}^{3}\right)$ with respect to $\left|F^{-1}\left(|\xi|^{s} \hat{f}(\xi)\right)\right|_{p}, f \in \mathcal{D}\left(\mathbb{R}^{3}\right)$. If $\phi \in \mathbb{H}^{s, p^{\prime}}=\left(\dot{H}^{s, p^{\prime}}\left(\mathbb{R}^{3}\right)\right)^{4}$ with $s=2-\frac{4}{p}$, we have (cf. [5] and [6]): $|S(t) \phi|_{p} \leq c|t|^{-1+\frac{2}{p}}\left(|\phi|_{\mathcal{H}^{B}}, p^{\prime}+|\alpha \cdot \nabla \phi|_{\mathcal{H}^{H} s-1, p^{\prime}}\right) ; t \neq 0$.
 completion of $\mathcal{D}\left(\mathbb{R}^{3}\right)$ with respect to $\left.\left|F^{-1}\left[\left(1+|\xi|^{2}\right)^{s / 2} \hat{f}(\xi)\right]\right|\right|^{\prime}$, we obtain

Furthermore, since $|S(t) \phi|_{\mathbb{N}} 1, p \leq c|S(t) \phi|_{\mathbb{H}^{3}}=c|\phi|_{\mathbb{H}} 3$, we have

$$
\begin{equation*}
|S(t) \phi|_{\mathbb{N}^{\prime}} 1, \mathrm{p} \leq c|\phi|_{\mathbb{H}^{3}}, \mathrm{t} \in \mathbb{R} . \tag{2.2}
\end{equation*}
$$

Hence, by (2.1) and (2.2), we obtain

if $\phi \in \mathbb{H}^{3} \cap \mathbb{N}^{3}, p^{\prime}$ and $t \in \mathbb{R}$.
We need the following result:
Lemma 2.1: Assume $\psi_{1}, \psi_{2} \in \mathbb{H}^{3}, J \psi=k\left(\psi^{+} B \psi\right) \beta \psi$. We have
(2.4) $\left|J \psi_{1}-J \psi_{2}\right|_{\mathbb{H}} \leq \operatorname{cr}\left(\psi_{1}, \psi_{2}\right)$, if $4<\mathrm{p}<+\infty$,
with $\gamma\left(\psi_{1}, \psi_{2}\right)=\left[\left(\left|\psi_{1}\right|_{\mathbb{N}}^{2} 1, p+\left|\psi_{2}\right|_{\mathbb{N}}^{2}\right], p\right)\left|\psi_{1}-\psi_{2}\right|_{\mathbb{H}^{3}}+$
$\left.\left.+\left(\left|\psi_{1}\right|_{\mathbb{H}}+\left|\psi_{2}\right|_{\mathbb{H}} 3\right)\left(\left|\psi_{1}\right|_{\mathbb{N}} 1,\left.p^{+\mid \psi_{2}}\right|_{\mathbb{N}}\right\rceil, \mathrm{p}\right)\left|\psi_{1}-\psi_{2}\right|_{\mathbb{N}} 1, \mathrm{p}\right]$.
(2.5) $\left|J \psi_{1}-J \psi_{2}\right|_{\mathbb{W}} 3, p^{\prime} \leq \mathrm{cy}\left(\psi_{1}, \psi_{2}\right)$ if $4<p \leq 6$.
(2.6) $\left|J \psi_{1}-J \psi_{2}\right|_{\mathbb{M}^{3}, p^{\prime} \leq} \leq$

$$
\begin{aligned}
& \leq c\left[\left(\left|\psi_{1}\right|_{\mathbb{N}} 1, p+\left|\psi_{2}\right|_{\mathbb{N} 1, p}\right)\left(\left|\psi_{1}\right|_{\mathbb{N} 1, r}+\left|\psi_{2}\right|_{\mathbb{N}} 1, r\right)\left|\psi_{1}-\psi_{2}\right|_{\mathbb{H}} 3+\right. \\
& \left.\quad+\left(\left|\psi_{1}\right|_{\mathbb{H}} 3+\left|\psi_{2}\right|_{\mathbb{H}} 3\right)\left(\left|\psi_{1}\right|_{\mathbb{N}} 1, r+\left|\psi_{2}\right|_{\mathbb{N}} 1, r\right)\left|\psi_{1}-\psi_{2}\right|_{\mathbb{N}} 1, p\right], \\
& \text { if } 6<p<+\infty, \text { with } \frac{1}{r}=\frac{1}{2}-\frac{2}{p} .
\end{aligned}
$$

Proof: Let us take the "simplified model" $J u=u^{3}, u \in H^{3}\left(\mathbb{R}^{3}\right)$ real and let us estimate $D^{3}\left(u^{3}-v^{3}\right)$, where $D^{3}$ is a third derivative. Let $w=u-v$. The terms of $D^{3}\left(u^{3}-v^{3}\right)$ are of the tivpe $\theta_{1}=D^{3} \omega u v, \theta_{2}=D^{2} \omega D u v$, $\theta_{3}=D \omega D^{2} u v, \theta_{4}=D \omega D u D v, \theta_{5}=\omega D^{3} u v, \theta_{6}=\omega D^{2} u D v$.
Let $4<p<+\infty$. We have, by Holder's inequality and Sobolev's imbedding theorem:

$$
\begin{aligned}
& \left|\theta_{1}\right|_{2} \leq\left|D^{3} \omega\right|_{2}|u|_{\infty}|v|_{\infty} \leq c \gamma_{1}(\omega, u, v), \\
& \left|\theta_{2}\right|_{2} \leq\left|D^{2} \omega\right|_{s}|D u|_{p}|v|_{\infty} \leq c \gamma_{1}(\omega, u, v),
\end{aligned}
$$

with $\gamma_{7}(\omega, u, v)=|\omega|_{H^{3}}|u|_{W} 1, p|v|_{W} 1, p$ and $\frac{1}{s}=\frac{1}{2}-\frac{1}{p}$,

$$
\begin{aligned}
& \left|\theta_{3}\right|_{2} \leq|D \omega|_{p}\left|D^{2} u\right|_{s}|v|_{\infty} \leq c \gamma_{1}(u, \omega, v), \\
& \left|\theta_{4}\right|_{2} \leq|D \omega|_{p}|D u|_{r}|D v|_{p} \leq c \gamma_{7}(u, \omega, v), \\
& \left|\theta_{5}\right|_{2} \leq|\omega|_{\infty}\left|D^{3} u\right|_{2}|v|_{\infty} \leq c \gamma_{7}(u, \omega, v),
\end{aligned}
$$

$$
\left|\theta_{6}\right|_{2} \leq|\omega|_{\infty}\left|D^{2} u\right|_{s}|D v|_{p} \leq c \gamma_{1}(u, \omega, v),
$$

with $\frac{1}{s}=\frac{1}{2}-\frac{1}{p}$ and $\frac{1}{r}=\frac{1}{2}-\frac{2}{\mathrm{p}}$.
Now, let $4<p<6, \frac{1}{p^{\prime}}+\frac{1}{p}=1$. We have

$$
\begin{aligned}
& \left|\theta_{1}\right|_{p^{\prime}} \leq\left.\left|D^{3} \omega_{2}\right| u\right|_{q}|v|_{q} \leq c \gamma_{1}(\omega, u, v), \\
& \left|\theta_{2}\right|_{p^{\prime}} \leq\left.\left|D^{2} \omega_{2}\right| D u\right|_{p}|v|_{r} \leq c \gamma_{1}(\omega, u, v), \\
& \left|\theta_{3}\right|_{p^{\prime}} \leq|D \omega|_{p}\left|D^{2} u\right|_{2}|v|_{r} \leq c \gamma_{1}(u, \omega, v), \\
& \left|\theta_{4}\right|_{p^{\prime}} \leq|D \omega|_{p}|D u|_{p}|D v|_{r_{1}} \leq c \gamma_{1}(v, \omega, u), \\
& \left|\theta_{5}\right|_{p^{\prime}} \leq|\omega|_{r}\left|D^{3} u\right|_{2}|v|_{p} \leq c \gamma_{1}(u, \omega, v), \\
& \left|\theta_{6}\right|_{p^{\prime}} \leq|\omega|_{r}\left|D^{2} u\right|_{2}|D v|_{p} \leq c \gamma_{1}(u, \omega, v),
\end{aligned}
$$

with $\frac{1}{q}=\frac{1}{4}-\frac{1}{2 p} ; \frac{1}{r}=\frac{1}{2}-\frac{2}{p}$ and $\frac{1}{r_{1}}=1-\frac{3}{p}$.
Finally, let $6<p<+\infty$. We have,

$$
\begin{aligned}
& \left|\theta_{1}\right|_{p^{\prime}} \leq\left|D^{3} \omega\right|_{2}|u|_{p}|v|_{r} \leq c \gamma_{2}(\omega, u, v), \\
& \left|\theta_{2}\right|_{p} \leq\left|D^{2} \omega\right|_{2}|D u|_{p}|v|_{r} \leq c \gamma_{2}(\omega, u, v), \\
& \left.\left|\theta_{3}!_{p^{\prime}} \leq|D \omega|_{p}\right| D^{2} u\right|_{2}|v|_{r} \leq c \gamma_{2}(\omega, u, v), \\
& \left|\theta_{4}\right|_{p^{\prime}} \leq|D \omega|_{p}|D u|_{2}|D v|_{r} \leq c \gamma_{2}(u, \omega, v), \\
& \left|\theta_{5}\right|_{p^{\prime}} \leq|\omega|_{p}\left|D^{3} u\right|_{2}|v|_{r} \leq c \gamma_{2}(u, \omega, v), \\
& \left|\theta_{6}\right|_{p^{\prime}} \leq|\omega|_{p}\left|D^{2} u\right|_{2}|D v|_{r} \leq c \gamma_{2}(u, \omega, v),
\end{aligned}
$$

with $\gamma_{2}(\omega, u, v)=|\omega|_{H^{3}}|u|_{W} 1, p|v|_{W} 1, r$ and $\frac{1}{r}=\frac{1}{2}-\frac{2}{p}$.
Notice that, if $6<p<+\infty$, we have $6>r>2$ and $\left(1-\frac{2}{p}\right)+\left(1-\frac{2}{r}\right)=1+\frac{2}{p}>1$. W Now, let $4<p<+\infty, \psi \in C\left(\mathbb{R} ; H^{3}\right)$. We put
(2.7) $\|\psi\|=\sup _{t \in \mathbb{R}}|\psi(t)|_{\mathbb{H}^{3}}{ }^{+} \sup _{t \in \mathbb{R}}\left[(1+|t|)^{1-\frac{2}{p}}|\psi(t)|_{\mathbb{W}} 1, p\right]$

Let $\Sigma=\left\{\phi \in \mathbb{H}^{3}!\|\phi\| \|_{\Sigma}<+\infty\right\}$, where
(2.8) $\left.\|\phi\|_{\Sigma}=\|S(t) \phi\|=|\phi|_{\mathbb{H}^{3}}+\sup _{t \in \mathbb{R}}\left[(1+|t|)^{1-\frac{2}{p}}|S(t) \phi|_{\mathbb{W}}\right], p\right]$

By (2.3), if $\phi \in \mathbb{H}^{3} \cap \mathbb{W}^{3}, p^{\prime}, \frac{1}{\rho^{1}}+\frac{1}{p}=1$, we have $\phi \in \Sigma$ and

$$
\|\phi\|_{\Sigma \leq} \leq\left(\left|\phi \|_{\mathbb{M}^{3}}+|\phi|_{\mathbb{N}^{3}, \mathrm{p}}\right)\right.
$$

We can now prove the following theorem (cf. [4]):
Theorem 2.1: Assume $\psi_{0} \in \mathbb{H}^{3}$ and suppose that, for $\left.p \in\right] 4,+\infty\left[,\left\|\psi_{0}\right\|_{\Sigma}\right.$ is small enough. Then, there exists an unique solution $\psi$ of the equation (1.1) $(m=0)$ verifying $\psi(0)=\psi_{0}$ and such that $\psi \in \mathbb{C}\left(\mathbb{R} ; \mathbb{H}^{3}\right) \cap C^{1}\left(\mathbb{R} ; \mathbb{H}^{2}\right)$. Furthermore, $\|\psi\|<+\infty$.
Proof: Let $\eta>0$ and $\psi_{0} \in \Sigma$ such that $\|\psi\|_{\Sigma \Sigma \eta \text {. We put }}$

$$
X\left(\eta, \psi_{0}\right)=\left\{\psi \in \mathcal{C}\left(\mathbb{R} ; \mathbb{H}^{3}\right)!\left\|\psi-S(t) \psi_{0}\right\| \leq n\right\},
$$

which is a complete metric space for the distance $\underset{t}{d}\left(\psi_{1}, \psi_{2}\right)=\left\|\psi_{1}-\psi_{2}\right\|$. We have $\|\psi\| \leq 2 \eta$ for $\psi \in X\left(\eta, \psi_{0}\right)$. Define $(T \psi)(t)=-i \int_{0}^{t} S(t-\tau) j(\psi(\tau)) d \tau$, $(M \psi)(t)=S(t) \psi_{0}+(T \psi)(t)$, for $\psi \in X\left(n, \psi_{0}\right)$, with ${ }^{0} J \psi=k\left(\psi^{+} \beta \psi\right) \beta \psi$. We assume that $4<p<6$, since the proof in the case $6<p<+\infty$ follows by an easy adaptation (cf. lemma 2.1). By reasons of symmetry we take $t \geq 0$. We have, by (2.4) and since $2-\frac{4}{p}>1$,

$$
\begin{gathered}
|(T \psi)(t)|_{H^{3}} \leq \int_{0}^{t}|J(\psi(\tau))|_{\mathbb{H}}{ }^{3} d \tau \leq \\
\leq c_{2}(2 \eta)^{3} \int_{0}^{t}(1+|\tau|)^{-2+\frac{4}{p}} d \tau \leq c_{3} \eta^{3} . \\
\text { Now, by }(2.3),(2.4) \text { and }(2.5) \text {, we deduce }
\end{gathered}
$$

$$
\begin{aligned}
& |(T \psi)(t)|_{\mathbb{W}} 1,0 \leq \int_{0}^{t}|S(t-\tau) J(\psi(\tau))|_{\mathbb{W}} 1, p d \tau \leq \\
& \leq c \int_{0}^{t}(1+|t-\tau|)^{-1+\frac{2}{p}}\left(|J(\psi(\tau))|_{\mathbb{H}^{3}}+|J(\psi(\tau))|_{\mathbb{N}^{3}, p^{\prime}}\right) d \tau \\
& \leq c_{A} \int_{0}^{t}(1+|t-\tau|)^{-1+\frac{2}{p}|\psi(\tau)|_{\mathbb{H}}|\psi(\tau)|_{\mathbb{M}}^{2} 1, p} d \tau \leq \\
& \leq c_{4}(2 \eta)^{3} \int_{0}^{t}(1+|t-\tau|)^{-1+\frac{2}{p}}(1+|\tau|)^{-2+\frac{4}{p}} d \tau \leq \\
& \leq c_{5} \eta^{3}(1+|t|)^{-1+\frac{2}{p}} \text {, by the } 1 \text { emma in page } 78 \text { of [7]. Hence, } \\
& \|T \psi\| \leq c_{6} \eta^{3} \text { if } \psi \in X\left(n, \psi_{0}\right) \text { and so }\left\|M \psi-S(t) \psi_{0}\right\| \leq c_{6} \eta^{3} \text {. Let us take } \\
& n_{0}>0 \text { such that } c_{6} \eta_{0}^{2} \leq 1 \text {. Hence, for } n \leq \eta_{0} \text {, we have proved that }
\end{aligned}
$$

$M: X\left(n, \psi_{0}\right) \rightarrow X\left(\eta, \psi_{0}\right)$. By a similar argument we can prove that, if $\psi_{1}$, $\psi_{2} \varepsilon X\left(n, \psi_{0}\right)$ with $n \leq n_{0}$, we have

$$
\begin{aligned}
& !\left(M \psi_{1}\right)(t)-\left.\left(M \psi_{2}\right)(t)\right|_{\mathbb{H}} 3 \leq c_{8} n^{2}\left\|\psi_{1}-\psi_{2}\right\| \\
& !\left(M \psi_{1}\right)(t)-\left.\left(M \psi_{2}\right)(t)\right|_{\mathbb{N}} 1, p \leq c_{9} n^{2}\left\|\psi_{1}-\psi_{2}\right\|(1+|t|)^{-1+\frac{2}{p}}
\end{aligned}
$$

Hence !! $M \psi_{1}-M \psi_{2}!!\leq c_{10} \eta^{2} \| \psi_{1}-\psi_{2}!\mid$. By choosing $\eta_{1} \leq \eta_{0}$ such that $c_{10} \eta_{1}^{2}<1$, we can apply Banach's fixed point theorem and this completes the proof.
3. The massive case $(m>0)$.

Assume $m>0, \phi \in H^{2}, 3<p \leq \frac{10}{3}$ (hence, $3\left(\frac{1}{2}-\frac{1}{p}\right)>\frac{1}{2}$ and $\frac{3}{2}-\frac{5}{p} \leq 0$ ). We have (cf. [2] and [6] ) if $\phi \in \mathbb{H}^{s}, p^{\prime}$, with $s=1+\frac{3}{2}-\frac{5}{p}, \frac{1}{p^{1}}+\frac{1}{p}=1$,

Hence, since $W^{\top}, P^{\prime}=H^{\top}, P^{\prime} \subset H^{s, P^{\prime}}$, we have

$$
\left.|S(t) \phi|_{p \leq c} \leq|t|^{-3\left(\frac{1}{2}-\frac{1}{p}\right)}|\phi|_{\mathbb{W}}\right], p^{\prime},|t| \geq 1
$$

if $\phi \in \mathbb{N}^{1}, \mathrm{p}^{\prime}$, and so
(3.1) $|S(t) \phi|_{\mathbb{W}} 1, p \leq c|t|^{-3\left(\frac{1}{2}-\frac{1}{p}\right)}|\phi|_{\mathbb{I} W^{2}, p^{\prime}},|t| \geq 1$,
if $\phi \in \mathbb{N}^{2}, \mathrm{p}^{\prime}$.
Furthermore, since $\left.|S(t) \phi|_{\mathbb{W}}\right\urcorner, p \leq c|S(t) \phi|_{\mathbb{H}} 2=c|\phi|_{\mathbb{H}}{ }^{2}$, we have
(3.2) $\left.|S(t) \phi|_{\mathbb{N}}\right\}, p \leq c|\phi|_{\mathbb{H}} 2, \quad t \in \mathbb{R}$.

Hence, by (3.1) and (3.2), we obtain
(3.3) $|S(t) \phi|_{\mathbb{W}} 1, p \leq c(1+|t|)^{-3\left(\frac{1}{2}-\frac{1}{p}\right)}\left(|\phi|_{\mathbb{H}^{2}}+|\phi|_{\mathbb{W}^{2}, p^{1}}\right)$,
if $\phi \in \mathbb{H}^{2} \cap \mathbb{N}^{2}, p^{\prime}$ and $t \in \mathbb{R}$.
We need the following result:
Lemma 3.1: Assume $\psi_{1}, \psi_{2} \in \mathbb{H}^{2}, J \psi=k\left(\psi^{+} \beta \psi\right) \beta \psi$. We have
(3.4) $\max \left(\left|J \psi_{1}-J \psi_{2}\right|_{\mathbb{H}}{ }^{2},\left|J \psi_{T}-J \psi_{2}\right|_{\mathbb{N}^{2}, \mathrm{P}^{1}}\right) \leq$

$$
\begin{aligned}
& \leq c\left[\left(\left|\psi_{1}\right|_{\mathbb{N}}^{2} 1, p+\left|\psi_{2}\right|_{\mathbb{W} 1, p}^{2}\right)\left|\psi_{1}-\psi_{2}\right|_{\mathbb{H}} 2^{+}\right. \\
& \left.+\left(\left|\psi_{1}\right|_{\mathbb{H}^{2}}+\left|\psi_{2}\right|_{\mathbb{H}} 2\right)\left(\left|\psi_{1}\right|_{\mathbb{W}} 1, \mathrm{p}+\left|\psi_{2}\right|_{\mathbb{W} 1, p}\right)\left|\psi_{\mathbb{1}}-\psi_{2}\right|_{\mathbb{I W}], p}\right] .
\end{aligned}
$$

Proof: Let us take again the "simplified model" $J u=u^{3}, u \in H^{3}\left(\mathbb{R}^{3}\right)$ real and let us estimate $D^{2}\left(u^{3}-v^{3}\right)$ where $D^{2}$ is a second derivative. Let $\omega=u-v$. The terms of $D^{2}\left(u^{3}-v^{3}\right)$ are of the type $\theta_{1}=D^{2} \omega u v, \theta_{2}=D \omega D u v$, $\theta_{3}=\omega D^{2}$ uv. Let $3<p \leq \frac{10}{3}$. We have, by Holder's inequality and Sobolev's imbedding theorem:

$$
\begin{aligned}
& \left|\theta_{1}\right|_{2} \leq\left|D^{2} \omega\right|_{2}|u|_{\infty}|v|_{\infty} \leq c \gamma_{3}(\omega, u, v) \\
& \left|\theta_{2}\right|_{2} \leq|D \omega|_{s}|D u|_{p}|v|_{\infty} \leq c \gamma_{3}(\omega, u, v) \\
& \left|\theta_{3}\right|_{2} \leq|\omega|_{\infty}\left|D^{2} u\right|_{2}|v|_{\infty} \leq c \gamma_{3}(u, \omega, v),
\end{aligned}
$$

with $\gamma_{3}(\omega, u, v)=|\omega|_{H^{2}}|u|_{W} 1, p|v|_{W}, p$ and $\frac{1}{s}=\frac{1}{2}-\frac{1}{p}$.
Furthermore, we have, with $\frac{1}{p^{1}}+\frac{1}{p}=1$,

$$
\begin{aligned}
& \left|\theta_{7}\right|_{p^{\prime}} \leq\left|D^{2} \omega\right|_{2}|u|_{q}|v|_{q} \leq c r_{3}(\omega, u, v), \\
& \left|\theta_{2}\right|_{p^{\prime}} \leq|D \omega|_{p}|D u|_{p}|v|_{r_{1}} \leq c r_{3}(v, \omega, u), \\
& \left|\theta_{2}\right|_{p^{\prime}} \leq|\omega|_{q}\left|D^{2} u\right|_{2}|v|_{q} \leq c \gamma_{3}(u, w, v),
\end{aligned}
$$

with $\frac{1}{q}=\frac{1}{4}-\frac{1}{2 p}$ and $\frac{l}{r_{1}}=1-\frac{3}{p}$.
Now, let $3<p \leq \frac{10}{3}, \psi \in \in\left(\mathbb{R} ; H^{2}\right)$. We put
(3.8) ||! $\psi\left|\left|\left|=\sup _{t \in \mathbb{R}}\right| \psi(t)\right|_{\mathbb{H}} 2^{+} \sup _{t \in \mathbb{R}}\left[(1+|t|)^{3\left(\frac{1}{2}-\frac{1}{p}\right)}|\psi(t)|_{\mathbb{N}} 1, p\right]\right.$

Let $\tilde{\Sigma}=\left\{\phi \in \mathbb{H}^{2} \mid\|\phi\| \tilde{\Sigma}^{<+\infty}\right\}$, where
(3.9) $\|\phi\|_{\tilde{\Sigma}}=\|||S(t) \phi||\left|=|\phi|_{\mathbb{H}^{2}}+\sup _{t \in \mathbb{R}}\left[(1+|t|)^{3\left(\frac{1}{2}-\frac{1}{p}\right)}|S(t) \phi|_{\mathbb{W}} 1, p\right]\right.$

By (3.3) if $\phi \in \mathbb{H}^{2} \cap \mathbb{W}^{2}, p^{\prime}, \frac{1}{p^{1}}+\frac{1}{p}=1$, we have $\phi \in \tilde{\Sigma}$ and

$$
\|\phi\|_{\tilde{\Sigma}} \leq c\left(|\phi|_{\mathbb{H}^{2}}+|\phi|_{\mathbb{M}^{2}, p}\right) .
$$

The next result has a proof similar to that of theorem 2.1, based on 1emma 3.1:

Theorem 3.1: Assume $\psi_{0} \in \mathbb{H}^{2}$ and suppose that, for $\left.\left.p \in\right] 3, \frac{10}{3}\right]$, $\left\|\psi_{0}\right\|_{\tilde{\Sigma}}$ is small enough. Then, there exists an unique solution $\psi$ of the equation (1.1) ( $m>0$ ) verifying $\psi(0)=\psi_{0}$ and such that $\psi \in C\left(\mathbb{R} ; \mathrm{IH}^{2}\right) \cap C^{\top}\left(\mathbb{R} ; \mathbb{I H}^{1}\right)$. Furthermore, $\|\mid \psi \psi\|<+\infty$.

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# THE CAUCHY PROBLEM FOR THE NON-LINEAR KLEIN-GORDON EQUATION 

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## 1. INTRODUCTION

The purpose of this lecture is to summarize some recent perturbative results on the Cauchy problem for the non linear Klein-Gordon equation

$$
\begin{equation*}
\square \varphi \equiv \ddot{\varphi}-\Delta \varphi=-f(\varphi) \tag{1.1}
\end{equation*}
$$

Here $\varphi$ is a complex valued function defined in space time $\mathbb{R}^{n+1}$, the upper dot denotes the time derivative, $\Delta$ is the Laplace operator in $\mathbb{R}^{n}$ and $f$ is a non linear complex valued function. Only the case $n \geqq 2$ will be considered since the special case $n=1$ is simpler and would require slightly modified statements.

The existence and the properties of the solutions $\varphi$ depend in a crucial way on the initial data and on the non linear term $f$. The situation can be best illustrated by considering the following typical form of $f$

$$
\begin{equation*}
f(\varphi)=\lambda_{0} \varphi+\lambda \varphi|\varphi|^{p-1} \tag{1.2}
\end{equation*}
$$

with $\lambda_{0}, \lambda \in \mathbb{R}$ and $l<p<\infty$. In this case, if one restricts conveniently the space where to operate and therefore the space of initial data, the Cauchy problem for the equation (l.l) has a ique local (in time) solution. The extension of a local solution to a global one requires, in general, additional assumptions since, as it appears already in elementary examples, local solutions may fail to exist beyond a finite time. The standard procedure to prevent this kind of phenomenon makes

[^0]use, in an essential way, of the conservation of the energy and of some of its positivity properties. For this reason, while there is a large flexibility in the choice of the space where to solve the local Cauchy problem, the available proofs of existence of solutions for the global Cauchy problem always request the initial data to belong to the energy space $X_{e} \equiv\left(H^{2}\left(\mathbb{R}^{n}\right) \cap L^{p+1}\left(\mathbb{R}^{n}\right)\right) \oplus L^{2}\left(\mathbb{R}^{n}\right)$. These solutions will be called finite energy solutions. The positivity property of the energy depends crucially on the $\operatorname{sign}$ of $\lambda$ : if $\lambda<0$, finite energy solutions, known to exist for short time, can blow up; if $\lambda \geqq 0$, any element of $X_{e}$ can be taken as the initial condition of a global solution in $X_{e}$.

Existence of global solutions of the Cauchy problem in $X_{e}$ for $\lambda \geqq 0$ can be obtained by a compactness technique, which is inherently non perturbative in character. Instead, all proofs of uniqueness require a perturbative argument which can be implemented in $X_{e}$ under the assumption

$$
\begin{equation*}
p-1<4 /(n-2) . \tag{1.3}
\end{equation*}
$$

The condition (1.3) can be relaxed at the price of proving uniqueness in a space smaller than $X_{e}$, in which case, however, the spaces of existence and uniqueness do not match any more.

Here a presentation is given to the chain of perturbative arguments which leads to the existence and uniqueness in $x_{e}$ of global solutions of the Cauchy problem for the equation (l.l) under the assumption (1.3) in the situation where the energy is suitably bounded from below ( $\lambda \geqq 0$ for the example (l.2)) and for any $n \geqq 2$. Proofs of the statements, details and references to previous work are given in [3]. Additional useful information can be found in [2].

We now provide the main notation. We denote by $\left\|\left\|\|_{r}\right.\right.$ the norm in $L^{r}=L^{r}\left(\mathbb{R}^{n}\right)$. With each $r$ it is convenient to associate the variables $\gamma(r)$ and $\delta(r)$ defined by

$$
\gamma(r) /(n-1)=\delta(r) / n=1 / 2-1 / r .
$$

For each integer $k$ we denote by $H^{k} \equiv H^{k}\left(\mathbb{R}^{n}\right)$ the usual Sobolev spaces. We shall use the homogeneous Besov spaces of arbitrary order and the associated Sobolev inequalities, for which we refer to the Appendix both for information and references. We use the notation $\dot{\mathrm{B}}_{\mathrm{r}}^{\rho} \equiv \dot{\mathrm{B}}_{\mathrm{r} \boldsymbol{2}}^{0}\left(\mathbb{R}^{n}\right)$ for those spaces. For any interval $I$, for any Banach space $B$, we denote by $\mathscr{C}(I, B)$ the space of strongly continuous functions from $I$ to $B$. For any $q, l \leqq q \leqq \infty$, we denote by $L^{q}(I, B)\left(r e s p . L_{l o c}^{q}(I, B)\right)$,
the space of measurable functions $\varphi$ from $I$ to $B$ such that $\|\varphi(\cdot) ; B\|$ $\in L^{q}(I)\left(r e s p .\|\varphi(\cdot) ; B\| \in L_{l o c}^{q}(I)\right)$.

We shall need the operators $k(t)=(-\Delta)^{-1 / 2} \sin (-\Delta)^{1 / 2} t$ and $\dot{K}(t)=\cos (-\Delta)^{1 / 2} t ;$ both are bounded and strongly continuous with respect to $t$ in $H^{k}$ for any $k$.

## 2. THE LOCAL CAUCHY PROBLEM

The study of the Cauchy problem for the equation (1.1) can be conveniently replaced by the study of the integral equation

$$
\begin{equation*}
\varphi=A\left(t_{0}, \varphi^{(0)} ; \varphi\right) \tag{2.1}
\end{equation*}
$$

where $t_{0} \in \mathbb{R}, \varphi^{(0)}$ is a solution of the free wave equation

$$
\begin{equation*}
\square \varphi^{(0)}=0 \tag{2.2}
\end{equation*}
$$

and

$$
A\left(t_{0}, \varphi^{(0)} ; \varphi\right) \equiv \varphi^{(0)}+F\left(t_{0} ; \varphi\right)
$$

with

$$
\begin{equation*}
\left(F\left(t_{0} ; \varphi\right)\right)(t) \equiv-\int_{t_{0}}^{t} d \tau k(t-\tau) f(\varphi(\tau)) \tag{2.3}
\end{equation*}
$$

At a formal level any solution of the equation (2.1) is a solution of the equation (1.1) and, conversely, any solution of (1.1) solves (2.1) with a suitable $\varphi^{(0)}$ (solution of (2.2)) which contains the information on the initial data at time $t_{0}$. The integral (2.3) and the subsequent ones may be understood in various senses. Generally speaking, when the functions involved are sufficiently regular, they are ordinary integrals and the estimates they satisfy determine their extension to a larger class of functions. Their meaning should be understood from the context and will not be mentioned explicitely.

In order to prove uniqueness of the solutions of the equation (2.1), we build a space based on two norms chosen in such a way that $F\left(t_{0},.\right)$, restricted to the bounded sets of the first one, is contracting with respect to the second one. Those norms involve space time integrations and are suggested by the following fundamental estimate for $K(t)$ [8][6][4]

$$
\begin{equation*}
\|K(t) u\|_{r} \leqq C|t|^{l-\delta(r)+\delta(s)}\|u\|_{s} \tag{2.4}
\end{equation*}
$$

which holds for

$$
\begin{aligned}
& 0 \leqq \delta(r)-\delta(s) \leqq \operatorname{Min}\{l+\gamma(r), n(1-\gamma(r))\} \\
& 1<s, r<\infty \quad \text { if } n=2 .
\end{aligned}
$$

This motivates the following definitions. For any interval $I$ and suitable values of $\ell, q, r$ and $q_{l}$, we define the spaces

$$
x_{0}(I)=L^{q}\left(I, L^{\ell}\right), \quad x_{1}(I)=L^{q}\left(I, L^{r}\right)
$$

and similarly the local ones in time.
The basic assumption on $f$ is expressed by a power law estimate of the following type
(Al) $f \in \mathcal{C}^{l}(\mathbb{C}, \mathbb{C}), f(0)=0$ and for some $p, 1 \leqq p<\infty$ and all $z \in \mathbb{C}$

$$
\begin{equation*}
\left|f^{\prime}(z)\right| \equiv \operatorname{Max}\{|\partial f / \partial z|,|\partial f / \partial \bar{z}|\} \leqq c\left(1+|z|^{p-1}\right) \tag{2.5}
\end{equation*}
$$

The first important property. of $F\left(t_{0}\right.$,.) with respect to the previously defined spaces is contained in the next lemma.

Lemma 2.l. Let $f$ satisfy (Al), let $I$ be a bounded interval of ime, let $t_{0} \in I$ and let $\varphi_{1}, \varphi_{2} \in X_{1}$ (I) $\cap \mathcal{X}_{0}$ (I). Then

$$
\begin{align*}
& \left\|F\left(t_{0}, \varphi_{1}\right)-F\left(t_{0}, \varphi_{2}\right) ; X_{1}(I)\right\| \leqq c\left\|\varphi_{1}-\varphi_{2} ; X_{1}(I)\right\| \\
& \times\left\{|I|^{2}+|I|^{\eta_{1}} \underset{i=1,2}{\left.\left(\sum_{i}\left\|\varphi_{i} ; X_{0}(I)\right\|^{p-1}\right)\right\}}\right.  \tag{2.6}\\
& \text { provided } \ell, r, q, q_{1} \text { satisfy }
\end{align*}
$$

$$
\begin{align*}
& l \leqq \ell, r, q, q_{1} \leqq \infty \\
& 1<r<\infty \quad \text { if } n=2 ; \quad|\gamma(r)| \leqq 1 \quad \text { if } n \geqq 3  \tag{2.7}\\
& (p-1) n / \ell \leqq \operatorname{Min}\{1+\gamma(r), n(1-\gamma(r))\}  \tag{2.8}\\
& (p-1) / q+1 / q_{1} \leqq 1  \tag{2.9}\\
& n_{1} \equiv 2-(p-1)(n / \ell+1 / q)>0 . \tag{2.10}
\end{align*}
$$

The proof is a consequence of (2.4) and of Hölder's and Young's inequalities.

Lemma 2.1 implies immediately the following uniqueness result.

Proposition 2.1. Let $f$ satisfy (Al), let $\quad, \quad r, q, q_{1}$ satisfy (2.7)(2.10), let $I$ be an open interval, let $t_{0} \in I$ and let $\varphi^{(0)} \in \mathcal{X}_{1 \text { loc }}$ (I) $\cap$ $X_{010 \mathrm{c}}(I)$. Then the equation (2.1) has at most one solution in $X_{1 l o c}$ (I) $n$ $x_{0 \text { loc }}(I)$.

By choosing suitable values of $\ell, r, q, q_{1}$, the conditions (2.7)(2.10) can be satisfied for any $p$; in particular large values of $p$ require large values of $\ell$ and $q$. In this case, in general, finite energy solutions do not belong to $X_{0 l o c}(\mathbb{R})$ so that the previous proposition does not apply to this important class of solutions, unless an upper bound on $p$ is imposed (this bound is expressed by (1.3)).

For the existence of local solutions it is sufficient to show that balls of arbitrary radius $R$ in $X_{0}(I)$ are left invariant by $F\left(t_{0}\right.$, .) for some $I$ depending on $R$, or, more generally, that this happens for the balls of an appropriate space $X_{2}(I)$ continuously embedded in $X_{0}$ (I). It turns out that a convenient choice for this new space is

$$
x_{2}(I)=L^{q}\left(I, \quad \dot{B}_{r}^{\rho}\right)
$$

for suitable values of $\rho, r$ and $q$. The Sobolev inequalities (see the Appendix) imply that $X_{2}(I)$ is continuously embedded in $X_{0}(I)$ provided $n / r-\rho=n / \ell, \ell \geqq 2$ and $\ell \geqq r$. In order to show that $F\left(t_{0},.\right)$ reproduces the space $\mathscr{X}_{2}(I)$, the following generalization of (2.4) plays an important role :

$$
\begin{equation*}
\left\|K(t) u ; \dot{B}_{r}^{\rho}\right\| \leqq c|t|^{-\mu}\left\|u ; \dot{B}_{r^{\prime}}^{\rho^{\prime}}\right\| \tag{2.11}
\end{equation*}
$$

which holds for all $r$ with $0 \leqq \lambda(r) \leqq l$ and all $\rho, \rho^{\prime}, r^{\prime}, \mu$ such that
$0 \leqq 1+\mu=\rho+\delta(r)-\rho^{\prime}-\delta\left(r^{\prime}\right) \leqq \frac{1}{2}\left(\gamma(r)-\gamma\left(r^{\prime}\right)\right)(1+1 / \gamma(r)) \leqq 1+\gamma(r)$.

Lemma 2.2. Let $f$ satisfy (Al), let $I$ be a bounded interval of time, let $t_{0} \in I$ and let $\varphi \in X_{2}$ (I). Then
$\left\|F\left(t_{0}, \varphi\right) ; X_{2}(I)\right\| \leqq c\left\{\left.I\right|^{2}\left\|\varphi ; X_{2}(I)\right\|+|I|^{\eta}\left\|_{\varphi}\right\| X_{2}(I) \|^{p}\right\}$
provided $\rho, r$ and $q$ satisfy $0 \leqq \rho<l$ and

$$
\begin{align*}
& 0 \leqq \gamma(r) \leqq(n-1) /(n+1)  \tag{2.13}\\
& 0 \leqq(p-1)(n / r-p) \leqq 1+\gamma(r)  \tag{2.14}\\
& p \leqq q  \tag{2.15}\\
& \eta_{2} \equiv 2-(p-1)(n / r-\rho+1 / q)>0 \tag{2.16}
\end{align*}
$$

The proof is similar to that of Lemma $2 . l$ where (2.4) is replaced by (2.11). The only new ingredient is the Leibnitz rule adapted to fraction. al derivatives (See Lemma Al).

We can now state the basic local existence and uniqueness result for the solutions of the equation (2.l). We shall denote by $B_{2}(I, R)$ the closed ball of radius $R$ in $X_{2}(I)$.

Proposition 2.2. Let $f$ satisfy (Al), let $\rho, r, q$ and $q_{1}$ satisfy $0 \leqq \rho<1$, $1 \leqq q \leqq q_{1} \leqq \infty$ and (2.13)-(2.16). Then for any $R>0$, there exists $T(R)>0$ such that, for any $t_{0} \in \mathbb{R}$ and for any $\varphi^{(0)} \in B_{2}(I, R) \cap \mathcal{X}_{1}(I)$, where $I=\left[t_{0}-T(R), t_{0}+T(R)\right]$, the equation (2.1) has a solution in $B_{2}(I, 2 R) \cap \mathcal{X}_{1}(I)$ with $\left\|\varphi ; X_{1}(I)\right\| \leqq 2\left\|\varphi^{(0)} ; X_{1}(I)\right\|$. That solution is unique in $X_{2}$ (I) $\cap X_{1}$ (I).

If $n \geqq 4$, the assumptions of Proposition 2.2 impose on $p$ the upper limit

$$
(p-1)(n / 2-3 / 2-1 / n)<2
$$

which is obtained by taking $\rho \sim 1, \gamma(r)=(n-1) /(n+1)$ and $q=q_{1}=\infty$. This condition is not sufficient to ensure that finite energy solutions belong to $X_{2}(\mathbb{R})$ since $\rho, r$ and $q$ are too large. A more stringent upper limit on $p$ (given by (1.3)) is necessary in order to allow values of $\rho, r, q$ which permit to accomodate finite energy solutions in $X_{2}(\mathbb{R})$.

We now discuss the behaviour of the finite energy solutions.
We recall that the energy space $X_{e}$ has been already defined in the Introduction. If $p$ satisfies (1.3), the Sobolev inequalities imply that

$$
\begin{equation*}
x_{e}=\left\{\left(\varphi_{0}, \psi_{0}\right): \varphi_{0} \in H^{1}, \psi_{0} \in L^{2}\right\}=H^{1} \oplus L^{2} \tag{2.17}
\end{equation*}
$$

This is the expression to which from now on we will refer for $X_{e}$ since in all what follows (1.3) will be always supposed to hold.

With any $\left(\varphi_{0}, \psi_{0}\right) \in X_{e}$ we can construct the finite energy solution

$$
\begin{equation*}
\varphi^{(0)}(t)=\dot{k}\left(t-t_{0}\right) \varphi_{0}+k\left(t-t_{0}\right) \psi_{0} \tag{2.18}
\end{equation*}
$$

of the equation (2.2). This solution belongs to $\mathscr{C}\left(\mathbb{R}, H^{l}\right)$ and its space-time integrability properties are expressed by the following lemma [7][9][5][2].

Lemma 2.3 Let $\rho, r, q$ satisfy

$$
\begin{align*}
& 0 \leqq \delta(r) \leqq n / 2 \\
& -1 \leqq \sigma=\rho+\delta(r)-1<1 / 2  \tag{2.19}\\
& \sigma \leqq \gamma(r) / 2 \\
& 1 / q=\operatorname{Max}(0, \sigma) \tag{2.20}
\end{align*}
$$

Then, for any $\left(\varphi_{0}, \psi_{0}\right) \in X_{e}, \varphi^{(0)}$, as defined by (2.18), belongs to $X_{2}(\mathbb{R})$ and satisfies the estimate

$$
\begin{equation*}
\left\|\varphi^{(0)} ; X_{2}(\mathbb{R})\right\| \leqq c\left(\left\|\psi_{0}\right\|_{2}+\left\|\nabla \varphi_{0}\right\|_{2}\right) \tag{2.21}
\end{equation*}
$$

This leads to the basic local existence and uniqueness result concerning finite energy solutions.

Proposition 2.3 Let $f$ satisfy (Al) and (1.3). Then
(1) There exist $\rho, r$ and $q$ satisfying $0 \leqq \rho<1,(2.13)-(2.16)$ and (2.19) and (2.20).
Let $X_{1}$ and $X_{2}$ correspond to the previous values of $\rho, r, q$ and to $q_{1} \geqq q$. Then
(2) For any $\left(\varphi_{0}, \psi_{0}\right) \in x_{e}$ there exists $T>0$ depending only on
$\left\|\left(\varphi_{0}, \psi_{0}\right) ; x_{e}\right\|$ such that for any $t_{0} \in \mathbb{R}$, the equation (2.1) with $\varphi^{(0)}$ defined by (2.18) has a unique solution in $X_{1}(I) \cap X_{2}(I)$ where $I=$ $\left[\mathrm{t}_{0}-\mathrm{T}, \mathrm{t}_{0}+\mathrm{T}\right]$.
(3) For any $\left(\varphi_{0}, \psi_{0}\right) \in X_{e}$, for any interval $I$, for any $t_{0} \in I$, the equation (2.1) with $\varphi(0)^{\text {en }}$ defined by (2.18) has at most one solution in $x_{1 \text { loc }}(I) \cap x_{2 \text { loc }}(I)$.

Once the local Cauchy problem is solved, the next natural question consists in ascertaining whether the solutions of the equation (l.l) obtained by the previous perturbative technique can be extended to all times. This continuation is possible if we can find an a priori bound on the norms of the local solutions and, for this purpose, the energy is a fundamental quantity, provided it satisfies an appropriate positivity condition. The relevant assumption can be formulated in the following way :
(A2) There exists a function $V \in \mathcal{C}^{l}(\mathbb{C}, \mathbb{R})$ such that $V(0)=0, V(z)$ $=V(|z|)$ for all $z \in \mathbb{C}$ and $f(z)=\partial V / \partial \bar{z}$. For all $R>0, V$ satisfies the estimate

$$
\begin{equation*}
V(R) \geqq-a^{2} R^{2} \tag{3.1}
\end{equation*}
$$

for some $a \geqq 0$.

For $(\varphi, \psi) \in X_{e}$ and such that $V(\varphi) \in L^{l}$ the energy is defined by

$$
\begin{equation*}
E(\varphi, \psi)=\|\psi\|_{2}^{2}+\|\nabla \varphi\|_{2}^{2}+\int d x V(\varphi(x)) \tag{3.2}
\end{equation*}
$$

The first part of (A2), namely the relation between $f$ and $V$, implies, at least at the formal level, that

$$
\begin{equation*}
\frac{d}{d t} E(\varphi(t), \dot{\varphi}(t))=0 \tag{3.3}
\end{equation*}
$$

where $\varphi$ is a solution of the equation (1.l), whereas the lower boundedness condition (3.1) prevents that an infinite compensation between the kinetic and potential parts of the energy takes place. Those two facts altogether yield a uniform bound for $\varphi$ in $H^{1}$ at finite times in terms of the initial data. The actual proof of energy conservation proceeds through the following path. One first introduces suitable cut-offs both in the equation and in the initial data, then proves a regularized form of the conservation law, and finally deduces (3.3) by removing the cut-offs. The next proposition summarizes the situation in exact terms.

Proposition 3.1. Let f satisfy (A1), (1.3) and (A2). Let ( $\varphi_{0}, \psi_{0}$ ) $\in X_{e}$ let $I$ be an open interval and let $t_{0} \in I$. Let $\rho, r$ and $q$ satisfy $0 \leqq \rho<1$, (2.13)-(2.15), (2.19), (2.20) and let $q_{1} \geqq q$. Let $\varphi$ (0) be defined by
(2.18) and let $\varphi$ be a solution of the equation (2.1) in $X_{1}$ (I) $\cap X_{2}$ (I). Then $(\varphi, \dot{\varphi}) \in \mathscr{C}\left(I, H^{l} \oplus L^{2}\right)$ and $\varphi$ satisfies the conservation of the energy

$$
\begin{equation*}
E(\varphi(t), \dot{\varphi}(t))=E\left(\varphi_{0}, \psi_{0}\right) \equiv E \tag{3.4}
\end{equation*}
$$

and the estimates

$$
\begin{align*}
& \|\varphi(t)\|_{2} \leqq e\left(E, t-t_{0}\right)  \tag{3.5}\\
& \|\dot{\varphi}(t)\|_{2}^{2}+\|\nabla \varphi(t)\|_{2}^{2} \leqq \dot{e}\left(E, t-t_{0}\right)^{2} \tag{3.6}
\end{align*}
$$

where

$$
e(E, \tau)=\left\|\varphi_{0}\right\|_{2} \cosh a|\tau|+\left(E+a^{2}\left\|\varphi_{0}\right\|_{2}^{2}\right)^{1 / 2} a^{-1} \sinh a|\tau|
$$

We can now state the basic global existence and uniqueness result for finite energy solutions.

Proposition 3.2. Let $f$ satisfy (Al, (1.3) and (A2). let ( $\varphi_{0}, \psi_{0}$ ) $\in X_{e}$, let $t_{0} \in \mathbb{R}$ and let $\varphi^{(0)}$ be defined by (2.18). Then the equation (2.1) has a unique solution $\varphi$ such that $(\varphi, \dot{\varphi}) \in \mathscr{G}\left(\mathbb{R}, X_{e}\right)$. That solution satisfies the conservation of the energy (3.3) and the bounds (3.4) and (3.5).

Let $\rho, r, q$ and $q_{1}$ be as in Proposition 3.l. Then the solution is unique in $X_{1 l o c}(\mathbb{R}) \cap{\underset{X}{2 l o c}}^{(\mathbb{R}) \text {. }}$

The proof is an immediate consequence of previous results. The existence of global solutions follows from Proposition 2.3 and the a priori estimates of Proposition 3.1, the uniqueness in $X_{1 l o c}(\mathbb{R}) \cap X_{2 l o c}(\mathbb{R})$ follows from Proposition 2.2 and the uniqueness in $\mathscr{C}\left(\mathbb{R}, x_{e}\right)$ follows from the fact that any solution in $L_{l o c}^{\infty}\left(\mathbb{R}, X_{e}\right)$ belongs to $X_{1 l o c}(\mathbb{R}) \cap$ $X_{2 \mathrm{loc}}(\mathbb{R})$ (see Lemma 3.3 of [2]).

## APPENDIX

In this appendix we collect some basic facts about the homogeneous Besov and Sobolev spaces. Additional information can be found in [1], [10] and in the Appendix of [2].

We define the closed subspace $Z=Z\left(\mathbb{R}^{n}\right)$ of $\varphi=\varphi\left(\mathbb{R}^{n}\right)$ by

$$
Z=\left\{u \in \mathscr{Y}:\left(D_{\hat{u}}^{\alpha}\right)(0)=0 \text { for any multiindex } \alpha\right\} .
$$

The dual of the inclusion map $Z \subset \mathcal{Y}$ is a surjection $\pi$ from $\mathcal{Y}$, to $Z^{\prime}$ the kernel of which are the polynomials $P$, so that $Z^{\prime}=\mathscr{\varphi}^{1} / P$.

Let now $\psi \in \mathscr{Q}_{0}^{\infty}\left(\mathbb{R}^{n}\right)$, with $0 \leqq \hat{\psi} \leqq 1, \psi(\xi)=1$ for $|\xi| \leqq 1$ and $\hat{\psi}(\xi)=0$ for $|\xi| \geqq 2$, and define for any $\dot{J} \in \mathbb{Z}$

$$
\hat{\varphi}_{\mathbf{j}}(\xi)=\hat{\psi}\left(2^{-\mathbf{j}_{\xi}}\right)-\hat{\psi}\left(2^{-(\dot{j}+1)} \xi\right) .
$$

With any $u \in \mathcal{S}^{\prime}$ we can associate the sequence of functions $\left\{\varphi_{j} * u_{j} \equiv u_{j}\right\}$, $j \in \mathbb{Z}$. That sequence is actually defined for $u \in Z^{\prime}$ since $u_{j}=0$ for $u \in P$. For any $\rho \in \mathbb{R}$ and any $r$ and $s$ with $l \leqq r, s \leqq \infty$, we define the homogeneous Besov spaces

and the useful auxiliary spaces
$\dot{F}_{r s}^{\rho}=\left\{u \in Z^{\prime}:\left\|\left\{\sum_{\dot{J}}\left|2^{j \rho} \varphi_{j} * u\right|^{s}\right\}^{1 / s}\right\|_{r} \equiv\left\|u ; \dot{F}_{r s}^{p}\right\|<\infty\right\}$,
with obvious modifications if $s=\infty$. The factor $2^{\rho}$ in the sums mimicks a derivative of order $p$. All those spaces are Banach spaces and satisfy the following continuous embeddings

$$
\dot{\mathrm{F}}_{\mathrm{rs}}^{\rho} \subset \dot{\mathrm{B}}_{\mathrm{rs}}^{\rho}
$$

if $l \leqq r \leqq s \leqq \infty$, and

$$
\dot{S}_{r s}^{\rho} \subset \dot{F}_{r s}^{\rho}
$$

if $1 \leqq s \leqq r \leqq \infty$. This result allows one to compare the homogeneous Besov spaces to the homogeneous Sobolev spaces to be defined below. For any $\rho \in \mathbb{R}$, for any $u \in Z$ we define the operator $\omega^{\rho}$ by

$$
\widehat{\omega}^{\rho} u(\xi)=|\xi|^{\rho} \hat{u}(\xi)
$$

One sees by inspection that $\omega^{\rho}$ maps $Z$ into $Z$ and therefore by duality, $Z^{\prime}$ into $Z^{\prime}$. The space $X \equiv U \quad L^{r}$ is embedded in $\varphi^{\prime}$ and therefore in $1 \leqq r<\infty$ $Z^{\prime}$ so that, since any $u \in \mathbb{T}(X)$ is the image of only one element in $X$, the space $X$ is canonically embedded in $Z^{\prime}$. For any $\rho$, for any $r$ with $l \leqq r<\infty$, we define the homogeneous Sobolev space $\dot{H}_{r}^{0}$ by

$$
\dot{H}_{\mathrm{r}}^{\rho}=\omega^{-\rho_{\mathrm{L}}} \mathbf{r}
$$

Equipped with the norm $\left\|u ; \dot{H}_{r}^{\rho}\right\|=\left\|\omega^{\rho} u\right\|_{r}, \dot{H}_{r}^{\rho}$ becomes a Banach space. A non trivial argument shows that $\dot{F}_{r 2}^{\rho}=\dot{H}_{r}^{\rho}$ for $1<r<\infty$, so that the previous embeddings imply

$$
\dot{\mathrm{B}}_{\mathrm{r} 2}^{\rho} \subset \dot{H}_{\mathrm{r}}^{\rho} \subset \dot{\mathrm{B}}_{\mathrm{rr}}^{\rho}
$$

if $2 \leqq r<\infty$, and

$$
\dot{\dot{P}}_{\mathrm{rr}}^{\rho} \subset \dot{H}_{\mathrm{r}}^{\rho} \subset \dot{\mathrm{B}}_{\mathrm{r} 2}^{\rho}
$$

if $1<r \leqq 2$.
For our purposes the useful Sobolev inequalities in Besov spaces take the form of the following continuous inclusion

$$
\dot{\mathrm{B}}_{\mathrm{r} 2}^{\rho} \subset \dot{\mathrm{B}}_{\ell 2}^{\sigma}
$$

with $n / r-\rho=n / \ell-\sigma$, and $l \leqq r \leqq \ell \leqq \infty$. In particular

$$
\dot{\mathrm{B}}_{\mathrm{r} 2}^{\rho} \subset \mathrm{L}^{0}
$$

with $n / r-\rho=n / \ell, 2 \leqq \ell$ and $l \leqq r \leqq \ell<\infty$.
Finally we quote a result of great use in dealing with non linear problems in Besov spaces (see Lemma 3.2 of [2]) :

Lemma A.l. Let $f \in \mathscr{C}^{1}(\mathbb{C}, \mathbb{C})$ with $\left|f^{\prime}(z)\right| \leqq C|z|^{p-1}$ for some $p, l \leqq p$ $<\infty$. Let $0<\rho<1$, let $l \leqq r \leqq k \leqq \infty$ and $l / s=1 / r-1 / k$. then the following inequality holds

$$
\left\|f(u) ; \dot{B}_{r 2}^{\rho}\right\| \leqq c\left\|u ; \dot{S}_{k 2}^{\rho}\right\|\left\||u|^{p-1}\right\|_{s}
$$

for all $u$ such that $u \in L^{m}$ for some $m, l \leqq m<\infty$ and such that the norms in the right hand side are finite.

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## CONFORMAL INVARIANCE AND TIME DECAY

FOR NONLINEAR WAVE EQUATIONS
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## 1. INTRODUCTION AND GENERAL BACKGROUND

In this lecture, we describe some implications of the approximate conformal invariance of the nonlinear wave ( $N L W$ ) equation on the time decay of its solutions. The equation under study is

$$
\begin{equation*}
\square \varphi \equiv \ddot{\varphi}-\Delta \varphi=-f(\varphi) \tag{1.1}
\end{equation*}
$$

where $\varphi$ is a complex valued function defined in space-time $\mathbb{R}^{n+1}$, the upper dot denotes the time derivative, $\Delta$ is the Laplace operator in $\mathbb{R}^{n}$, and $f$ is a nonlinear complex valued function, a typical form of which is

$$
\begin{equation*}
f(\varphi)=\lambda \varphi|\varphi|^{p-1} \tag{1.2}
\end{equation*}
$$

with $1 \leq p<\infty$. The question under study is a special case of the more general question of determining the asymptotic behaviour in time of the solutions of an evolution equation for which the global Cauchy problem can be solved in a reasonably general fashion. It is known that an important tool for that purpose consists of a priori estimates derived from exact or approximate conservation laws associated with the equation. Another example which will be considered here for comparison, is that of the nonlinear Schrödinger (NLS) equation

$$
\begin{equation*}
i \dot{\varphi}=-\frac{1}{2} \Delta \varphi+f(\varphi) \tag{1.3}
\end{equation*}
$$

[^1]with a similar interaction $f$, together with its approximate pseudoconformal conservation law [4].

It is known (see [10] and references therein contained) that for initial data $\left(\varphi(0)=\varphi_{0}, \dot{\varphi}(0)=\psi_{0}\right)$ in the energy space $x_{e} \equiv H^{1} \oplus L^{2}$, the global Cauchy problem for the equation (1.1) has a unique solution $\varphi$ which is a continuous function of time with values in $X_{e}$ under assumptions on $f$ that reduce to $\lambda \geq 0$ and $p<1+4 /(n-2)$ in the special case (1.2). Furthermore $\varphi$ satisfies the conservation of the energy

$$
\begin{equation*}
E(\varphi(t), \quad \dot{\varphi}(t))=E\left(\varphi_{0}, \psi_{0}\right) \tag{1.4}
\end{equation*}
$$

for all $t \in \mathbb{R}$, where

$$
\begin{equation*}
\mathrm{E}(\varphi, \Psi)=\|\psi\|_{2}^{2}+\|\nabla \varphi\|_{2}^{2}+\int \mathrm{dx} V(\varphi(\mathrm{x})) \tag{1.5}
\end{equation*}
$$

$\|.\|_{2}$ denotes the norm in $L^{2} \equiv L^{2}\left(\mathbb{R}^{n}\right)$ and $V(\varphi)=2 \lambda(p+1)^{-1}|\varphi|^{p+1}$ in the special case (1.2) (See the general definition in Section 2 below). Similar results hold for the NLS equation (1.3).

Conservation laws for equations such as (1.1) and (1.3) formally result from the fact that the equation (i) is derived from a variational problem and (ii) is invariant under a one parameter continuous group of transformations, through the use of Noether's theorem (see [3] and references therein quoted). One is generally given a Lagrangian density $\mathscr{L}\left(\varphi,\left\{\partial_{\mu} \varphi\right\}\right)$ where $\left\{\partial_{\mu} \varphi ; 0 \leq \mu \leq n\right\} \equiv\left\{\partial_{0} \varphi \equiv \dot{\varphi},\left\{\partial_{j} \varphi ; 1 \leq j \leq n\right\}\right.$ $\equiv \nabla \varphi\}$ is the set of first order derivatives of $\varphi$, and the equation is the Euler - Lagrange equation associated with the action $A_{\Lambda}$ defined for any open subset $\Lambda \subset \mathbb{R}^{\text {n+1 }}$ with smooth boundary by

$$
A_{\Lambda}=\int \mathscr{L}\left(\varphi,\left\{\partial_{\mu} \varphi\right\}\right) d t d x
$$

With a one parameter transformation group of $\mathbb{R}^{n+1}$ under which the equation is invariant, there is associated an infinitesimal transformation of $\mathbb{R}^{n+1}$, namely a vector field in $\mathbb{R}^{n+1}\left\{\delta x^{\mu} ; 0 \leq \mu \leq n\right\}=$ $\left\{\delta \mathrm{x}^{0} \equiv \delta \mathrm{t},\left\{\delta \mathrm{x}^{j} ; 1 \leq j \leq \mathrm{n}\right\} \equiv \delta \mathrm{x}\right\}$ and an infinitesimal transformation of the field. $\delta \varphi$. If $\varphi$ satisfies the variational equation, the infinitesimal transformation of the action is easily seen to be

$$
\delta A_{\Lambda}=-\int_{\partial \Lambda} J^{\mu} d \sigma_{\mu}
$$

where $d \sigma_{\mu}$ is the area $n$-form in $\mathbb{R}^{n+1}$, $J^{\mu}$ is the current vector

$$
J^{\mu}=-\mathscr{L} \delta x^{\mu}-\left[\partial \mathscr{L} / \partial\left(\partial_{\mu} \varphi\right)\right] \delta \varphi
$$

we use the sumation convention of Relativity Theory, and the Minkowski metric $g_{\lambda \mu}$ used to raise and lower indices is defined by $g_{00}=-g_{i i}=1$, $g_{\lambda \mu}=0$ for $\lambda \neq \mu$. If the action is invariant, namely $\delta A_{\Lambda}=0$ for any $\Lambda$, then the current $J^{\mu}$ is conserved, namely $\partial_{\mu} J^{\mu}=0$. Taking for $\Lambda$ the region $s \leq x^{0} \leq t$ and assuming sufficient decay at infinity in space, one obtains the conservation law $Q(t)=Q(s)$ for all $s$ and $t$, where the charge $Q(t)$ is defined by

$$
Q(t)=\int_{x^{0}=t} J^{0} d x .
$$

If the invariance is only approximate, one obtains only the approximate conservation law

$$
\begin{equation*}
Q(t)-Q(s)=\int_{s}^{t} R(\tau) d \tau \tag{1.6}
\end{equation*}
$$

with

$$
R(\tau)=\int_{x^{0}=\tau} \partial_{\mu} J^{\mu} d x
$$

In the case of the NLW equation (1.1), the Lagrangian density is $\mathscr{L}\left(\varphi,\left\{\partial_{\mu} \varphi\right\}\right)=|\dot{\varphi}|^{2}-|\nabla \varphi|^{2}-v(\varphi) \equiv \partial^{\mu} \bar{\varphi}_{\mu} \varphi-v(\varphi)$, where $v$ is the same potential function as in the energy (1.5). The relevant transformations of space-time are those of the conformal group, generated by the Poincaré transformations, space-time dilations, and the transformations obtained therefrom by conjugation under the external automorphism $x^{\mu} \rightarrow x^{\mu}\left(x_{\lambda} x^{\lambda}\right)^{-1}$. Of special interest are the transformations thereby obtained from space-time translations. The corresponding infinitesimal transformation of space-time is given by

$$
\delta x^{\mu}=-2 a^{\lambda} x_{\lambda} x^{\mu}+x^{\lambda} x_{\lambda} a^{\mu}
$$

where $a=\left\{a^{\mu}\right\}$ is a space-time vector, and the transformation of the field $\varphi$ is given by

$$
\delta \varphi=-\delta x^{\lambda} \partial_{\lambda} \varphi+(n-1) a^{\lambda} x_{\lambda} \varphi .
$$

Taking in particular for a the vector $\{1,0, \ldots, 0\}$, one obtains the relevant conservation law in the form (1.6) with

$$
\begin{align*}
& Q(t, \varphi, \dot{\varphi})=Q_{0}(t, \varphi, \dot{\varphi})+Q_{1}(t, \varphi),  \tag{1.7}\\
& Q_{0}(t, \varphi, \psi)=\|r \psi\|_{2}^{2}+t^{2}\|\psi\|_{2}^{2}+\|r \nabla \varphi\|_{2}^{2}+t^{2}\|\nabla \varphi\|_{2}^{2} \\
& \quad+2 t \operatorname{Re}\left\langle\psi, 2 x \cdot \nabla \varphi+(n-1) \varphi>-(n-1)\|\varphi\|_{2}^{2},\right.  \tag{1.8}\\
& Q_{1}(t, \varphi)=\int d x\left(t^{2}+r^{2}\right) V(\varphi),  \tag{1.9}\\
& R(t, \varphi)=2 t \int d x W(\varphi)  \tag{1.10}\\
& W(\varphi)=(n+1) V(\varphi)-(n-1) \operatorname{Re} \bar{\varphi} f(\varphi), \tag{1.11}
\end{align*}
$$

where $r=|x| ;<., .>$ denotes the scalar product in $L^{2}$. $Q$ as given by (1.7)-(1.9) will be called the conformal charge, and $Q_{0}$ and $Q_{1}$ will be called the kinetic and potential parts thereof. By an elementary computation, $Q_{0}$ can be rewritten in the form

$$
\begin{equation*}
Q_{0}(t, \varphi, \psi)=\|t \nabla \varphi+x \psi\|_{2}^{2}+\|L \varphi\|_{2}^{2}+\|x \cdot \nabla \varphi+t \psi+(n-1) \varphi\|_{2}^{2} \tag{1.12}
\end{equation*}
$$

where $L=x \times \nabla$ is the angular momentum operator, which shows in particular that $Q_{0}$ is non negative.

The fact that the conformal conservation law implies some time decay of the solutions appears immediatly in the case of the single power interaction (1.2), where $V(\varphi)=2 \lambda(p+1)^{-1}|\varphi|^{p+1}$ and

$$
\begin{equation*}
W(\varphi)=[n+1-(p+1)(n-1) / 2] V(\varphi) . \tag{1.13}
\end{equation*}
$$

In fact for $\dot{\lambda} \geq 0$ and $p+1 \geq \ell_{1} \equiv 2(n+1) /(n-1)$, or equivalently $p \geq 1+$ $4 /(n-1)$, one has $W(\varphi) \leq 0, t R(t, \varphi) \leq 0$ and therefore $Q(t) \leq Q(0)$ for all $t \in \mathbb{R}$. Since $Q_{0} \geq 0$, one obtains by using only $Q_{1}$

$$
\begin{equation*}
\|\varphi(t)\|_{p+1} \leq c t^{-2 /(p+1)} \tag{1.14}
\end{equation*}
$$

where \|. $\|_{l}$ denotes the norm in $L^{l} \equiv L^{\ell}\left(\mathbb{R}^{n}\right)$. That decay property was used extensively in previous works on the NLW equation [24].

In order to assess the strength of the decay property (1.14) it is useful to compare it with the decay available for solutions of the free wave equation $\square \varphi=0$, for which one obtains at best

$$
\begin{equation*}
\|\varphi(t)\|_{\ell} \leq C(l+|t|)^{-\gamma(\ell)} \tag{1.15}
\end{equation*}
$$

with

$$
\begin{equation*}
\gamma(\ell)=(n-1)(1 / 2-1 / \ell) . \tag{1.16}
\end{equation*}
$$

That the decay (1.15) is optimal is easily seen in the special case of the dimension $n=3$ where the solution of $\square \varphi=0$ with initial data $\left(\varphi_{0}=0, \psi_{0}\right)$ is given by

$$
\varphi(t)=(4 \pi t)^{-1} \delta(r-t) * \psi_{0}
$$

where * denotes the convolution in $\mathbb{R}^{n}$, and can be explicitly computed for $\psi_{0}=X(r \leq a)$, the characteristic function of the ball of radius $a$, to be

$$
\varphi(t)=(4 r)^{-1}\left[a^{2}-(t-r)^{2}\right]_{+}
$$

for $t \geq a$. More generally solutions of $\square \varphi=0$ with the decay (1.15) can be constructed with the help of the following lemma. There we use the notation $\omega=\sqrt{-\Delta}$, and in addition to $\gamma(\ell)$ defined by (1.16), we also use $\alpha(\ell)$ and $\delta(\ell)$ defined by

$$
\begin{equation*}
\alpha(\ell) \equiv \mathrm{n}^{-1} \delta(\ell) \equiv 1 / 2-1 / \ell . \tag{1.17}
\end{equation*}
$$

Lemma 1.1 $[17,20,26]$. Let $n \geq 2$, let $\ell$ and $s$ satisfy

$$
\begin{equation*}
1+\delta(s) \leq \delta(\ell) \leq \operatorname{Min}(1+\alpha(s), \mathrm{n}(1+\delta(s))) \tag{1.18}
\end{equation*}
$$

and $2<\ell<\infty$ if $n=2$. Then the following estimate holds

$$
\begin{equation*}
\left\|\omega^{-1} e^{ \pm i \omega t} \varphi\right\|_{\ell} \leq c|t|^{1-\delta(\ell)+\delta(s)}\|\varphi\|_{s} \tag{1.19}
\end{equation*}
$$

for all $t \neq 0$ and all $\varphi \in L^{s}$.

In particular, the optimal decay (1.15) for solutions

$$
\begin{equation*}
\varphi(t)=K(t) \psi_{0}+\dot{K}(t) \varphi_{0} \tag{1.20}
\end{equation*}
$$

of $\square \varphi=0$, where $K(t)=\omega^{-1}$ sinwt and $\dot{K}(t)=c o s \omega t$, can be obtained for suitable initial data $\left(\varphi_{0}, \psi_{0}\right)$ by using Lemma $1 . l$ in the borderline case $\alpha(\ell)=1+\delta(s)$.

Since $\gamma(p+1)=2 /(p+1)$ for $p+1=l_{1}$, the decay (1.14) obtained from the conformal conservation law coincides with the optimal decay (l.15) in that limiting case, while it is generally weaker in the allowed range $p+1 \geq \ell_{1}$. That result can then be improved (actually has been improved in previous works [24, 29]) by substituting the decay (1.14) in the integral equation associated with the equation (1.1) and performing various estimates thereof. It is however a natural question to ask whether stronger results can be obtained more directly by a more efficient application of the conformal conservation law, using in particular the kinetic part $Q_{0}$ of the conformal charge. For that purpose, it is useful to compare the case of the NLW equation with the simpler case of the NLS equation (1.3). The solutions of that equation satisfy the pseudo-conformal conservation law which is similar to the conformal conservation law of the NLW equation [4]. That law takes again the form (1.6) where however (the subscript $S$ stands for Schrödinger) $Q_{S}=Q_{O S}$ $+Q_{1 S}$ with

$$
\begin{align*}
Q_{0 S}(t, \varphi) & =\frac{1}{2}\|(x+i t \nabla) \varphi\|_{2}^{2}  \tag{1.8}\\
Q_{1 S}(t, \varphi) & =t^{2} \int d x V(\varphi)  \tag{1.9}\\
R_{S}(t, \varphi) & =t \int d x W_{S}(\varphi)  \tag{1.10}\\
W_{S}(\varphi) & =(n+2) V(\varphi)-n \operatorname{Re} \bar{\varphi} f(\varphi) \tag{1.11}
\end{align*}
$$

In the case of a single power interaction (l.2), one obtains

$$
\begin{equation*}
W_{S}(\varphi)=[n+2-(p+1) n / 2] V(\varphi) \tag{1.13}
\end{equation*}
$$

so that for $\lambda \geq 0$ and $p+1 \geq 2(n+2) / n$ or equivalently $p \geq 1+4 / n$, one has as before $W_{S}(\varphi) \leq 0, t R_{S}(t, \varphi) \leq 0, Q_{S}(t) \leq Q_{S}(0)$ for all $t$ in $\mathbb{R}$, and the solution $\varphi$ again satisfies the decay (1.14). The optimal time decay available for solutions of the free Schrödinger equation i $\dot{\varphi}=$ $(-1 / 2) \Delta \varphi$ is known to be

$$
\begin{equation*}
\|\varphi(t)\|_{\ell} \leq c(1+|t|)^{-\delta(l)} \tag{1.15}
\end{equation*}
$$

where $\delta(\ell)$ is defined by (l.17). The decay (1.14) is again identical with the optimal one in the limiting case $p+1=2(n+2) / n$, and weaker in the allowed range $p+1 \geq 2(n+2) / n$. In the NLS case however, one can obtain easily a better result by using directly the kinetic part $Q_{0 S}$
of the pseudo-conformal charge. For that purpose we introduce the free Schrödinger group $U(t)=\exp (i(t / 2) \Delta)$. For each $t \neq 0, U(t)$ can be represented as the operator of convolution with the function $(2 \pi i t)^{-n / 2}$ exp[ $\left.i x^{2} /(2 t)\right]$. Furthermore, the following identity holds

$$
\begin{equation*}
J \equiv \mathrm{x}+i t \nabla=\mathrm{U}(\mathrm{t}) \mathrm{x} \mathrm{U}(-t) . \tag{1.21}
\end{equation*}
$$

One can then prove the following estimate.
Lemma 1.2 [4]. Let $2 \leq \ell \leq 2^{*} \equiv 2 n /(n-2), \ell<\infty$ if $n=2$. Then

$$
\begin{equation*}
\|\varphi\|_{\ell} \leq c|t|^{-\delta(\ell)}\|x U(-t) \varphi\|_{2}^{\delta(\ell)}\|\varphi\|_{2}^{1-\delta(\ell)} \tag{1.22}
\end{equation*}
$$

for all $t \neq 0$ and all $\varphi$ for which the right hand side is finite. (Here c is a Sobolev constant).

Proof By an homogeneity argument, it is sufficient to prove the result for $t=-1$. In that case $U(1)=\theta S \mathcal{F}$ where $\mathcal{F}$ is the Fourier transform, $\theta$ is a constant phase factor, and $S$ is the operator of multiplication by $\exp \left(i x^{2} / 2\right)$. Then

$$
\|\mathrm{xU(1)} \varphi\|_{2}=\|\mathrm{x} \mathcal{G} \varphi\|_{2}=\|\nabla S \varphi\|_{2}
$$

by the Plancherel theorem, while

$$
\|\varphi\|_{\ell}=\|s \varphi\|_{\ell} \leq \mathrm{c} \quad\|\nabla \mathrm{~S} \varphi\|_{2}^{\delta(\ell)}\|\mathrm{S} \varphi\|_{2}^{1-\delta(\ell)}
$$

by a standard Sobolev inequality. This proves the lemma.
Q.E.D.

Applying Lemma 1.2 with $\varphi=\varphi(t)$ to a solution of the NLS equation (1.3) yields the optimal decay (1.15) for $2 \leq \ell \leq 2^{*}$ whenever $Q_{0 S}$ is uniformly bounded in time, and in particular in the case of a single power interaction (1.2) with $\lambda \geq 0$ and $p+1 \geq 2(n+2) / n$ considered previously.

The first main result reported in this lecture is the analogue of Lemma 1.2 in the case of the NLW equation, yielding in particular an estimate of the $L^{\ell}$ norm of its solutions for $2 \leq \ell \leq 2^{*}$ with the optimal time decay (1.15) in terms of the kinetic part $Q_{0}$ of the conformal charge (See Proposition 3.2 below and [8] for more details). Actually we shall present a stronger result, since the basic estimate involves space-time weighted $L^{2}$ - norms and thereby yields an additional decay in
space. As in the NLS case, that estimate will emerge as a variant of a Sobolev inequality, but of a much more complicated type than the elementary estimate (1.22).

The previous considerations and results cover the "easy" case where the conformal conservation law immediately implies the boundedness of the conformal charge, corresponding to $p+1 \geq l_{1}$ in the special case (1.2). In order to guess what time decay should be expected in the more difficult case $p+1<\ell_{1}$, it is useful to consider again the situation for the NLS equation. The problem of the time decay of solutions can be regarded as part of the general theory of Scattering for that equation $[4,7,27,28]$. Another part of that theory is the construction of dispersive solutions, namely of solutions that behave asymptotically in time as solutions of the corresponding free equation, obtained by dropping the interaction term f. Dispersive solutions can be constructed by solving the Cauchy problem with infinite initial time by a contraction method, in suitable functional spaces that include some time decay in their definition. The method can be implemented under assumptions on $f$ which reduce to a lower bound on $p$ in the special case (1.2). That lower bound depends on the choice of the time decay that appears in the relevant functional space. If that decay is chosen to be as strong as the optimal decay (1.15) $S^{\prime}$, the lower bound $p_{S}(n)$ on $p$ in the NLS case turns out to be the (positive) root of the equation

$$
\begin{equation*}
p \delta(p+1)=1 \tag{1.23}
\end{equation*}
$$

or equivalently

$$
\begin{equation*}
n p^{2}-(n+2) p-2=0 \tag{1.24}
\end{equation*}
$$

namely $p_{S}(2)=1+\sqrt{2}, p_{S}(3)=2$, etc. Remarkably enough, it can be shown that the kinetic part of the conformal charge still remains bounded in time and therefore that the optimal decay still holds for all $p>p_{S}(n)[14,27,28]$. We state the result in the special case (1.2) only, in a somewhat loose way.

Proposition 1.1 $[14,27,28]$. Let $p_{S}(n)<p \leq 2(n+2) / n$, let $\varphi_{0} \in H^{1}$, with $\times \varphi_{0} \in \mathrm{~L}^{2}$, and let $\varphi$ be the (unique $H^{1}$-valued) solution of the NLS equation (1.3) with $f$ given by (1.2) and $\lambda \geq 0$, with initial data $\varphi_{0}$ at time zero. Then
(1) $Q_{0 S}(t, \varphi(t))$ is uniformly bounded in time.
(2) $\varphi$ satisfies the decay estimate (1.15) ${ }_{S}$ for $2 \leq \ell \leq 2^{*}$ ( $\ell<\infty$ if $n=2$ ).

Sketch of proof. It suffices to prove part (1), from which part (2) follows by Lemma l.2, The proof proceeds in two steps.
(1) By reestimating $W_{S}$ in terms of $V$ in the pseudo-conformal conservation law, one obtains a differential inequality for the pseudo-conformal charge which by integration yields (among other) the decay (1.15)s for the special value $\ell=p+1$.
(2) One has to estimate $J \varphi$ (see (1.21)) in $L^{2}$ uniformly in time. Now $J$ is easily seen to commute with i $d / d t+(1 / 2) \Delta$, so that $J \varphi$ satisfies the differential equation

$$
\begin{equation*}
i \frac{d}{d t} J \varphi=-\frac{1}{2} \Delta J \varphi+J f(\varphi) \tag{1.25}
\end{equation*}
$$

and therefore the associated integral equation

$$
\begin{equation*}
J \varphi(t)=U(t) \times \varphi_{0}-i \int_{0}^{t} d \tau U(t-\tau) J £(\varphi(\tau)) . \tag{1.26}
\end{equation*}
$$

Let now $S \equiv S(t) \equiv \exp \left(i x^{2} / 2 t\right)$. Then $J$ also satisfies the relation

$$
\begin{equation*}
J=S \text { it } \nabla S^{-1} \tag{1.27}
\end{equation*}
$$

so that

$$
\begin{aligned}
J f(\varphi) & =S \text { it } \nabla S^{-1} f(\varphi)=S \text { it } \nabla f\left(S^{-1} \varphi\right) \\
& =\operatorname{sit}\left(\nabla S^{-1} \varphi\right) \frac{\partial f}{\partial \varphi}\left(S^{-1} \varphi\right)+S i t\left(\nabla \overline{S^{-1} \varphi}\right) \frac{\partial f}{\partial \bar{\varphi}}\left(S^{-1} \varphi\right) \\
& =(J \varphi) \frac{\partial f}{\partial \varphi}-(\overline{J \varphi}) \frac{\partial f}{\partial \bar{\varphi}}
\end{aligned}
$$

Therefore (1.26) is a linear equation for $J \varphi$ regarded as an independent function. One can then use that equation to estimate $J \varphi$ in $L^{2}$ uniformly in time by using known estimates on the operator

$$
h \rightarrow \int_{0}^{t} d \tau U(t-\tau) h(\tau)
$$

[16] together with the estimate (1.15)s with $\ell=p+1$ obtained in step 1 in order to control the factors $\partial f / \partial \varphi$ and $\partial f / \partial \bar{\varphi}$ coming from Jf $(\varphi)$. That step turns out of work precisely under the condition $p>p_{S}(n)$.
Q.E.D.

The preceding line of argument can be extended to some extent to the NLW equation in the difficult case $p+1 \leq \ell_{1}$, but one encounters several difficulties which somehow restrict the scope of the results, as we now explain. The optimal decay rate (1.15) in the NLW case has $\delta(\ell)$ replaced by $\gamma(\ell)$ as compared with the NLS case (or for that matter, with the case of the massive nonlinear Klein-Gordon (NLKG) equation $\square \varphi+\varphi+f(\varphi)=0)$. The shift from $n$ to $n-1$ is best understood by noticing that dispersion in the massless case $\square \varphi=0$ occurs (at least in odd dimensions by the Huygens principle) in (n-l)-dimensional submanifolds instead of $n$ dimensional space in the massive case and in the Schrödinger case. Correspondingly, one expects the contraction argument leading to the proof of existence of dispersive solutions of the NLW equation (1.1) with interaction (l.2) to work for $p>p_{0}(n) \equiv$ $p_{S}(n-1)\left(\right.$ i.e $p_{0}(3)=1+\sqrt{2}, p_{0}(4)=2, \ldots$ ). That lower bound is known to be optimal, in view of the existing blow up results for attractive interactions and small initial data in the opposite case $p \leq p_{0}(n)$ $[11,15,22]$. A natural tool to be used in order to implement the contraction argument is the estimate of Lemma l.l., which actually produces the optimal decay (1.15) in the borderline case $\alpha(\ell)=1+\delta(s)$. Unfortunately, the use of that lemma produces the expected result only under the stronger condition $p>p_{1}(n)$ where $p_{1}(n)>p_{0}(n)$ is the (larger) root of the equation

$$
\begin{equation*}
n(n-1) p^{2}-p\left(n^{2}+3 n-2\right)+2=0 \tag{1.28}
\end{equation*}
$$

$[6,18,19,23]$. Extending the contraction argument from $p_{1}(n)$ down to $p_{0}(n)$ is a highly non trivial matter and has been done so far only in dimensions 2 [.12] and 3 [15, 2l] for special classes of solutions, by using in an essential way the positivity properties of the propagator for the free equation $\square \varphi=0$, and by using space-time weighted norms instead of simply $L^{\ell}$-norms.

The general scheme leading to the analogue of Proposition 1.1 can be implemented in the NLW case; in particular the expression (1.12) for $Q_{0}$ is an adequate substitute for the expression $Q_{0 S}=(1 / 2)\|J \varphi\|_{2}^{2}$. Unfortunately the NLW analogue of the first step in the proof of that proposition does not yield (for $\ell=p+1$ ) the optimal decay (1.15) but only the weaker decay

$$
\begin{equation*}
\|\varphi(t)\|_{\ell} \leq c|t|^{1-2 \delta(\ell)} \tag{1.29}
\end{equation*}
$$

A simple power counting argument then shows that the remaining part of the proof can be expected to hold only under an assumption on $p$ which allows for the contraction argument mentioned previously to work when only the decay (1.29) is used in the definition of the relevant functional space. That assumption turns out to be $p>p_{2}(n)$ where $p_{2}(n)$ $>p_{1}(n)$ is the (positive) root of the equation [25]

$$
\begin{equation*}
p(2 \delta(p+1)-1)=1 \tag{1.30}
\end{equation*}
$$

or equivalently

$$
(n-1) p^{2}-(n+2) p-1=0
$$

Thus, although one may conjecture that the analogue of Proposition l.l holds in the NLW case for $p>p_{0}(n)$, the available method of proof is restricted to $p>p_{2}(n)$.

A final difference between the NLS and NLW cases is the technical fact that the estimates on the integral equation are substantially more difficult in the latter case than in the former. As a consequence the expected results have been proved only for certain dimensions, namely $n=3$ and $n=4$ [9]. The case of higher dimension is technically more complicated and has not been worked out in detail. The case of dimension $n=2$ can be analyzed rather completely down to $p_{2}(2)$, but is plagued with special difficulties in the form of logarithmic factors [8]. The case $n=1$ does not seem to be interesting in the present context.

The second main result reported in this lecture is the analogue of Proposition $1 . l$ in the case of the NLW equation, in the case of space dimension $n=2,3$ and 4 and under assumptions on $f$ that reduce to $p_{2}(n)<p<1+4 /(n-2)$ in the special case (1.2), according to the preceding discussion.

The remaining part of this lecture is organized as follows. In Section 2 we state the conformal conservation law for the NLW equation with the relevant functional analytic details, including in particular the choice of the relevant space of initial data and the natural assumptions on the interaction $f$. In section 3 , we state the basic apriori estimates in terms of the kinetic part of the conformal charge, extending Lemma 1.2 to the NLW case, and we give a brief sketch of their proof. Finally in Section 4, we derive therefrom the resulting decay estimates of the solutions of the NLW equation, first for general $n \geq$ 2 in the easy case corresponding to $p+1 \geq \ell_{1}$, and then for $n=2,3$, 4 in the harder case $p_{2}(n)<p \leq \ell_{1}-1$. We refer the reader to [8]
for a complete treatment of the special case $n=2$ and to $[8,9]$ for the details of the proofs.

We conclude this section with the remark that, in the framework of Scattering Theory, the results of Sections 3 and 4 are the essential ingredients of the proof of asymptotic completeness (for arbitrarily large initial data).

## 2. THE CONFORMAL CONSERVATION LAW.

In order to convert the formal derivation of the conformal conservation law sketched in Section $l$ into an actual proof, it is convenient to apply the following general method. One first regularizes the equation by introducing suitable cut-offs. The solutions of the regularized equation have sufficient smoothness and decay at infinity in space to allow for the proof of a regularized version of the conservation law by the same computations as used in the preceding heuristic derivation. One finally removes the cut-offs by a limiting procedure.

The regularization uses a local cut-off $h$ and a space cut-off $g$ at large distances. The cut-off $h$ is taken as a non negative even function in $\mathscr{C}_{0}^{\infty}\left(\mathbb{R}^{n}\right)$, such that $\|h\|_{1}=1$, and the cut-off $g$ as a function in $\mathscr{U}^{1}\left(\mathbb{R}^{n}\right)$ with compact support and such that $0 \leq g \leq 1$ and $g=1$ in some region around the origin. Because of the finite propagation speed for the equation (1.1), the space cut-off can be introduced either in the initial data or in the interaction. We choose the second possibility because of the intrinsic interest of localized interactions. We replace the equation (1.1) by the regularized equation

$$
\begin{equation*}
\square \varphi+h * g f(h * \varphi)=0 \tag{2.1}
\end{equation*}
$$

and consider the Cauchy problem for (2.1) with regularized initial data $(\varphi(0), \dot{\varphi}(0))=\left(h * \varphi_{0}, h * \psi_{0}\right)$.

The limiting procedure consists in letting $h$ tend to a $\delta$ function and $g$ tend to 1 in the following sense : we choose fixed $h_{1}$ and $g_{1}$ as described above; for any positive integer $j$, we define $h_{j}(x)=j^{n_{h}}{ }_{1}(j x)$ and $g_{j}(x)=g(x / j)$, we take $h=h_{j}, g=g_{j}$ and we let $j$ tend to infinity. The solutions of the regularized equation (2.l) tend to solutions of (1.1) in a sense which makes it possible to take the limit in the conservation law, provided the interaction $f$ and the initial data ( $\varphi_{0}$, $\psi_{0}$ ) satisfy suitable assumptions. These assumptions basically ensure that (i) the Cauchy problem for (1.1) can be solved in a unique way, and (ii) the conservation law has a meaning.

The natural condition on $(\varphi, \psi)$ for the energy $E(\varphi, \psi)$ and the kinetic part of the conformal charge $Q_{0}(t, \varphi, \psi)$ to be defined is that $(\varphi, \psi) \in \Sigma \equiv \Sigma_{1} \oplus \Sigma_{0}$, where

$$
\begin{equation*}
\Sigma_{1}=\left\{\varphi \in \mathrm{L}^{2}: \nabla \varphi, r \nabla \varphi \in \mathrm{~L}^{2}\right\} \tag{2.2}
\end{equation*}
$$

and

$$
\begin{equation*}
\Sigma_{0}=\left\{\psi \in \mathrm{L}^{2}: r \psi \in \mathrm{~L}^{2}\right\} \tag{2.3}
\end{equation*}
$$

The spaces $\Sigma_{1}$ and $\Sigma_{0}$ are Hilbert spaces with norms
$\left\|\varphi ; \Sigma_{1}\right\|^{2}=\|\varphi\|_{2}^{2}+\|\nabla \varphi\|_{2}^{2}+\|r \nabla \varphi\|_{2}^{2}$
$\left\|\psi ; \Sigma_{0}\right\|^{2}=\|\psi\|_{2}^{2}+\|r \psi\|_{2}^{2} \quad$.
We shall consider only initial data $\left(\varphi_{0}, \psi_{0}\right) \in \Sigma$.
The assumptions on the interaction $f$ consist of those required to solve the global Cauchy problem with uniqueness (namely (Al) and (A2, a, b) below) supplemented with an assumption which prevents $Q_{1}(t, \varphi)$ to tend to $-\infty$ while $Q_{0}(t, \varphi, \psi)$ tends to $+\infty$ for fixed $Q(t, \varphi, \psi)$ (namely (A2c)). They can be stated as follows.
(A1). $£ \in \mathscr{C}^{1}(\mathbb{C}, \mathbb{C})$ and $f(0)=0$. If $n \geq 2$, $f$ satisfies the estimate

$$
\begin{equation*}
\left|f^{\prime}(z)\right| \equiv \operatorname{Max}(|\partial f / \partial z|,|\partial f / \partial \bar{z}|) \leq C\left(|z|^{p_{1}^{-1}}+|z|^{p_{2}-1}\right) \tag{2.4}
\end{equation*}
$$

for some $p_{1}, p_{2}$ with $l \leq p_{1} \leq p_{2}<1+4 /(n-2)$, and all $z \in \mathbb{C}$.
(A2). (a) There exists a function $V \in \mathscr{Q}^{1}(\mathbb{C}, \mathbb{R})$ such that $V(0)=0$, $V(z)=V(|z|)$ for all $z \in \mathbb{C}$, and $f(z)=\partial V / \partial \bar{z}$.
(b) V satisfies the estimate
$V(R) \geq-C R^{2}$
for some $\mathrm{C} \geq 0$ and all $\mathrm{R} \geq 0$.
(c) V satisfies the estimate

$$
\begin{equation*}
V(R) \geq-C R^{2+4 / n} \tag{2.6}
\end{equation*}
$$

```
for some C \geq 0 and all R \geq0.
```

We can now state the conservation law, first with the cut-off $g$ still included, and finally with both cut-offs removed.

Proposition 2.1 Let $f$ satisfy (A1) and (A2 a, b), let ( $\varphi_{0}, \psi_{0}$ ) $\Sigma$, let $t_{0} \in \mathbb{R}$ and let $\varphi$ be the (unique $H^{1}$-valued) solution of (l.1) with initial data $\left(\varphi_{0}, \psi_{0}\right)$ at time $t_{0}$. Then $(\varphi, \dot{\varphi}) \in \mathscr{C}(\mathbb{R}, \Sigma)$ and for all $s, t \in \mathbb{R}$, the following identity holds:

$$
\begin{align*}
& Q_{0}(t, \varphi(t), \dot{\varphi}(t))+\int d x\left(t^{2}+r^{2}\right) g V(\varphi(t))= \\
& Q_{0}(s, \varphi(s), \dot{\varphi}(s))+\int d x\left(s^{2}+r^{2}\right) g V(\varphi(s)) \\
& +\int_{s}^{t} 2 \tau d \tau \int d x\{g W(\varphi(\tau))+(x . \nabla g) V(\varphi(\tau))\} \tag{2.7}
\end{align*}
$$

Proposition 2.2 Let $f$ satisfy (Al) and (A2 $a, b, c)$, let $\left(\varphi_{0}, \psi_{0}\right) \in \Sigma$. Assume that

$$
\int \mathrm{dx} \mathrm{r}^{2}\left|\mathrm{v}\left(\varphi_{0}\right)\right|<\infty
$$

Let $t_{0} \in \mathbb{R}$ and let $\varphi$ be the (unique $H^{1}$-valued) solution of (l.l) with initial data $\left(\varphi_{0}, \psi_{0}\right)$ at time $t_{0}$. Then $(\varphi, \dot{\varphi}) \in \mathscr{C}(\mathbb{R}, \Sigma)$, the integral $\int d x r^{2} V(\varphi)$ is absolutely convergent for each $t \in \mathbb{R}$ and continuous with respect to $t$, and $\varphi$ satisfies the conformal conservation law (1.6) - (1.ll) for all $s, t \in \mathbb{R}$, i.e. (2.7) with $g=1$.

We refer to [8], Section 2 for the details of the proofs.

## 3. CONFORMAL ESTIMATES

In this section, we derive some a priori estimates involving the kinetic part of the conformal charge, thereby providing an analogue of Lemma 1.2 in the case of the wave equation. The starting point is the expression (1.l2) for $Q_{0}(t, \varphi, \psi)$, where we regard $t$ as a real positive parameter and $\varphi, \psi$ as two independent functions in $\Sigma_{1}$ and $\Sigma_{0}$ respective -ly. We use two different methods.

The first method exploits Lemma 1.l. For that purpose, we first rewrite $Q_{0}$ in the form

$$
\begin{equation*}
Q_{0}(t, \varphi, \psi)=\|t \nabla \varphi+x \psi\|_{2}^{2}+\left\|x \omega \varphi-\omega^{-1} t \nabla \psi\right\|_{2}^{2} \tag{3.1}
\end{equation*}
$$

with $\omega=\sqrt{-\Delta}$, after an elementary computation involving the commutator of x with $\omega$. We then recombine $\varphi$ and $\psi$ into two new functions $\varphi_{ \pm}$ which would be the positive and negative frequency parts of $\varphi$ if $\varphi$ were a solution of $\square \varphi=0$ and $\psi$ its time derivative, namely

$$
\begin{equation*}
\varphi_{ \pm}=\left(\varphi \pm i \omega^{-1} \psi\right) / 2 \tag{3.2}
\end{equation*}
$$

One can then rewrite $Q_{0}$ as

$$
\begin{align*}
Q_{0}(t, \varphi, \psi) & =2 \sum_{ \pm}\left\|(x \omega \pm i t \nabla) \varphi_{ \pm}\right\|_{2}^{2}  \tag{3.3}\\
& =2 \sum\left\|x \omega \exp ( \pm i \omega t) \varphi_{ \pm}\right\|_{2}^{2} \tag{3.4}
\end{align*}
$$

where we have used the identity

$$
\begin{equation*}
\exp (\mp i \omega t) x \exp ( \pm i \omega t)=x \pm i t \omega^{-1} \nabla \tag{3.5}
\end{equation*}
$$

We can then prove the following estimate.

Proposition 3.1 Let $n \geq 2,2<\ell \leq \ell_{1}$, and $(\varphi, \psi) \in \Sigma$. Then $\varphi$ and $\omega^{-l} \psi$ belong to $L^{\ell}$ and satisfy the estimates

$$
\begin{equation*}
\|\varphi\|_{\ell},\left\|\omega^{-1} \psi\right\|_{\ell} \leq C t^{-\gamma(\ell)} E_{0}(\varphi, \psi)^{\alpha(\ell) / 2} Q_{0}(t, \varphi, \psi)^{(1-\alpha(\ell)) / 2} \tag{3.6}
\end{equation*}
$$

where $\mathrm{E}_{0}(\varphi, \psi)=\|\psi\|_{2}^{2}+\|\nabla \varphi\|_{2}^{2}$.

Proof. We express $\varphi$ and $\psi$ in terms of $\varphi_{ \pm}$by using (3.2). We then estimate $\varphi_{ \pm}$in $L^{\ell}$ by Lemma l.l with $\alpha(\ell)=1+\delta(s)$ and obtain

$$
\begin{align*}
\left\|\varphi_{ \pm}\right\|_{\ell} & \leq c t^{-\gamma(\ell)}\left\|\omega e^{ \pm i \omega t} \varphi_{ \pm}\right\|_{S} \\
& \leq c t^{-\gamma(\ell)}\left\|\omega e^{ \pm i \omega t} \quad \varphi_{ \pm}\right\|_{2}^{\alpha(\ell)} \quad\left\|x \omega e^{ \pm i \omega t} \quad \varphi_{ \pm}\right\|_{2}^{1-\alpha(\ell)} \tag{3.7}
\end{align*}
$$

where the second inequality is obtained by applying the Holder inequality in the form

$$
\|x\|_{s} \leq\left\|\left(a^{2}+x^{2}\right)^{-1 / 2}\right\|_{m}\left\|\left(a^{2}+x^{2}\right)^{1 / 2} x\right\|_{2}
$$

with $1 / s=1 / m+1 / 2$ and optimizing with respect to a. The result then follows from (3.4), (3.7) and the fact that

$$
E_{0}(\varphi, \psi)=2 \underset{ \pm}{\Sigma}\left\|\omega \varphi_{ \pm}\right\|_{2}^{2}
$$

Q.E.D.

Although Proposition 3.1 is a result of the type we are looking for, it has some shortcomings. The limiting case $\ell=2$ is excluded by the use of the Hölder inequality, whereas (3.6) for $\ell=2$ and $n \geq 3$ is simply the Hardy inequality. Furthermore the upper limit $\ell \leq \ell{ }_{1}$ is not optimal, as one expects the result to hold for all $\ell \leq 2^{*}(\ell<\infty$ for $n=2$ ). The upper limit $\ell=2^{*}$ for $n \geq 3$ is sharp however, as shown by the following argument : for $\varphi=\varphi_{+},(3.7)$ takes the form

$$
\begin{aligned}
\|\varphi\|_{\ell} & \leq c t^{-\gamma(\ell)}\|\omega \varphi\|_{2}^{\alpha(\ell)}\|(x \omega+i t \nabla) \varphi\|_{2}^{1-\alpha(\ell)} \\
& =c t^{1-\delta(\ell)}\|\omega \varphi\|_{2}^{\alpha(\ell)}\left\|\left(x t^{-1} \omega+i \nabla\right) \varphi\right\|_{2}^{1-\alpha(\ell)}
\end{aligned}
$$

which for smooth and rapidly decaying $\varphi$ implies $\|\varphi\|_{\ell}=0$ and therefore $\varphi=0$ if $\delta(\ell)>1$, or equivalently $\ell>2^{*}$, by letting $t$ tend to infinity.

The second method will remedy the previous defects. We restrict our attention to the case $n \geq 3$ and refer to [8] for a detailed treatment of the case $n=2$ which has additional complications in the form of logarithmic factors in the estimates. Since we want to estimate $\varphi$ in terms of $Q_{0}$ which contains another function $\psi$ independent of $\varphi$, we lose nothing by first minimizing $Q_{0}$ as a function of $\psi$. We introduce an auxiliary radial function $h \in \mathscr{C}((0, \infty) ; \mathbb{R})$ and we use the notation $\hat{\mathrm{x}}=\mathrm{x} / \mathrm{r}, \quad \theta=\left(\mathrm{t}^{2}+\mathrm{r}^{2}\right)^{1 / 2}$.

Lemma 3.1 Let $n \geq 3$, let $\left.h \in \varphi^{\prime}(0, \infty) ; \mathbb{R}\right)$ with $h$ and $r d h / d r$ in $L^{\infty}\left(\mathbb{R}^{+}\right)$ and let $\varphi \in \Sigma_{1}$. Then

$$
\begin{align*}
& Q_{m}(t, \varphi) \equiv \operatorname{Inf}_{\psi} Q_{0}(t, \varphi, \psi)=\left\|\theta r^{-1} \mathrm{~L} \varphi\right\|_{2}^{2} \\
& +\left\|\theta^{-1}\left\{\left(t^{2}-r^{2}\right) \hat{x} \cdot \nabla \varphi-\left(n-1-\theta^{2} r^{-2} h\right) r \varphi\right\}\right\|_{2}^{2} \\
& +\left\langle\varphi,\left\{\theta^{2} r^{-2} h(n-2-h)+\left(t^{2} r^{-2}-1\right) r d h / d r\right\} \varphi\right\rangle . \tag{3.8}
\end{align*}
$$

We refer to [8] for the proof of Lemma 3.1, which is an elementary computation since $Q_{0}(t, \varphi, \psi)$ is a quadratic function of $\psi$. That lemma provides a control in the $L^{2}$-norm sense of (i) the angular derivatives of $\varphi$ with an extra factor $\theta / r$, (ii) the radial derivative of $\varphi$ with a factor which vanishes on the light cone, since

$$
\theta^{-1}\left|t^{2}-r^{2}\right| \hat{x} \cdot \nabla \varphi \sim|t-r| d \varphi / d r
$$

(by choosing $h=(n-1) r^{2} / \theta^{2}$ ), and (iii) $\varphi$ itself with an extra factor $\theta / r$ (by choosing $h=(n-2) / 2)$. The last result is a variant of the Hardy inequality.

We can now state the main result of this section, which improves over Proposition 3.1.

Proposition 3.2 Let $n \geq 3$, let $2 \leq \ell \leq 2^{*}$, let $\tilde{\ell}=\operatorname{Max}(\ell, 2 n /(n-1))$, let $a \geq 0$, let $t>0$ and $\varphi \in \Sigma_{1}$. Then $\varphi$ satisfies the estimate

$$
\begin{align*}
& \left\|\left(\theta r^{-1}\right)^{1-\delta(\ell)} \theta^{\gamma(\ell)}\left(\theta^{-1}\left|t^{2}-r^{2}\right|+a\right)^{\alpha(\ell)} \varphi\right\|_{\ell} \leq c_{0}^{\delta(\ell)} \\
& \times\left\|\theta r^{-1} \varphi\right\|_{2}^{1-\delta(\ell)}\left\{\left\|\theta r^{-1} L \varphi\right\|_{2}^{2}+\left\|\theta r^{-1} \varphi\right\|_{2}^{2}\right\}^{\gamma(\ell) / 2} \\
& \left.\times\left\{\| \theta^{-1} \mid t^{2}-r^{2}\right\}\left(\frac{d \varphi}{d r}+\frac{n}{\tilde{\imath r}} \varphi\right)\left\|_{2}^{2}+a^{2}\right\| \frac{d \varphi}{d r} \|_{2}^{2}\right\}^{\alpha(\ell) / 2} \\
& \leq C_{1} Q_{m}(t, \varphi)^{(1-\alpha(\ell)) / 2}\left(Q_{m}(t, \varphi)+a^{2}\left\|\frac{d \varphi}{d r}\right\|_{2}^{2}\right)^{\alpha(\ell) / 2} \tag{3.9}
\end{align*}
$$

where the constants $C_{0}$ and $C_{1}$ depend only on $n$.

The proof is a complicated variant of the elementary proof [1] of the usual Sobolev inequality $\|u\|_{2} * \leq C\|\nabla u\|_{2}$. In the latter one starts with the estimates

$$
|v(x)| \leq(1 / 2) \int d x_{j}^{\prime}\left|\partial_{j} v\left(x_{1} \ldots, x_{j}^{\prime}, \ldots, x_{n}\right)\right|
$$

for $l \leq j \leq n$, one applies Hölder's inequality repeatedly, one substitutes $v=|u|^{p}$ for a suitable $p$, and one estimates the derivatives of $u$ in $\mathrm{L}^{2}$. Here we proceed similarly, but we use angular and radial instead of cartesian coordinates, and we introduce suitable weight factors in the starting point inequalities so as to reconstruct the weight factors occuring in $Q_{m}(t, \varphi)$ at the end of the computation. Eventually the angular derivatives of $\varphi$ are controlled by the term $(\theta / r) L \varphi$ in $Q_{m}$, and the radial derivative is controlled by the radial term in $Q_{m}$
(which yields (3.9) with $a=0$ ) or is left untouched (which gives the term with a in (3.9)). We refer to [8] for the details.

The estimate (3.9) for $\varphi$ is stronger than the estimate (3.6)
of Proposition 3.1 in every respect. In fact, the allowed values of $\ell$ range over the interval $\left[2,2^{*}\right]$ for (3.9) instead of the interval ( $\left.2, \ell_{l}\right]$ for (3.6). In the left-hand side of (3.9), the $L^{\ell}$ - norm is improved by various additional factors : the factor ( $\left.\theta r^{-1}\right)^{l-\delta(\ell)}$ yields the Hardy inequality, the factor $\theta^{\gamma(\ell)}$ yields the time decay $t^{-\gamma(\ell)}$ and an additional decrease at infinity in space; finally the factor $\left(\theta^{-1}\left|t^{2}-r^{2}\right|\right)^{\alpha(\ell)}$ yields an additional decay away from the light cone.

It is interesting to remark that the norm that appears in the left-hand side of (3.9) has some similarity with the norm used in [15] for $n=3$ to prove the existence of global solutions for small data or equivalently the existence of dispersive solutions of the equation (1.l) down to the optimal values $P_{0}(3)=1+\sqrt{2}$, namely (see (4.9a) of [15])

$$
\left\|(t+r)(1+|t-r|)^{p-2} \varphi\right\|_{\infty}
$$

We finally remark that the estimates of Proposition 3.2 still
hold if $\varphi$ is a scalar field minimally coupled with an external YangMills field, with the ordinary derivatives replaced by covariant derivatives. That property follows from the fact that for any function $v$ with values in the space of Yang-Mills potentials or in the space relevant for the coexisting scalar fields, the following inequality holds

$$
\partial_{\mu}|v| \leq\left|D_{\mu} v\right|
$$

where $|$.$| denotes the norm in the relevant space, and D_{\mu}$ the covariant derivative corresponding to $\partial_{\mu}$ (see [5], especially the Appendix).

## 4. TIME DECAY

In this section, we apply the previous results to the derivation of time decay properties of the solutions of the equation (l.1). Those decay properties will result in a straightforward way from combining the estimate of Proposition 3.2 (for $n \geq 3$ ) and its analogue for $n=2$ with boundedness properties of $Q_{0}(t, \varphi, \dot{\varphi})$, and we concentrate our attention on the derivation of the latter. We assume throughout this section that $\left(\varphi_{0}, \psi_{0}\right) \in \Sigma$ and that $f$ satisfies the assumptions (Al) and (A2). We need in addition a repulsivity condition (A3) :
(A3) f satisfies the inequalities

$$
\begin{equation*}
0 \leq\left(p_{1}+1\right) V(z) \leq 2 \operatorname{Re} \bar{z} f(z) \tag{4.1}
\end{equation*}
$$

for all $z \in \mathbb{C}$. (Equivalently, $V(R)=R^{p_{1}+1} v(R)$ where $v$ is a non negative non decreasing function from $\mathbb{R}^{+}$to $\mathbb{R}$ ).

In the case of a single power interaction (1.2) with $p=p_{1}$, (A3) simply says that $\lambda \geq 0$. Note also that (A3) implies (A2 b, c).

We first consider the easy case where $p_{1}+1 \geq \ell_{1}$.

Proposition 4.1 Let $n \geq 2$, let $f$ satisfies (A1), (A2) and (A3) with $p_{1}+1 \geq \ell_{1}$. Let $\left(\varphi_{0}, \psi_{0}\right) \in \Sigma$, and let $(\varphi, \dot{\varphi})$ be the solution of (l.l) in $\mathscr{(} \operatorname{R}, \Sigma)$ with initial data $\left(\varphi_{0}, \psi_{0}\right)$ at time zero, as described in Proposition 2.2. Then $Q_{0}(t, \varphi, \dot{\varphi})$ is uniformly bounded in time :

$$
\begin{align*}
Q_{0}(t, \varphi, \dot{\varphi}) & \leq Q(t, \varphi, \dot{\varphi}) \leq Q\left(0, \varphi_{0}, \psi_{0}\right) \\
& =\left\|x \psi_{0}\right\|_{2}^{2}+\left\|r \nabla \varphi_{0}\right\|_{2}^{2}+\int d x r^{2} v\left(\varphi_{0}\right) \tag{4.2}
\end{align*}
$$

Proof. The result follows imediately from the fact that under the assumption (A3), one has $W(\varphi) \leq 0$ and $t R(t, \varphi) \leq 0$ (See (1.10)-(1.11)).
Q.E.D.

We next turn to the difficult case $p_{1}+1 \leq \ell_{1}$. We assume nevertheless that $p_{1} \geq 1+4 / n$, a condition which is satisfied by the values of $p, p_{1}(n)$ and $p_{2}(n)$, for all $n$. That condition also makes the assumption $\int d x r^{2} V\left(\varphi_{0}\right)$ in Proposition 2.2 unnecessary. We need a slightly stronger version of the repulsivity condition (A3) :
(A3') $f$ satisfies the inequalities

$$
\begin{equation*}
c|z|^{p_{1}+1} \leq\left(p_{1}+1\right) V(z) \leq 2 \operatorname{Re} \bar{z} f(z) \tag{4.3}
\end{equation*}
$$

for some $C>0$ and all $z \in \mathbb{C}$. (Equivalently, if $V(R)=R_{1}^{p_{1}^{+1}} v(R)$ one assumes in addition that $v(0)>0)$.

We follow the method sketched in the introduction. The first task consists in extracting some preliminary decay from the conservation law (See (1.29)).

Lemma 4.1 Let $n \geq 2$, let $\ddagger$ satisfy (A1), (A2) and (A3') and let

$$
\begin{equation*}
\mu \equiv n+1-\left(p_{1}+1\right)(n-1) / 2 \geq 0 \tag{4.4}
\end{equation*}
$$

Let $\left(\varphi_{0}, \psi_{0}\right) \in \Sigma$, let $t_{0} \in \mathbb{R}$ and let $(\varphi, \dot{\varphi})$ be the solution of (1.1) in $\mathscr{C}(\mathbb{R}, \Sigma)$ with initial data $\left(\varphi_{0}, \psi_{0}\right)$ at time $t_{0}$. Then

$$
\begin{align*}
& m(t) \equiv Q(t, \varphi, \dot{\varphi})+E(\varphi, \dot{\varphi}) \leq m(0)\left(1+t^{2}\right)^{\mu}  \tag{4.5}\\
& \int d x V(\varphi) \leq m(0)\left(1+t^{2}\right)^{\mu-1}  \tag{4.6}\\
& \|\varphi(t)\|_{p_{1}+1} \leq c m(0)^{1 /\left(p_{1}+1\right)}(1+|t|)^{1-2 \delta\left(p_{1}+1\right)} \tag{4.7}
\end{align*}
$$

Proof. From the conservation law and from (A3), we obtain the integral form of the differential inequality

$$
\begin{align*}
d m / d t & =2 t \int d x W(\varphi) \leq 2 \mu t \int d x V(\varphi) \\
& \leq 2 \mu t\left(1+t^{2}\right)^{-1} m(t) \tag{4.8}
\end{align*}
$$

which implies (4.5) and therefore (4.6). Finally (4.7) follows from (4.6) and (A3') .
Q.E.D.

In order to estimate $Q_{0}(t, \varphi, \varphi)$, we first rewrite it by using (1.12) as

$$
\begin{equation*}
Q_{0}(t, \varphi, \dot{\varphi})=\|M \varphi\|_{2}^{2}+\|L \varphi\|_{2}^{2}+\|D \varphi\|_{2}^{2}=\sum_{A}\left\|\phi_{A}\right\|_{2}^{2} \tag{4.9}
\end{equation*}
$$

where in addition to the angular momentum $L=x \times \nabla$, we have introduced the operators

$$
\begin{align*}
& M=t \nabla+x(d / d t)  \tag{4.10}\\
& D=x \cdot \nabla+t(d / d t)+n-1 \tag{4.11}
\end{align*}
$$

and the sum in the last member of (4.9) runs over $A=L, M, D$, with $\phi_{A}=A \varphi$. The operators $L, M$ and $D$ are the infinitesimal generators of space rotations, of pure Lorentz transformations, and space-time dilations respectively. The expression (4.9) is the analogue in the NLW case of the expression $Q_{0 S}=(1 / 2)\|J \varphi\|_{2}^{2}$ in the NLS case (see (1.8) $S$ and (1.2l)). Note that in that case, $J$ is the infinitesimal generator of pure Galilei transformations. From the transformation
properties of the free wave equation, or more simply from a direct computation, it follows that

$$
\begin{equation*}
[\square, A]=0 \text { for } A=L, M ; \square D=(D+2) \square \tag{4.12}
\end{equation*}
$$

so that the functions $\phi_{A}$ satisfy the equations

$$
\left\{\begin{array}{l}
\square \phi_{A}+A f=0 \text { for } A=L, M  \tag{4.13}\\
\square \phi_{D}+(D+2) f=0
\end{array}\right.
$$

The interaction term in (4.13) can be rewritten as

$$
\begin{align*}
& A f=\phi_{A} \frac{\partial f}{\partial \varphi}+\bar{\phi}_{A} \frac{\partial f}{\partial \bar{\varphi}}=\phi_{A} f^{\prime}(\varphi) \quad \text { for } A=L, M  \tag{4.14}\\
& (D+2) f=\phi_{D} f^{\prime}(\varphi)+(n+1) f-(n-1) \varphi f^{\prime}(\varphi)
\end{align*}
$$

The Cauchy problem for (1.1) with initial data ( $\varphi_{0}, \psi_{0}$ ) at time zero can be rewritten in the form of the integral equation

$$
\begin{equation*}
\varphi(t)=\varphi(0)(t)-\int_{0}^{t} d \tau K(t-\tau) f(\varphi(\tau)) \tag{4.15}
\end{equation*}
$$

where

$$
\begin{equation*}
\varphi(0)(t)=\dot{K}(t) \varphi_{0}+K(t) \psi_{0} \tag{4.16}
\end{equation*}
$$

Similarly, the functions $\phi_{A}(t)$ satisfy the integral equations

$$
\phi_{A}(t)=\phi_{A}^{(0)}(t)-\int_{0}^{t} d \tau K(t-\tau) \phi_{A}(\tau) f^{\prime}(\varphi(\tau)) \quad \text { for } A=L, M
$$

and a similar equation for $\phi_{D}$, with the functions $\phi_{A}^{(0)}$ expressible in terms of $\left(\varphi_{0}, \psi_{0}\right)$. One sees easily that $\phi_{A}^{(0)}$ is bounded in $L^{2}$ uniformIy in time (with an additional Log $t$ factor for $\phi_{M}^{(0)}$ if $n=2$ ) and it remains to estimate $\phi_{A}$ in $L^{2}$ by using the integral equations. The function $f^{\prime}(\varphi)$ in the integrand is estimated by the use of the estimates (4.5)-(4.7) and possibly additional estimates obtained therefrom by substituting them in the integral equation (4.15) for $\varphi$. In space dimension $n \geq 3$, the method requires the use of homogeneous Besov spaces [2] and of estimates of the operator $K(t)$ acting between such spaces. It is reasonably simple in dimensions 3 and 4. The case of higher dimensions is more complicated and has not been worked out in details.

Here we only state the final results for $n=2,3,4$, and refer to [8] and [9] for the details of the proofs in dimensions $n=2$ and $n=3$, 4 respectively. In the latter case, one needs a slight reinforcement of the assumption (Al), in the following form
(A1') $f \in \mathscr{C}^{2}(\mathbb{C}, \mathbb{C}), f(0)=f^{\prime}(0)=0$, and $f$ satisfies the estimate

$$
\left|f^{\prime \prime}(z)\right|=\operatorname{Max}\left(\left|\partial^{2} f / \partial z^{2}\right|,\left|\partial^{2} f / \partial z \partial \bar{z}\right|,\left|\partial^{2} f / \partial \bar{z}^{2}\right|\right) \leq C\left(|z|^{p_{1}^{-2}}+\mid z^{p_{2}}\right)
$$

for some $p_{1}, p_{2}$ with $1 \leq p_{1} \leq p_{2}<1+4 /(n-2)$ and all $z \in \mathbb{C}$.

Proposition 4.2 If $n=2$, let $f$ satisfy (A1), (A2) and (A3') with $2+\sqrt{5} \equiv p_{2}(2)<p_{1}(\leq 5)$. If $n=3,4$, let $f$ satisfy (Al'), (A2) and $\left(A 3^{\prime}\right)$ with $p_{2}(n)<p_{1}(\leq 1+4 /(n-1))$ and $p_{1} \leq p_{2} \leq 1+4 n[(n-2)(n+1)]^{-1}$. Let $\left(\varphi_{0}, \Psi_{0}\right) \in \Sigma$, let $t_{0} \in \mathbb{R}$ and let $(\varphi, \dot{\varphi}) \in \mathscr{C}(\mathbb{R}, \Sigma)$ be the solution of (1.1) with initial data $\left(\varphi_{0}, \psi_{0}\right)$ at time $t_{0}$. Then

$$
\begin{array}{ll}
Q_{0}(t, \varphi, \dot{\varphi}) \leq C\left(1+\log _{+}|t|\right)^{2} & \text { if } n=2 \\
Q_{0}(t, \varphi, \dot{\varphi}) \leq C & \text { if } n=3,4
\end{array}
$$

The time decay properties of the solutions of (l.l) follow from Proposition 3.2, its analogue for $n=2$, and Propositions 4.1 and 4.2. In particular all solutions of (1.l) with initial data in $\Sigma$ satisfy the optimal decay (1.15) for $2 \leq \ell \leq 2$ under the assumptions of proposition 4.1 for $n \geq 3$ and of Proposition 4.2 for $n=3$, 4 .

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# Energy Forms and White Noise Analysis 

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In this article we give a short introduction to white noise analysis, with special emphasis to generalized and positive generalized white noise functionals and the causal calculus. These will be the main ingredients for the discussion of Dirichlet and eneray forms in this framework as examples of application [8,9]. For more details on white noise analysis see e.g. $[6,7,11,12]$ and literature quoted there.

## 1. White Noise

For simplicity we shall work in this article mostly with white noise with one dimensional time parameter. The generalization of notions and statements to higher dimensional "time" will be obvious. Thus we consider the Gel'fand triple

$$
\begin{equation*}
S^{\prime}(\mathbb{R}) \not L^{2}(\mathbb{R}, d t) \curvearrowright S(\mathbb{R}) \tag{1.1}
\end{equation*}
$$

which by Minlos' theorem [3,5] induces a Gaussian measure du on the $\sigma$-algebra $B$ over $S^{\prime}(\mathbb{R})$ generated by the cylinder sets such that for all $\xi \in S(\mathbb{R})$

$$
\begin{equation*}
\int_{S^{\prime}(\mathbb{R})} e^{i\langle x, \xi\rangle} d \mu(x)=e^{-1 / 2(\xi, \xi)} \tag{1.2}
\end{equation*}
$$

where we denoted the pairing of $S^{\prime}(\mathbb{R})$ and $S(\mathbb{R})$ by $<\cdot, \cdot>$ and the $L^{2}(\mathbb{R}, d t)$ scalar product by ( $\cdot, \cdot$ ). In fact one can prove that $d \mu$ is supported by the Sobolev space $S_{-\alpha}(\mathbb{R}, d t), \alpha>1 / 2$, which is defined as the completion of $S(\mathbb{R})$ under the norm

$$
\begin{equation*}
\|f\|_{-\infty}=\left\|H^{-\alpha} f\right\| \tag{1.3}
\end{equation*}
$$

where the last norm is the one of $L^{2}(\mathbb{R}, d t)$ and $H$ is the Hamiltonian of the harmonic oscillator: $H=-d^{2} / d t^{2}+t^{2}+1$ (we have added 1 to the usual Hamiltonian for later purposes).

For $p \geq 1$, we shall denote the Banach space $L^{p}\left(S^{\prime}(\mathbb{R}), B, d \mu\right)$ by $\left(L^{p}\right)$.

It is well-known (e.g. [5]) that ( $\mathrm{L}^{2}$ ) is isometric to the symmetric Fock space over $L^{2}(\mathbb{R}, d t)$ (complexified)

$$
\begin{equation*}
\left(L^{2}\right) \simeq \underset{n=0}{\infty} L^{2}\left(\mathbb{R}^{n}, n!d^{n} t\right) \tag{1.4}
\end{equation*}
$$

where . denotes symmetrization and if $n=0$ we mean $\phi$. Therefore
 be visualized in the following way

$$
\begin{equation*}
\varphi^{(n)}(x)=\int_{\mathbb{R}^{n}} f^{(n)}(t): x^{\otimes n}:(t) d^{n} t, \quad x \in S^{\prime}(\mathbb{R}) \tag{1.5}
\end{equation*}
$$

where $f^{(n)} \in L^{2}\left(\mathbb{P}^{n}, n!d^{n} t\right)$ and $: \quad$ denotes a "Wick ordered product.". It is defined recursively by

$$
\begin{align*}
: x^{\circ}: \quad & 1 \\
: x:\left(t_{1}\right)= & x\left(t_{1}\right), \quad t_{1} \in \mathbb{R}  \tag{1.6}\\
: & x^{\otimes n}:(t)= \\
& x\left(t_{n}\right): x^{\otimes(n-1)}:\left(t_{1}, \ldots, t_{n-1}\right) \\
& \quad \sum_{k=1}^{n-1} \delta\left(t_{n}-t_{k}\right): x^{\otimes(n-2)}:\left(t_{1}, \ldots, k_{k}, \ldots, t_{n-1}\right)
\end{align*}
$$

( $t=\left(t_{1}, \ldots, t_{n}\right)$ ). Then it takes a moment's thought to see that for $f^{(n)} \in S\left(\mathbb{R}^{n}\right)(1.5)$ is well-defined, if one interpretes the integral on the right as dual pairing. For $f(n) \in L^{2}\left(\mathbb{R}^{n}\right)$ (1.5) has to be interpreted as the ( $L^{2}$ )-limit of expressions like (1.5) with $f^{(n)}$ approximated by functions in $S\left(\mathbb{R}^{n}\right)$. In terms of $\varphi=\Sigma \varphi^{(n)} \in\left(\mathbb{L}^{2}\right)$, $\varphi^{(n)}$ as in (1.5), the isometry (1.4) reads

$$
\begin{equation*}
\|\varphi\|_{\left(L^{2}\right)}^{2}=\sum_{n=0}^{\infty} n!\left\|f^{(n)}\right\|_{L^{2}\left(\mathbb{R}^{n}, d^{n} t\right)}^{2} \tag{1.7}
\end{equation*}
$$

Let us now introduce a useful transformation on (L) [11]:

$$
\begin{equation*}
(S \varphi)(\xi):=\int \varphi(x+\xi) d \mu(x), \quad \xi \in S(\mathbb{R}) \tag{1.8}
\end{equation*}
$$

One readily checks that $\varphi^{(n)}$ in (1.5) transforms as:

$$
\begin{equation*}
\left(S \varphi^{(n)}\right)(\xi)=\int_{\mathbb{R}^{n}} f^{(n)}(t) \xi^{\otimes n}(t) d^{n} t \tag{1.9}
\end{equation*}
$$

which we may view as the evaluation of the tensor $f^{(n)} \in\left(L^{2}(\mathbb{R}, d t)\right)^{\hat{\otimes n}}$ on $(\xi, \ldots, \xi) \in\left(L^{2}(\mathbb{R}, d t)\right)^{8 n}$; in other words, we identify $S \varphi^{(n)}$ in a unique way with an element in $\left(L^{2}(\mathbb{R}, d t)\right)^{\hat{\theta} n}$ and therefore the range of $S$ is the symmetric Fock space over $L^{2}(\mathbb{R}, d t): S$ implements the isometry (1.4).

We conclude this subsection by giving attention to a special element in $\left(L^{2}\right)$, in fact in $H^{(1)}$, namely $\left.B(t ; x)=\langle x, 1[0, t)\rangle, t\right\rangle 0$, (understood as an ( $\mathrm{L}^{2}$ )-limit, s.a.) which is a version of Brownian motion and due to this relation we interprete $\dot{B}(t ; x) \equiv x(t)$ as the time derivative of Brownian motion.

## 2. Generalized Functionals

There are many motivations for the introduction of generalized functionals of white noise, but here we confine ourselves to mention that in section 4 we shall construct energy forms based on certain postive generalized functionals. For other motivations we refer the reader to the literature quoted in the references. Also, it should be pointed out that the type and construction of such spaces of generalized
functionals - like in finite dimensions - will vary with the application one has in mind. For our purposes it is convenient to use a construction based on second quantized operators (cf. also [11, 15, 19]).

Consider again the symmetric Fock space over $L^{2}(\mathbb{R}, d t)$ :
(which we identify with the r.h.s. of (1.4)). If A is a linear, closable operator we may define it second quantization $\Gamma(A)$ on $\hat{F}$ by setting

$$
\begin{equation*}
\left.\Gamma(A)\right|_{\left(I^{2}(\mathbb{R}, d t)\right)^{\hat{\otimes} n}}=A^{\otimes n} \tag{2.2}
\end{equation*}
$$

and extending (2.2) linearly [16]. We choose $A=H$, $H$ being the Hamiltonian of the harmonic oscillator (cf. section 1). Note that

$$
\begin{equation*}
\left.\Gamma(H)\right|_{\left(L^{2}(\mathbb{R}, d t)\right)^{\hat{\otimes} n} \geq 2^{n}} \tag{2.3}
\end{equation*}
$$

The isometry (1.4) defines now a unitary image of $\Gamma(\mathrm{H})$ on ( $\mathrm{L}^{2}$ ) and we denote it by the same symbol for simplicity. Moreover we set for $p \in \mathbb{N}_{o}$

$$
\begin{equation*}
\left(S_{p}\right):=D\left(\Gamma\left(H^{p}\right)\right) \subset\left(L^{2}\right) \tag{2.4}
\end{equation*}
$$

$D\left(\Gamma\left(H^{p}\right)\right.$ ) denoting the completed domain of $\Gamma\left(H^{p}\right)$. Clearly $\left(S_{p}\right)$ is a Hilbert subspace of ( $L^{2}$ ). Also it is not hard to see that the system of norms $\left(\|\cdot\|_{2, p} ; p \in N_{o}\right)$ of the spaces $\left(S_{p}\right)$ is a compatible system (see e.g. [3]). Let us define the space

$$
\begin{equation*}
(S):=\cap_{p \in \mathbb{N}_{0}}\left(S_{p}\right) \tag{2.5}
\end{equation*}
$$

as the projective limit of the family $\left(\left(S_{p}\right) ; p \in \mathbf{N}_{o}\right)$. (S) is a countably Hilbert nuclear space [3,11]. Its dual is given by [3]

$$
\begin{equation*}
(S)^{*}=\underset{p \in \mathbb{N}_{0}}{u}\left(S_{-p}\right) \tag{2.6}
\end{equation*}
$$

where $\left(S_{-p}\right)$ is the dual of ( $S_{p}$ ). We have therefore obtained a nuclear rigging

$$
\begin{equation*}
(S)^{*} p\left(L^{2}\right) p(S) \tag{2.7}
\end{equation*}
$$

and it is easy to check that (S) is dense in (L) (cf. [8] for an argument). White noise functionals in (S) are called testfunctionals, while those in ( $S)^{*}$ are called generalized functionals. The dual pairing between $(S)^{*}$ and $(S)$ will also be denoted by <•,•>. A very useful property of (S) is [8,11]

Lemma 2.1:
$(S)$ is an algebra.

In order to give the reader an intuitive idea about the space ( $S$ ), we remark that in the representation (1.5) of $\varphi^{(n)}$ belonging to (S) $\cap H^{(n)}$ $f(n) \quad$ is a member of $\quad \hat{S}\left(\mathbb{R}^{n}\right)$ (symmetric Schwartz space) and $\left\|\varphi^{(n)}\right\|\left(L^{2}\right)$ falls off in $n$ faster than any exponential.

From the last remark it should be clear that there are elements in $(S) *$ which correspond formally to functionals like (1.5) with $f^{(n)}$ $\in S^{\prime}\left(\mathbb{R}^{n}\right)$, e.g. $: x(t)^{n}$ : and the like.

Another class of functionals in (S) ${ }^{*}$, so-called Gauss kernels, will be described next. This class will later provide the first examples of singular measures which yield energy forms.

Formally speaking we want to construct elements like

$$
\begin{equation*}
\Phi_{\mathrm{un}}(\mathrm{x})=\exp (-1 / 2<\mathrm{x}, \mathrm{Kx}>) \tag{2.8}
\end{equation*}
$$

in $(S)^{*}$, where $K$ is some operator on $L^{2}(\mathbb{R}, d t)$. In [18] (2.8) had been studied for $K$ of Hilbert-Schmidt type with the following result: if $K>-1,\left\|K(1+K)^{-1}\right\|_{\text {H.S. }}<1$ and $K$ is trace class, then (2.8) exists (as it stands) in ${ }^{\left(L^{2}\right)}$. If the trace class condition is dropped, $\Phi_{u n}$ belongs after a multiplicative renormalization ("division by the expectation") to ( $L^{2}$ ): we denote this element by

$$
\begin{equation*}
\Phi(x)=N \exp (-1 / 2<x, K x\rangle) \tag{2.9}
\end{equation*}
$$

For more general linear selfadjoint operators $K$ on $L^{2}(\mathbb{R}, d t)$ we
proved in [9] the following

Lemma 2.2:
Assume that $K>-1$ and that for some $p \in \mathbb{N}$
(i)

$$
\begin{equation*}
\left\|H^{-p} \cdot(1+K)^{-2} H^{-p}\right\|_{\text {H.S. }}<\infty \tag{ii}
\end{equation*}
$$

Then there is an element $\Phi$ in (S)* (in fact in ( $S_{-p}$ )), formally denoted as in (2.9), with characteristic functional

$$
\begin{equation*}
\langle\Phi, \exp i<\cdot, \xi\rangle>=\exp \left(-1 / 2\left(\xi,(1+K)^{-1} \xi\right)\right) \tag{2.11}
\end{equation*}
$$

$(\xi \in S(\mathbb{R}))$.

Remark: Without condition (2.10.b) this has been shown in [9] for $K$ with $K \geq-1+\varepsilon, \varepsilon>0$. The above lemma can be proved on the basis of this result by a simple limiting argument, which makes use of (2.10.b).

Gauss kernels have an important property: they are positive in the sense that they map every positive (a.c.) element in (S) into a positive number. Elements in $(S)^{*}$ with this property will be called from now on positive generalized functionals. A useful theorem about positive generalized functionals is found in a paper by Yokoi [19]:

Theorem 2.3:
If $\Phi \in(S)^{*}$ is positive, then there exists a unique finite positive measure $v_{\Phi}$ on $B$ so that for every $F \in(S)$

$$
\begin{equation*}
\langle\Phi, F\rangle=\int \tilde{F}(x) d \nu_{\Phi}(x) \tag{2.12}
\end{equation*}
$$

where $\tilde{F}$ is the continuous version of $F$.

## 3. White Noise Calculus

This subsection will be rather sketchy, since by now the white
noise calculus can be considered as being well-established. For further details we refer the interested reader to $[5,9,11,12]$ and references quoted there.

The idea is to regard the generalized random variable $x(t)$, i.e. white noise "at time $t$ ", as a continuously indexed system of coordinates and to build a differential calculus with these coordinates. Thus we introduce a differential operator $\partial_{t}=\partial / \partial x(t)$ as follows

## Definition 3.1:

On (S) we set

$$
\begin{equation*}
\partial_{t}:=s^{-1} \frac{\delta}{\delta \xi(t)} s, \quad t \in \mathbb{R} \tag{3.1}
\end{equation*}
$$

where $\delta / \delta \xi(t)$ is the Fréchet functional derivative, and $s$ is the transformation (1.8).

With this definition one can check that $\partial_{t}$ is indeed a derivation on (S) (i.e. it admits the Leibniz rule) and also one can prove a chain rule [14]. Of course the domain of definition of $\partial_{t}$ can be largely extended. Here we confine ourselves to mention the following results

Lemma 3.2:
(i) $\quad \partial_{t} \operatorname{maps}(S)$ continuously into itself
(ii) $\partial_{t}$ is an annihilation operator: if we denote $H_{+}^{(n)}:=(S)$ $n H^{(n)}$ (cf. section 2), then

$$
\begin{equation*}
\partial_{t}: H_{+}^{(n)} \rightarrow H_{+}^{(n-1)} \tag{3.2}
\end{equation*}
$$

The adjoint operator $\partial_{t}^{*}$ of $\partial_{t}$ is defined as that

$$
\begin{equation*}
\left\langle\partial_{t}^{*} \Phi, F\right\rangle=\left\langle\Phi, \partial_{t} F\right\rangle \tag{3.3}
\end{equation*}
$$

for all $\Phi \in(S)^{*}$ and all $F \in(S)$.

Lemma 3.3:
(i) $\quad \partial_{t}^{*} \operatorname{maps}(S)^{*}$ continuously (in the weak-*-topology) into itself
(ii) $\partial_{t}^{*}$ is a creation operator: denote $H_{-}^{(n)}:=\left(H_{-}^{(n)}\right)^{*}$, then

$$
\begin{equation*}
\partial_{t}^{*}: H_{-}^{(n)} \rightarrow H_{-}^{(n+1)} \tag{3.4}
\end{equation*}
$$

(iii) we have the canonical commutation relations

$$
\begin{align*}
& {\left[\partial_{t}, \partial_{s}^{*}\right]=\delta(t-s)}  \tag{3.5.a}\\
& {\left[\partial_{t}, \partial_{s}\right]=\left[\partial_{t}^{*}, \partial_{s}^{*}\right]=0} \tag{3.5.b}
\end{align*}
$$

(iv) multiplication by $x(t)$, $t \in \mathbb{R}$, is defined as a mapping from $(S)$ into $(S)^{*}$ by

$$
\begin{equation*}
x(t) \cdot=\partial_{t}^{*}+\partial_{t} \tag{3.6}
\end{equation*}
$$

Formula (3.6) is the starting point of a reformulation of stochastic integrals within the white noise calculus, which leads to important generalizations of the standard stochastic integrals, cf. [13].

## 4. Energy Forms

In this subsection we shall sketch how energy forms [2] can be handled within the framework of white noise analysis. For more details and references for other approaches of energy forms in infinite dimension we refer the reader to our paper [8]. A review about the finite dimensional case is given in [17].

Roughly speaking the main idea is to generalize the finite dimensional form

$$
\begin{equation*}
\varepsilon(f, g):=\int_{\mathbb{R}^{n}}(\nabla f)(x) \cdot(\nabla g)(x) d v(x) \tag{4.1}
\end{equation*}
$$

where $d v$ is some positive o-finite measure - the "ground-state measure" - to an infinite dimensional situation. If $\varepsilon$ in (4.1) is closf able, then it defines uniquely a selfadjoint positive operator $\hat{H}$ on
$L^{2}\left(\mathbb{R}^{n}, d \nu\right)$ [10]. This representation of quantum mechanics allows for constructions of Hamiltonians $\hat{H}$ which in the Schrödinger representation with $H=-\Delta+V$ would have very singular potentials $V$. Thus this strategy seems promising for quantum field theory.

We shall work in this section with white noise with d-dimensional time parameter, i.e. with the basic triplet

$$
\begin{equation*}
S^{\prime}\left(\mathbb{R}^{d}\right) \curvearrowright L^{2}\left(\mathbb{R}^{d}, d t\right) \curvearrowright S\left(\mathbb{R}^{d}\right) \tag{4.2}
\end{equation*}
$$

and without further mentioning we shall use the notions introduced in the preceding sections in this context.

From how on let $\Phi$ be some positive generalized functional in $(S)^{*}$. By theorem $2.3 \Phi$ defines a measure on $S^{\prime}\left(\mathbb{R}^{\mathrm{d}}\right)$ which we shall denote by $\Phi(x) d \mu(x)$, $d \mu$ being the white noise measure, although this measure is in general not absolutely continuous with respect to d .

On the basis of lemma 2.1 and lemma 2.2 (cf. also the remark after lemma 2.1) it is not hard show that the following lemma holds.

Lemma 4.1:
If $F, G \in(S)$, then also $\nabla F \cdot \nabla G \in(S)$ with

$$
\begin{equation*}
\nabla F:=\left(\partial_{t} F ; t \in \mathbb{R}^{d}\right) \tag{4.3}
\end{equation*}
$$

and

$$
\begin{equation*}
\nabla F \cdot \nabla G=\int_{\mathbb{R}^{d}} d t\left(\partial_{t} F\right)\left(\partial_{t} G\right) \tag{4.4}
\end{equation*}
$$

Therefore we may define the following sesquilinear form on (S):

$$
\begin{align*}
\varepsilon(F, G) & :=\langle\Phi, \nabla \bar{F} \cdot \nabla G\rangle  \tag{4.5}\\
& =\int(\nabla \bar{F}(x)) \cdot(\nabla G(x)) \Phi(x) d \mu(x) \tag{4.6}
\end{align*}
$$

Equality (4.6) follows from theorem 2.3 (upon identification of elements in (S) with their continuous representant). Since (S) is dense in $\left(L^{2}\right)_{\Phi}:=L^{2}\left(S^{\prime}\left(\mathbb{R}^{d}\right), \Phi d \mu\right), \varepsilon$ is a densely defined, symmetric form on $\left(L^{2}\right)_{\Phi}$.

Next arises the question of closability of the form $\varepsilon$ in ( $\left.L^{2}\right)_{\Phi^{\prime}}$,
which would guarantee the existence of a unique selfadjoint operator $\hat{H}$ on $\left(L^{2}\right)_{\Phi}$ with

$$
\begin{equation*}
\bar{\varepsilon}(F, G)=(F, \hat{H} G)_{\Phi} \tag{4.7}
\end{equation*}
$$

for all $F, G \in D\left(\hat{H}^{1 / 2}\right)=D(\bar{\varepsilon})$, where $\bar{\varepsilon}$ is the closure of $\varepsilon$ and $(\cdot, \cdot)_{\Phi}$ is the $\left(L^{2}\right)_{\Phi}$ scalar product, [10,p.322].

Definition 4.2:
Let $\Phi \in(S)^{*}$ be positive. $\Phi$ is called admissible iff its associated symmetric form $\varepsilon$ (cf. (4.5)) is closable.

We can prove the following criterium $[8,9]$

## Theorem 4.3:

If $\Phi \in(S)^{*}$, positive, is such that for every $t \in \mathbb{R}^{d} \partial_{t} \Phi=B(t) \Phi$ with

$$
\begin{equation*}
\int B(t) \eta(t) d t \in(S) \tag{4.8}
\end{equation*}
$$

for every $n \in S\left(\mathbb{R}^{d}\right)$, then $\Phi$ is admissible.

## Example 4.4:

The Gauss kernels of section 2 are admissible, if the operator $K$ (cf. (2.9)) on $L^{2}\left(\mathbb{R}^{d}, d t\right)$ is such that $K$ maps $S\left(\mathbb{R}^{d}\right)$ into itself. Specially, one can choose $K=\left(-\Delta+m^{2}\right)^{1 / 2}, \mathrm{~m}^{2}>0$. As is shown in [9], this provides a representation of the free relativistic massive boson field in $d$ space dimensions as a generalized white noise functional.

For a criterium in the case that the measure induced by $\Phi$ is absolutely continuous w.r.t. du, we refer to our paper [8].

In [1] also non-Gaussian measures will be treated.
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# Principles of Solitary Wave Stability 

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## 1. Introduction

A lot of physical systems can be described by nonlinear differential equations which admit solutions in the form of so-called solitary waves. By a solitary wave we mean a localized wave which keeps its form or shape. Physically, a basic problem is to understand the role of these localized nonlinear objects. One of the most important and natural requirements for solitary waves is the condition of stability.

For example, in particle physics classical nonlinear wave equations form a basis for the construction of quantum objects. In all investigations the first order approximation deals with the equations as if they were describing classical field configurations, rather than quantum operator fields. In such a case the demand for classical stability is motivated by the requirement that the corresponding quantum state should be stable.[ J ]

Unfortunately the meaning of stability is not unique in the sense that there exist a lot of different definitions (or notions) of stability used in the literature and the information about the relations between these concepts of stability are very meager. The aim of this contribution is therefore to give a brief survey about the stability problem and to report on recent progress in this field. Furthermore using simple conditions we prove some relations between several concepts of solitary wave stability.

In Section 2 we present a general framework for the stability of solitary waves. We restrict ourselves to Hamiltonian systems since the Hamiltonian structure will allow to extend all the stability methods developed in classical mechanics (i.e. for systems having finite degrees of freedom) to systems with infinite degrees of freedom.

In Section 3 we study the relations between several concepts of stability and prove some simple theorems on them. In particular, we are interested in the relation between energetic and nonlinear stability.

In Section 4 we report on recent numerical work concerning soliton stability in nonintegrable systems.
2. Concepts of Stability

A pointed out in the introduction we shall consider Hamiltonian systems only. In the standard form they can be written as

$$
\begin{equation*}
\frac{\mathrm{du}}{\mathrm{dt}}=J E^{\prime}(u) \tag{H}
\end{equation*}
$$

on a function space (e.g. a Hilbert space) where $E$ denotes the energy, $E^{\prime}$ its derivative 'with respect to $u$ ' and $J$ is a skew-symmetric linear operator.

Typical examples for ( $H$ ) on an infinite dimensional space which arise from classical field theory are nonlinear Klein-Gordon equations

$$
\begin{equation*}
\psi_{t t}-\Delta \psi+m^{2} \psi-g\left(|\psi|^{2}\right) \psi=0 \quad \text { on } R^{N} \tag{NLKG}
\end{equation*}
$$

nonlinear Schrödinger equations

$$
\begin{equation*}
i \phi_{t}+\Delta \phi+f\left(|\phi|^{2}\right) \phi=0 \quad \text { on } R^{N} \tag{NLS}
\end{equation*}
$$

and nonlinear Dirac equations

$$
\begin{align*}
& i \gamma^{0} \partial_{t} \psi+i \gamma^{k} \partial_{k} \psi-m \psi+\frac{\partial G(\bar{\psi} \Gamma \psi)}{\partial \bar{\psi}}=0 \quad \text { on } R^{N}  \tag{NLD}\\
& 1 \leq k \leq N, \quad \Gamma=\text { element of the } \gamma \text {-al gebra. }
\end{align*}
$$

An interesting property of all these equations is that they possess additional symmetries like translation invariance and (global) gauge invariance. These symmetries generate additional conserved quantities like momentum (coming from translation invariance) and charge (coming from gauge invariance). Therefore we may assume that $(H)$ is invariant under certain group representations. We follow now a recent paper where a rather general stability theory has been presented [GSS]. We assume That ( $H$ ) is invariant under a one-parameter group of operators $U(\cdot)$ (an extension to more dimensional abelian groups seems not to be difficult). Let $Q$ be the conserved quantity associated to $U(\cdot)$. A solitary wave solution is a solution of the form

$$
\begin{equation*}
u(t)=U(\omega t) \phi_{\omega} \tag{2.1}
\end{equation*}
$$

It may be viewed as a critical point of the energy $E$ subject to constant $Q$. This leads to the definition of energetic stability.

Definition 1: A solution of $(H)$ given by (2.1) is called energetically stable if its stationary part $\phi_{\omega}$ is a local minimum of the energy $E$ subject to constant $Q$.

A sufficient condition for energetic stability is that the linearized operator $H_{\omega}=E^{\prime \prime}\left(\phi_{\omega}\right)-\omega Q\left(\phi_{\omega}\right)$ is nonnegative. One should note that $H_{\omega}$ cannot be positive since the symmetry causes a nontrivial nullspace (see [GSS]).

Orbital (or nonlinear) stability is defined as follows.
Definition 2: The solution $U(\omega t) \phi_{\omega}$ is called orbitally stable if for any tubular neighborhood of $0_{\omega}=\left\{U(s) \phi_{\omega}\right.$,s real\} there exists a neighborhood $V$ of $\phi_{\omega}$ such that all trajectories $u(t)$ of $(H)$ which start in $V$ remain in the given tube for all time. Otherwise we call $U(\omega t) \phi_{\omega}$ unstable.

In most cases one uses energetic stability to prove orbital stability by taking the energy as a Liapunov functional [Be] but energetic stability is not a necessary condition for orbital stability [JR,SV].

Another method to study the stability properties of solitary wave solutions is to linearize the system around a solitary wave $U(\omega t) \phi_{\omega}$. The linearized dynamics is then described by the equation
$\left(H_{1 i n}\right) \quad \frac{d w}{d t}=J H_{\omega} w$
with $H_{\omega}=E^{\prime \prime}\left(\phi_{\omega}\right)-\omega Q^{\prime \prime}\left(\phi_{\omega}\right)$.
Definition 3: A solitary wave $U(\omega t) \phi_{\omega}$ is called linear dynamically stable if any solution of the linearized system remains bounded for $t \geq 0$.

Unfortunately this definition is not directly applicable since one has to solve the so called zero-mode problem [ ML]: The nontrivial nullspace of the linearized operator in the case of symmetries will generate solutions which grow polynomial in time. Physically, however, these zero modes (or secular modes) have no meaning for the stability properties. Apart from some particular cases [W1 ] we do not have a rigorous result for the linear dynamical stability.

Physically, an interesting problem is to study the 'collision' of two solitary waves. If one interprets solitary waves as particles then the interaction between two solitary waves yields information on the 'stability' of the considered formation. If the solitary waves interact elastically then they have (at least asymptotically) the same shape as before. In such a case these solitary waves are called soliton. Mathematically the soliton property is related to the complete integrability of the corresponding field equation. Solitons have been found analytically by the inverse scattering transform, e.g. for the sine-Gordon equation

$$
\begin{equation*}
\psi_{t t}-\psi_{x x}+\sin \psi=0 \tag{S-G}
\end{equation*}
$$

This method also permits an analytically study of soliton interactions.
Unfortunately nonintegrable systems, e.g. the $\phi^{4}$-model

$$
\begin{equation*}
\phi_{t t}-\phi_{x x}-\phi+\phi^{3}=0 \tag{4}
\end{equation*}
$$

abound in physics and the powerful methods available for integrable systems do not apply to these models.

## 3. Relations between energetic linear and nonlinear stability

To discuss the relation between energetic and nonlinear stability we start with a rather general but simple theorem.

Theorem 3.1: Let $\phi_{0}$ be a solution of $(H)$, i.e. $E^{\prime}\left(u_{0}\right)=0$ such that the 1 inear operator $E^{\prime \prime}\left(\phi_{0}\right)$ satisfies

$$
\begin{equation*}
<E^{\prime \prime}\left(\phi_{0}\right) w, w>\geq c\|w\|^{2} \text { for some } c>0 \tag{3.1}
\end{equation*}
$$

Then $\phi_{0}$ is linearly and nonlinearly stable.
Proof: First of all we note that (3.1) implies the energetic stability. To prove the linear stability we use the linearized energy < $E^{\prime \prime}\left(\phi_{0}\right) w, W>$ as a Liapunov functional. The linearized energy is conserved for $\left(H_{i n}\right)$ and the norm can be controlled by estimate (3.1). Each solution of the linearized system remains bounded which implies the linear stability of $u_{0}$.

To prove nonlinear stability we observe that (3.1) implies the existence of $\varepsilon>0$ such that

$$
\begin{equation*}
E(u)-E\left(\phi_{0}\right) \geq c\left\|u-\phi_{0}\right\|^{2} \tag{3.2}
\end{equation*}
$$

for all $u$ with $\left\|u-\phi_{0}\right\|<\varepsilon$. Estimate (3.2) will then imply the nonlinear stability. 口

If there is an additional conserved quantity $Q$, e.g. the charge, then an estimate of the type (3.1) for all perturbations which keep the quantity $Q$ fixed will be sufficient to prove nonlinear stability.

To illustrate this fact let us consider the following dimensional Klein Gordon equation

$$
\begin{equation*}
u_{t t}-U_{x x}+U-|U|^{2 p} u=0 \quad \text { on } \mathbb{R} \tag{3.3}
\end{equation*}
$$

It can be easily checked that for each $\omega$ with $\omega^{2}<1$ equation (3.3) possesses a unique solution (up to translations in space and phase) of the form $e^{i \omega t}{u_{\omega}}^{(x)}$ with $u_{\omega}$ symmetric, positive function strictly decreasing on $\mathbb{R}^{+}$. One can prove that if

$$
\begin{equation*}
d(\omega) \equiv \int u_{\omega}^{\prime 2}(x) d x \tag{3.4}
\end{equation*}
$$

satisfies $d^{\prime \prime}(\omega)>0$ then $u_{\omega}$ is energetically and nonlinearly stable.
The energy and the charge associated to (3.3) are given by

$$
\begin{aligned}
& E\left(u, u_{t}\right)=\frac{1}{2} \int\left|u_{t}\right|^{2}+\left|u_{x}\right|^{2}-\frac{1}{p+1}|u|^{2 p+2} d x \\
& Q\left(u, u_{t}\right)=\operatorname{Im} \int \bar{u} u_{t} d x
\end{aligned}
$$

Indeed in [BSV1] we show that estimate (3.1) is valid for all tangent vectors of curves $\left(u_{1}(\lambda), u_{2}(\lambda)\right)$ which keep the charge fixed and are orthogonal to the orbit of $u_{\omega}(x)$

$$
\begin{aligned}
& \left.\frac{d^{2}}{d \lambda^{2}}\right|_{\lambda=0} E\left(u_{1}(\lambda), u_{2}(\lambda)\right) \geq C\left(\left\|y_{1}\right\|\left\|^{2}+\right\| y_{2} \|^{2}\right) \\
& \text { where } y_{j}=\frac{d}{d \lambda} \quad u_{j}(x) \quad \lambda=0 .
\end{aligned}
$$

Results in the same direction have been obtained for a lot of nonlinear field equations [ $\mathrm{Be}, \mathrm{Bo}, \mathrm{GSS}, \mathrm{W} 2]$.

Instability results for nonlinear Klein Gordon equation have been proved if $d^{\prime \prime}(\omega)<0$ in [SS].

For nonlinear Dirac equations the situation is much more difficult. Solitary waves of the Gross-Neve model or the Thirring model are expected to be nonlinearly stable but as we showed in [BSV2] they are always energetically unstable which can be traced back to the indefinite 'Kinetic' term in the energy.

## 4. Soliton stability

We finish our lecture with a few remarks on soliton stability. As pointed out in the introduction the soliton stability is related to the complete integrability of the corresponding wave equation. In such case when the solitary waves interact, they are always scattered elastically, preserving asymptotically their shape. The integrability also permits an analytical study of the multisoliton interactions. [ DEGM]

In many real physical systems the basic models are not integrable, and in other cases the integrability condition (and then the solitonic stability) is destroyed due to physical perturbations (impurities, external fields, ...). In certain cases of small perturbations, the solitary wave interactions can be treated analytically, but, in general, we must study numerically the dynamics of the solitary waves collisions. The understanding of the interaction mechanisms can be used to given a measure of how far is the system from an integrable one. The natural mathematical
frame of the collision phenomenology would be a KAM theorem for infinite dimensional systems.

There are many numerical studies, begun in the mid 1970's, about the collisions of solitary waves in non-integrable systems [AKL, M, CP, CPS]. Here we summarize some of the obtained results:
(1) Collisions of solitary waves, in one spatial dimension, in confining models [ST,AKL] : In such collisions new localized objects are generated, which are pulsating in time, and they appear to be stable.
(2) Collisions of the solitary waves associated to the one-dimensional Dirac field with a scalar self-interaction [AC] :

$$
i \gamma^{\mu} \phi_{u} \psi-m \psi+2 \lambda(\bar{\psi} \psi) \psi=0 .
$$

The numerical experiments showed inelastic interactions and bound state production in binary collisions. Also it was observed charge and energy interchange except for some particular initial velocities of the solitary waves.
(3) Kink-Antikink ( $K \bar{K}$ ) collisions in the one-dimensional nonlinear Klein-Gordon equations

$$
\begin{aligned}
& \phi_{t t}-\phi_{x x}-\phi+\phi^{3}=0 \\
& \phi_{t t}-\phi_{x x}+\sin \phi+2 \lambda \sin \frac{\phi}{2}=0
\end{aligned}
$$

The following features are observed in the center mass frame (in which a Kink with velocity $V$ collides with an antikink with velocity $-V$ )
(a) For lower velocities than a critical value $\left(V_{c}\right)$ there is a"trapping" and an oscillatory state is created.
(b) Inelastic interactions: the final velocity of the solitary waves is less then their initial velocity.
(c) For well-defined velocities, below $V_{c}$, the $K \bar{K}$ reflect once, escape to finite separation, and finally reflect again before separating to infinity.
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